

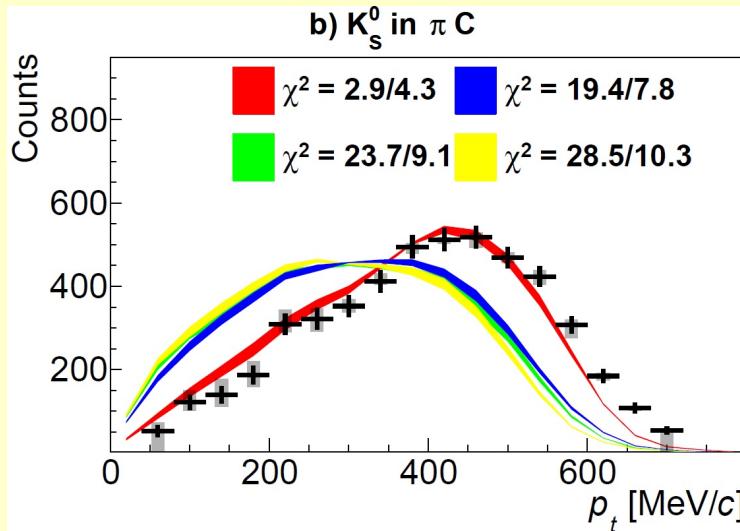
First test for hyperon potentials & implications to neutron star matter

Theodoros Gaitanos



ΤΜΗΜΑ ΦΥΣΙΚΗΣ

ΑΡΙΣΤΟΤΕΛΕΙΟ
ΠΑΝΕΠΙΣΤΗΜΙΟ
ΘΕΣΣΑΛΟΝΙΚΗΣ



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ΠΑΝΕΠΙΣΤΗΜΙΟ
ΘΕΣΣΑΛΟΝΙΚΗΣ

DAAD

Deutscher Akademischer Austausch Dienst
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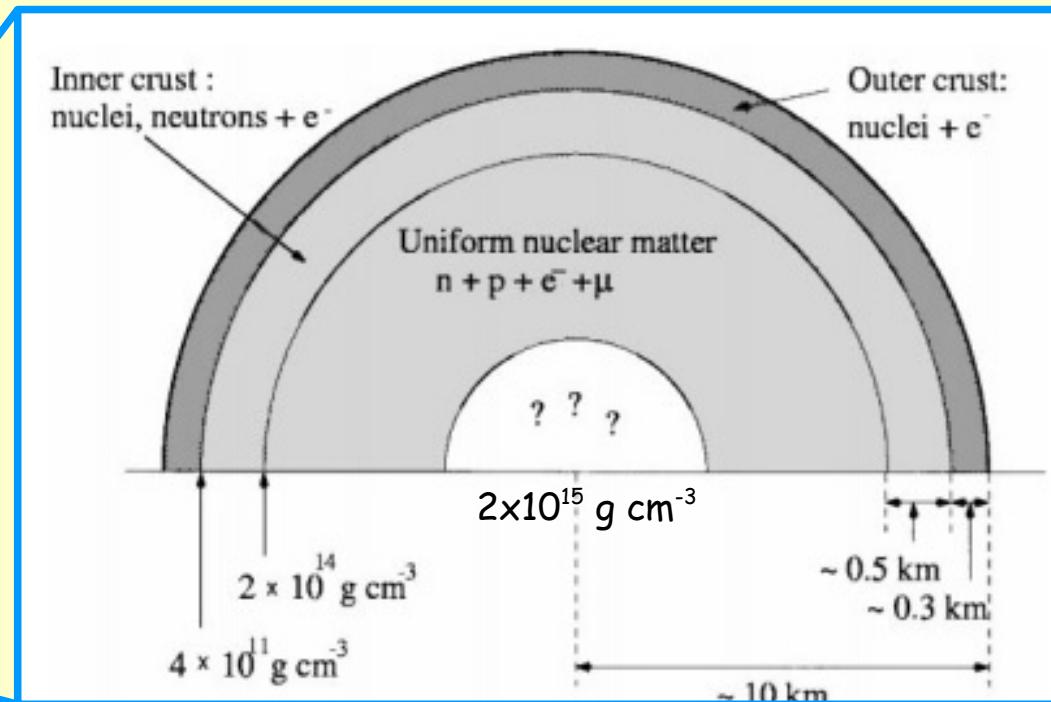
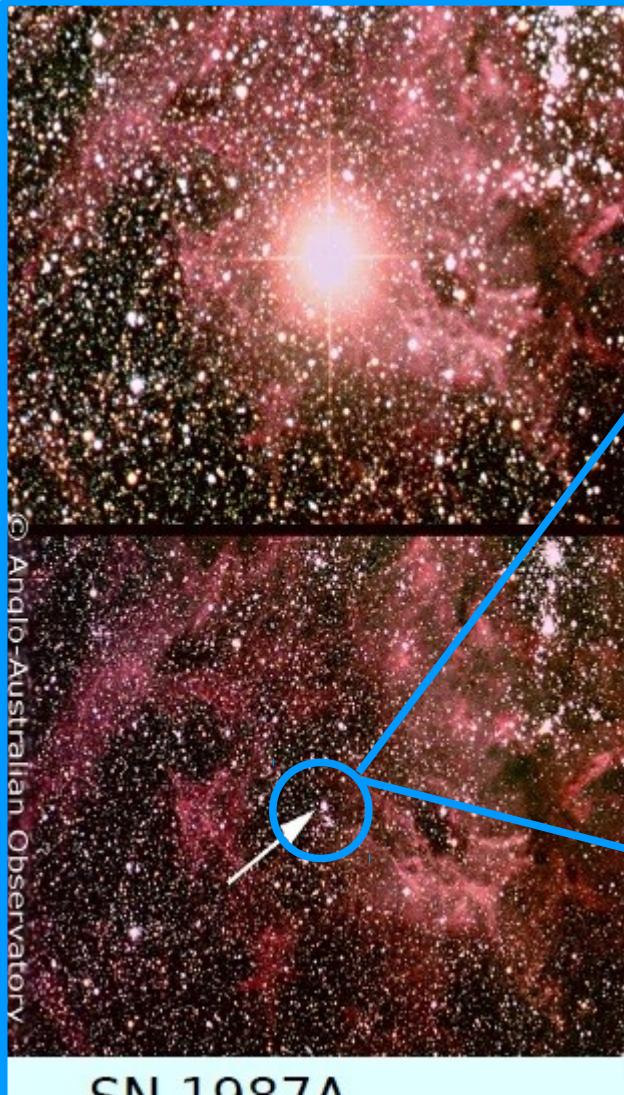
HIC | FAIR

Helmholtz International Center

Introduction...

Neutron Stars (NS)

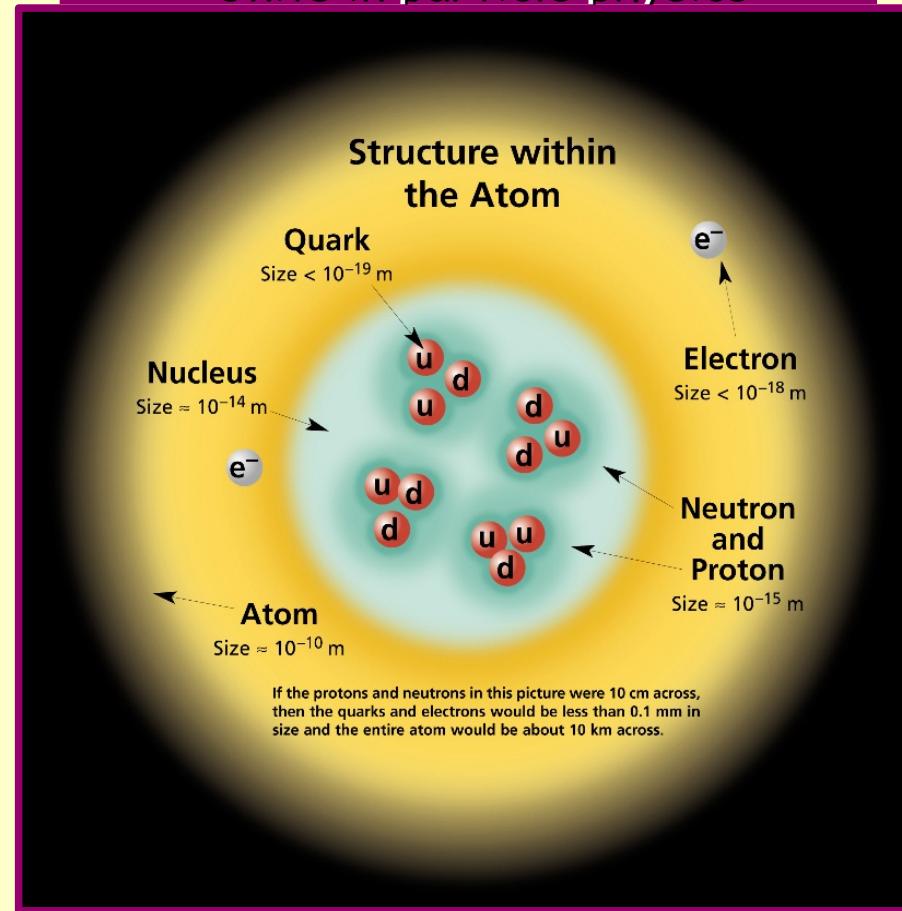
- stable "cosmic nucleus" made of p,n,e⁻,...
- densest visible matter in the universe
1.5-2 x solar mass inside 10km radius



earth's average density ~5.5 g cm⁻³
earth's radius ~6.371 km

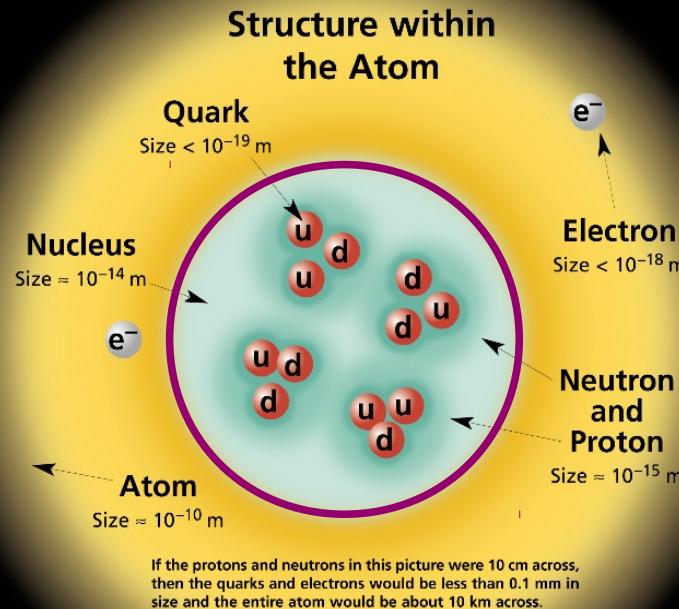
Introduction...

Units in particle physics



Introduction...

Units in particle physics



Units:

$$\hbar = c = 1$$

$$1 \text{ GeV} = 1.8 \cdot 10^{-24} \text{ g}$$

$$1 \text{ fm} = 10^{-15} \text{ m}$$

Temperature units:

$$[1 \text{ MeV}] \sim [10^{10} \text{ K}]$$

nuclei: $T \sim 0 \text{ MeV}$

(earth: $T \sim 180 \text{ K} - 330 \text{ K}$)

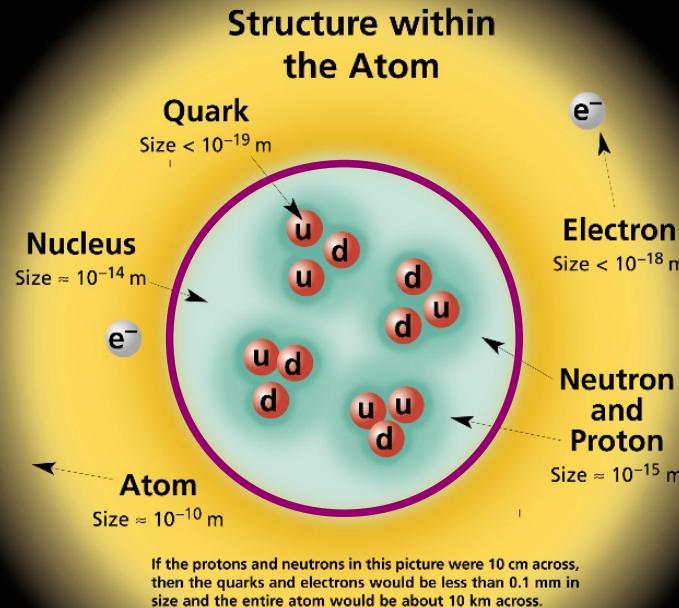
Density units:

$$\text{nuclei: } \rho_0 \sim 2.5 \cdot 10^{14} \text{ g cm}^{-3} = 0.165 \text{ fm}^{-3}$$

(earth: $\rho \sim \text{g cm}^{-3}$)

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Scales in hadron physics:

Baryon masses $M \sim 1000 \text{ MeV} = 1 \text{ GeV}$

Baryon potentials $V \sim 500 \text{ MeV} (\sim M!)$

Nuclear ground state $\sim 8 \text{ MeV}$ (small number!)

That means for us...

$T \sim 1 \text{ MeV}$ is a cold relativistic system

$T \sim 50 \text{ MeV}$ is a rather hot relativistic system

→ static neutron stars are cold objects

→ neutron star binaries can be cold/hot

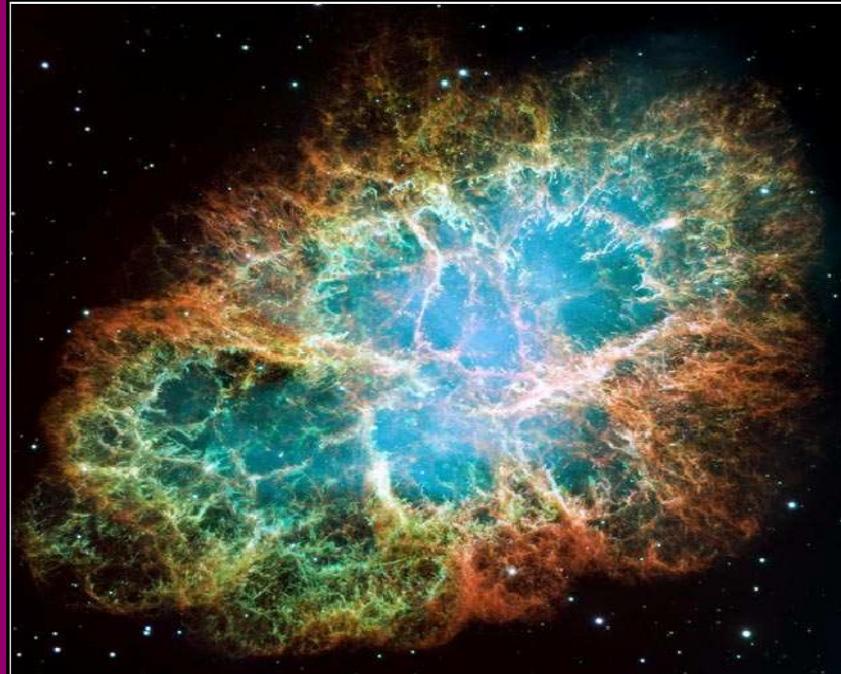
Introduction...

explore astrophysics...

- Properties of low/high density warm matter (supernovae, neutron stars)

Crab Nebula • M1

HST • WFPC2



NASA, ESA, and J. Hester (Arizona State University)

STScI-PRC05-37

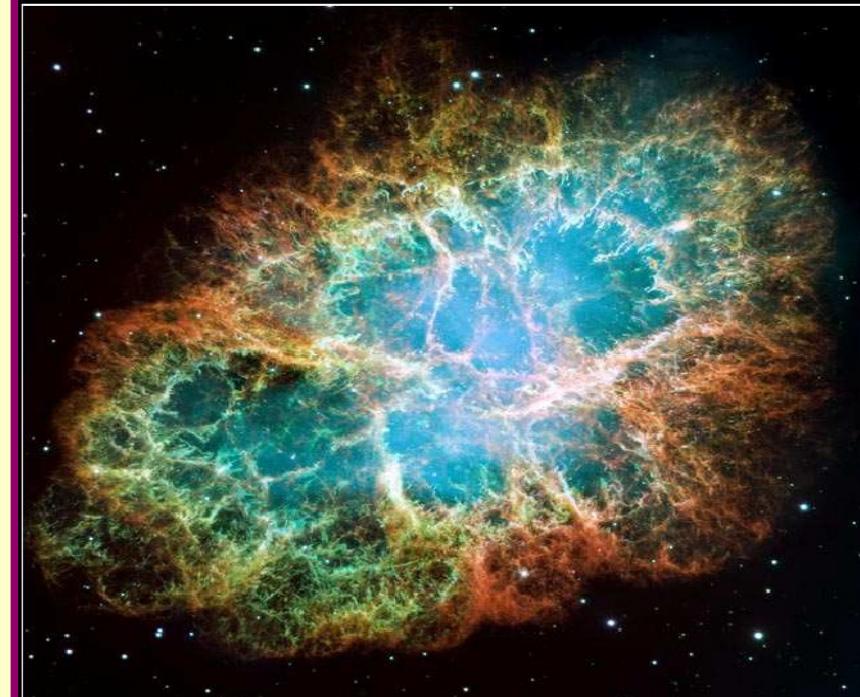
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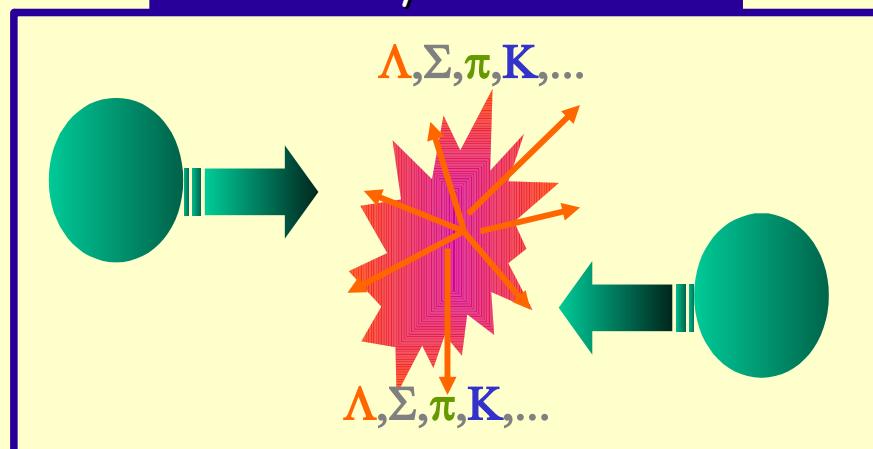


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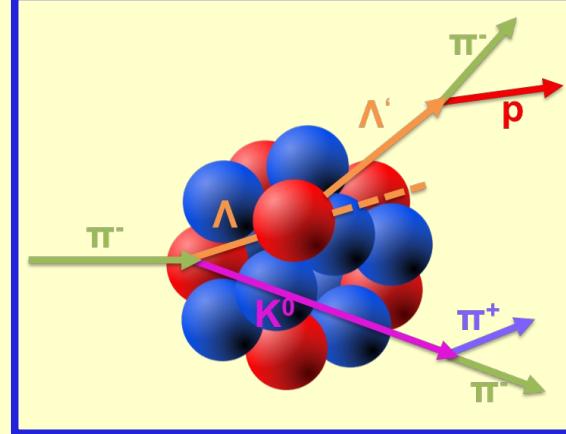
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...in the laboratory

HIC: Heavy-Ion Collisions

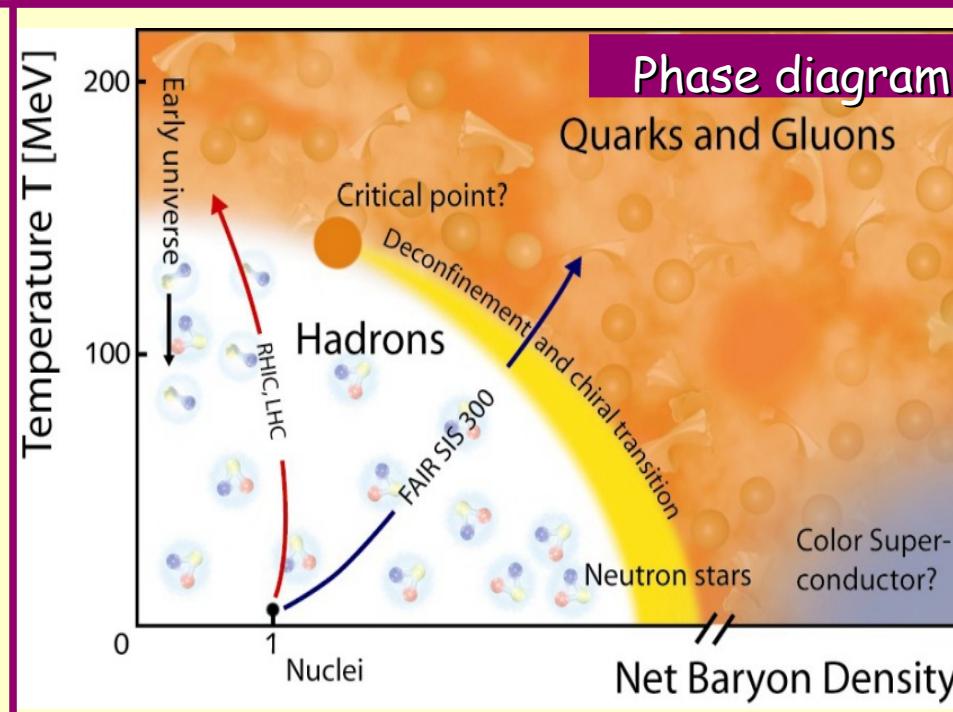
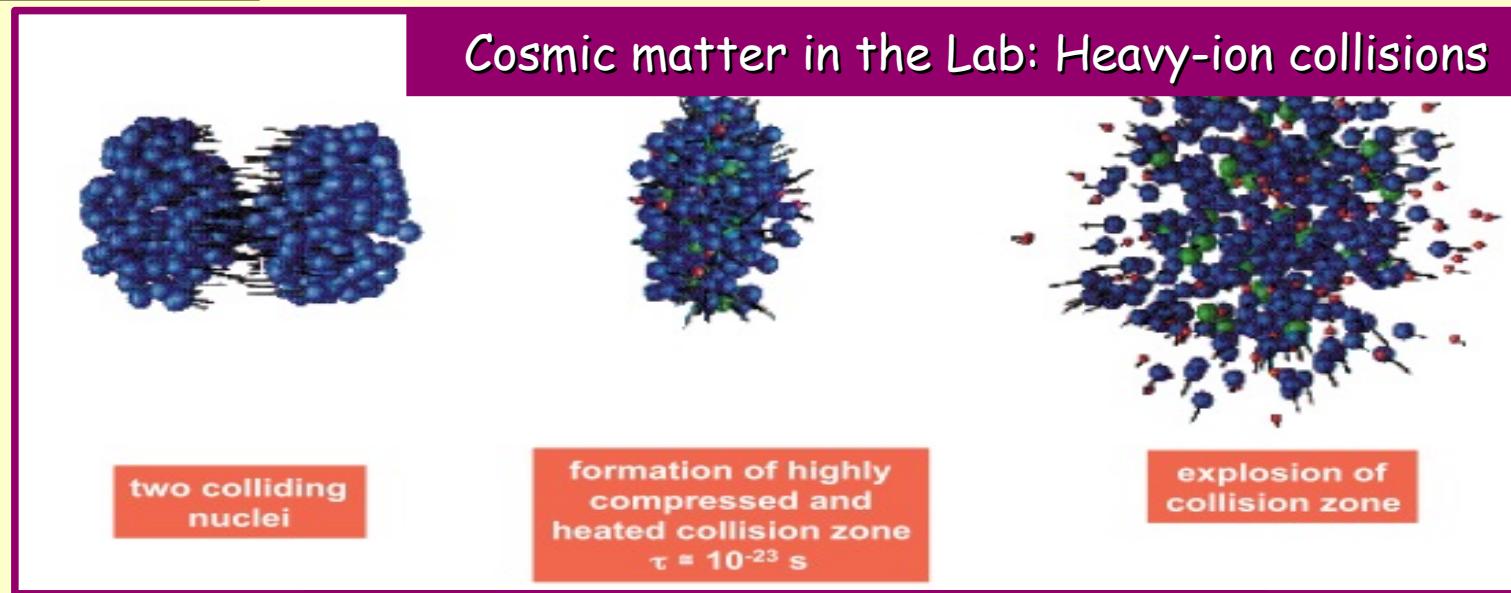


Hadron-induced reactions



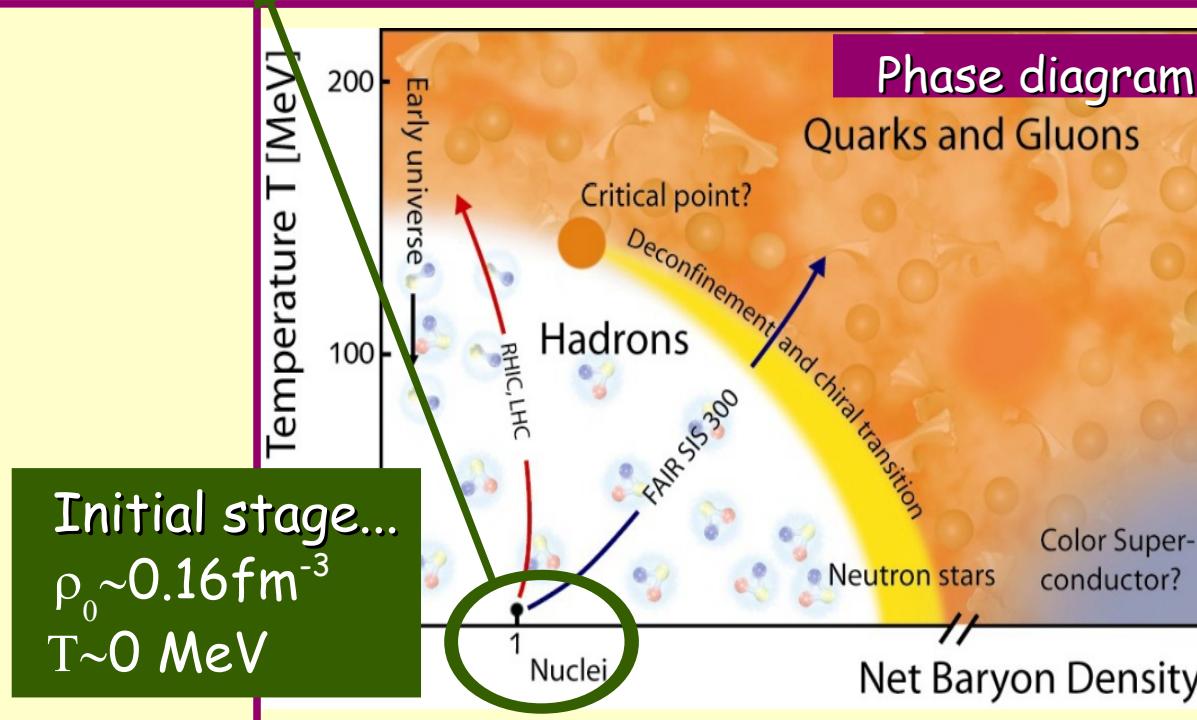
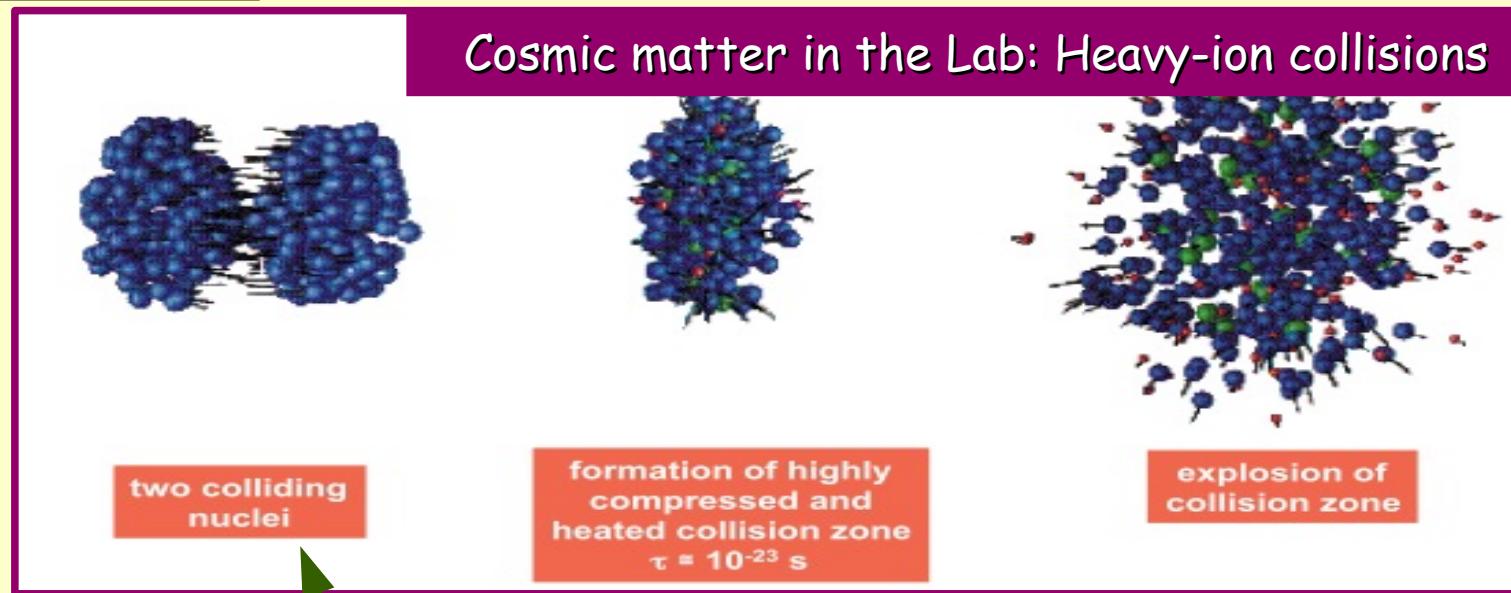
Introduction...

Cosmic matter in the Lab: Heavy-ion collisions



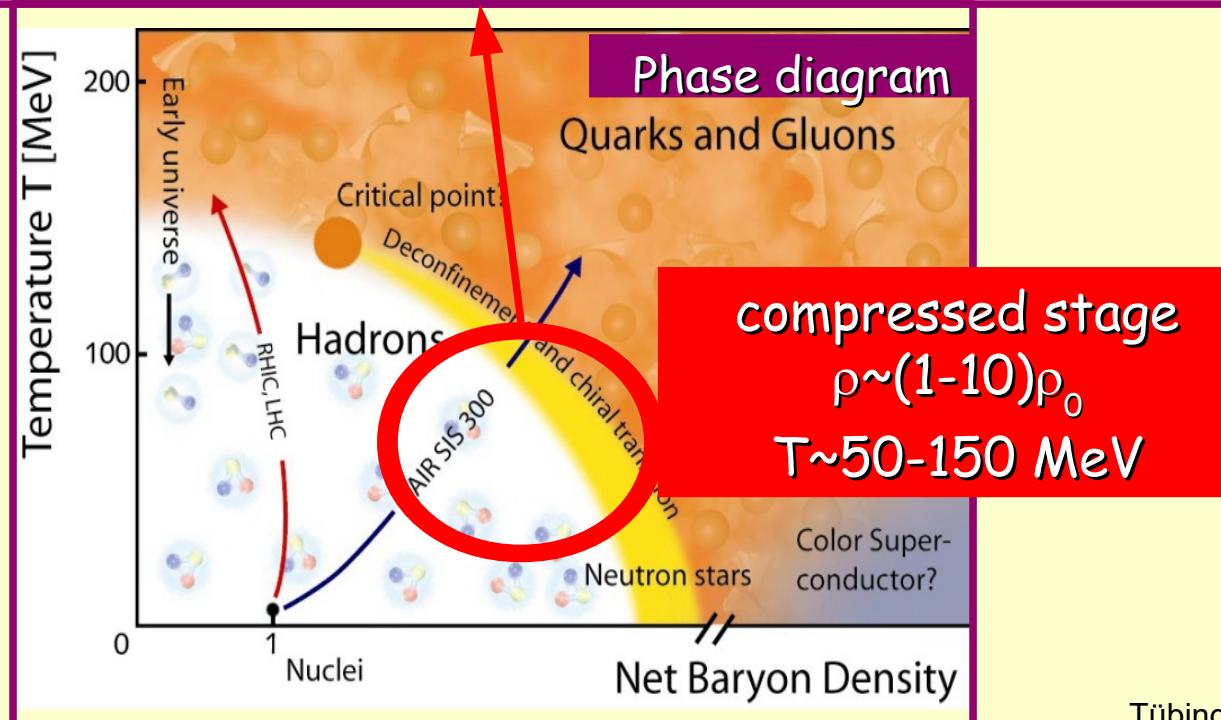
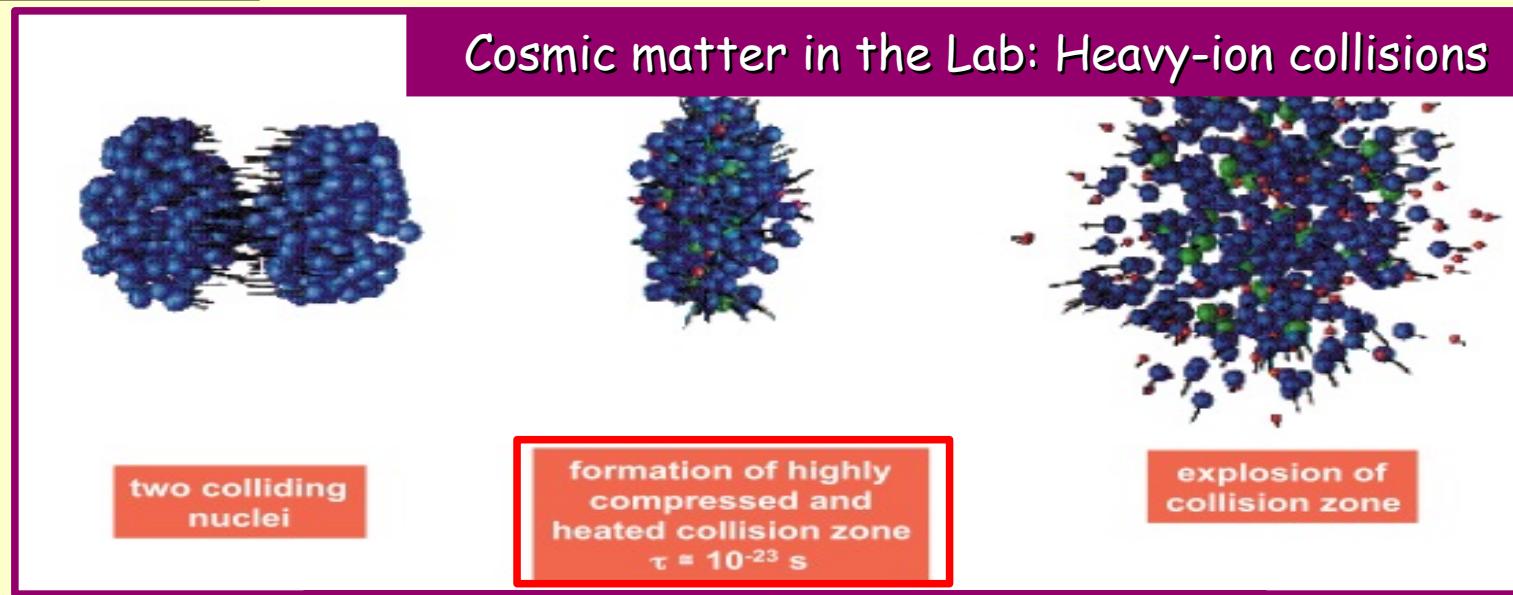
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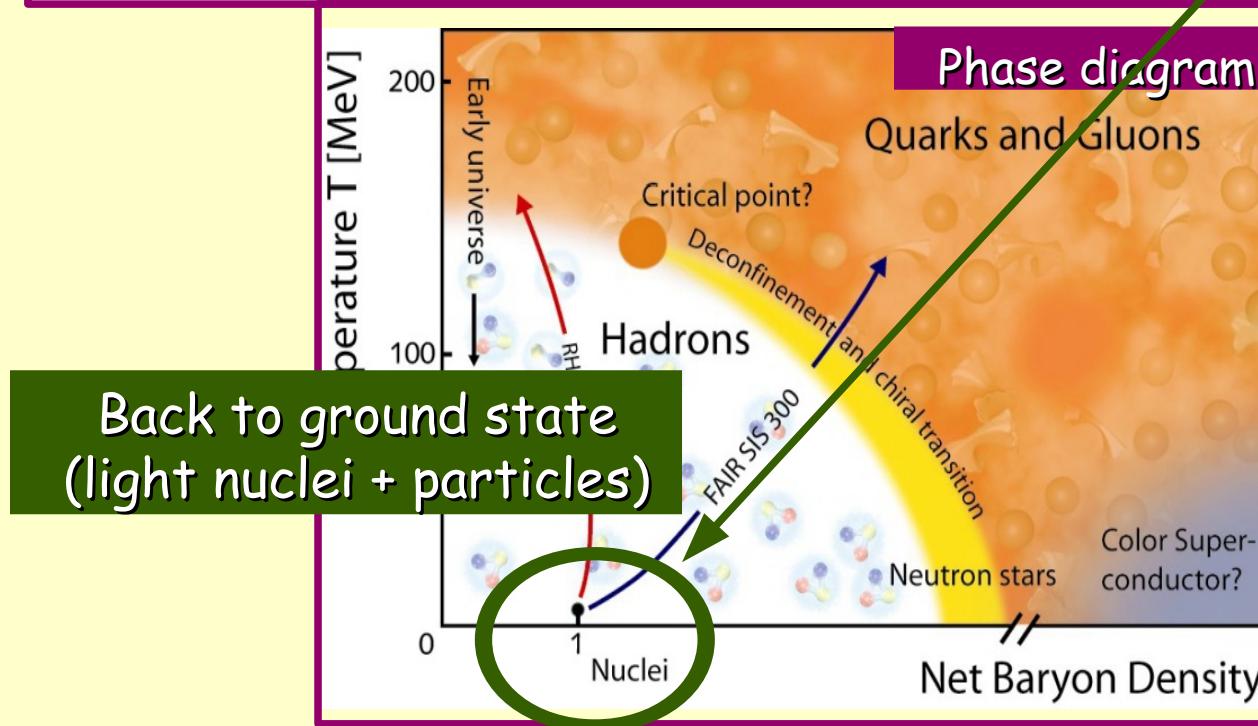
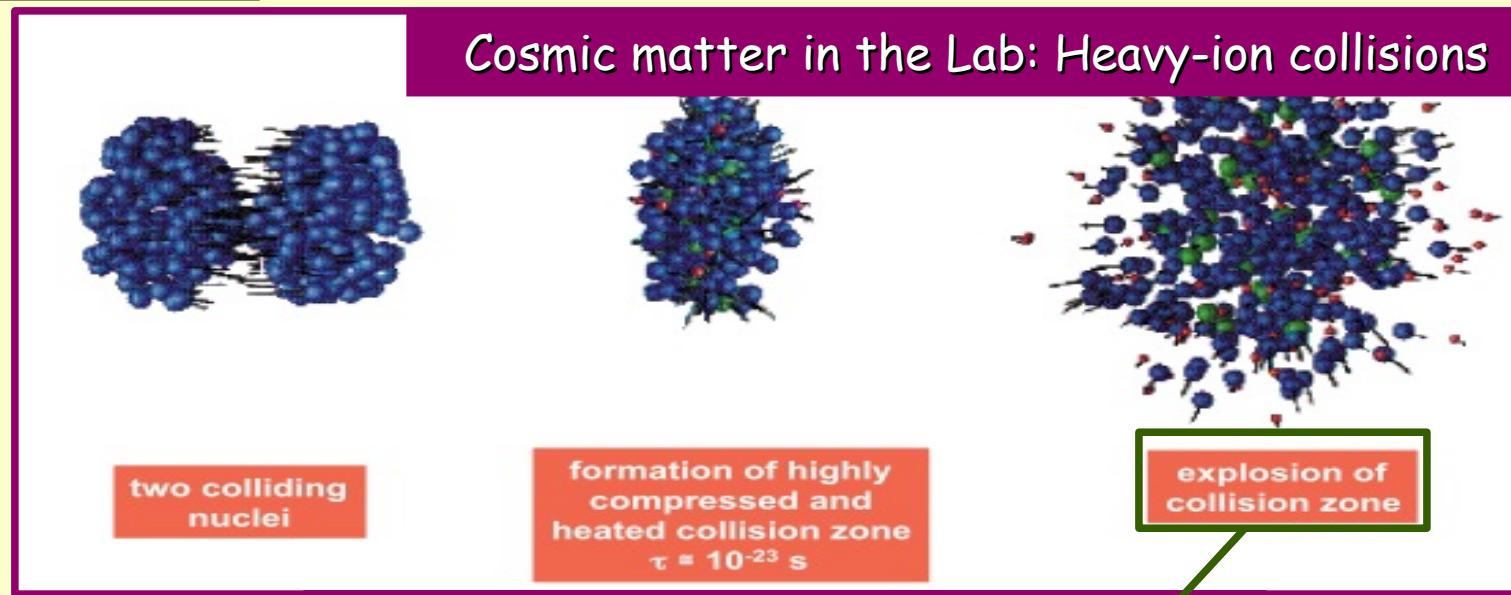
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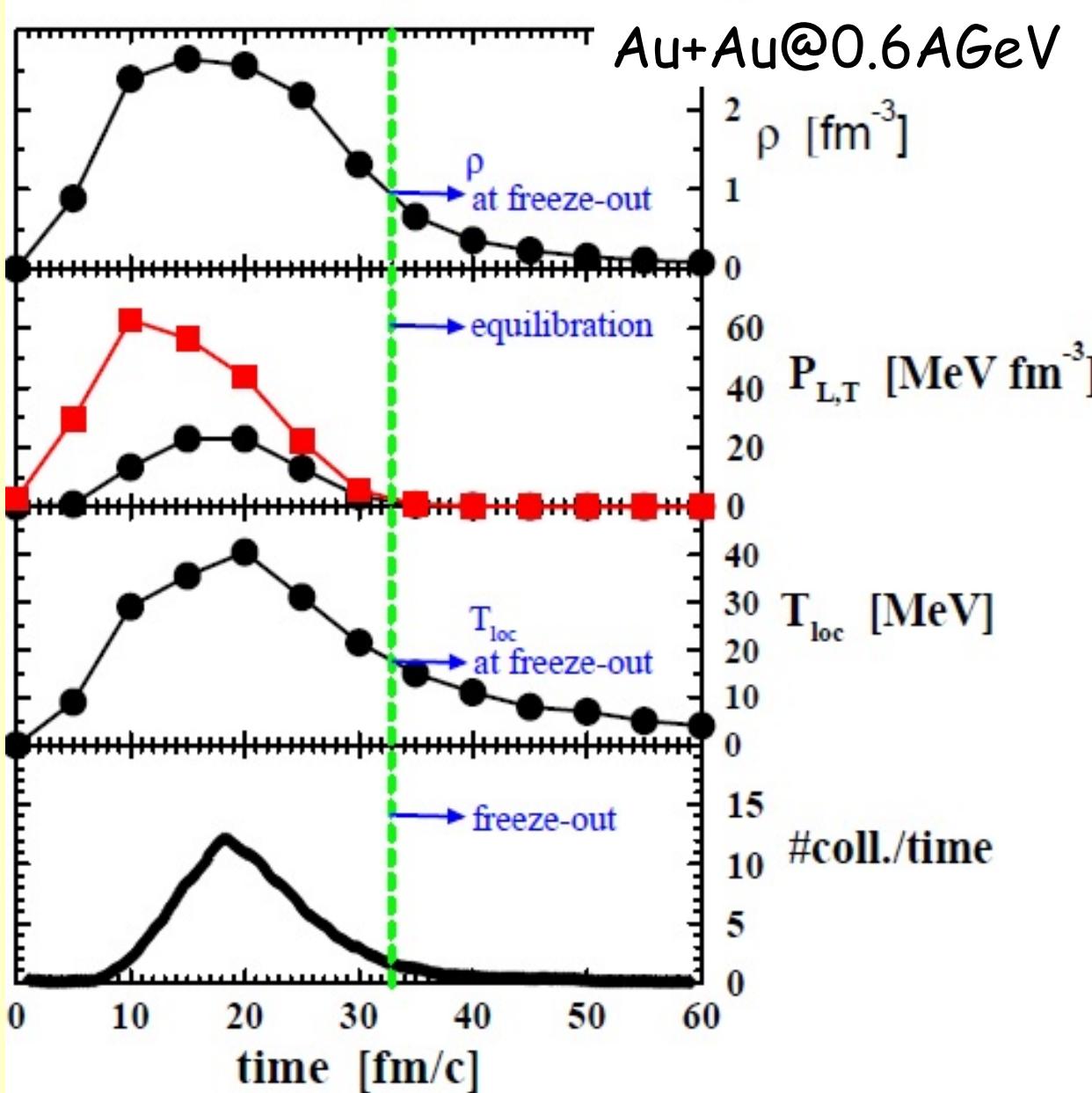


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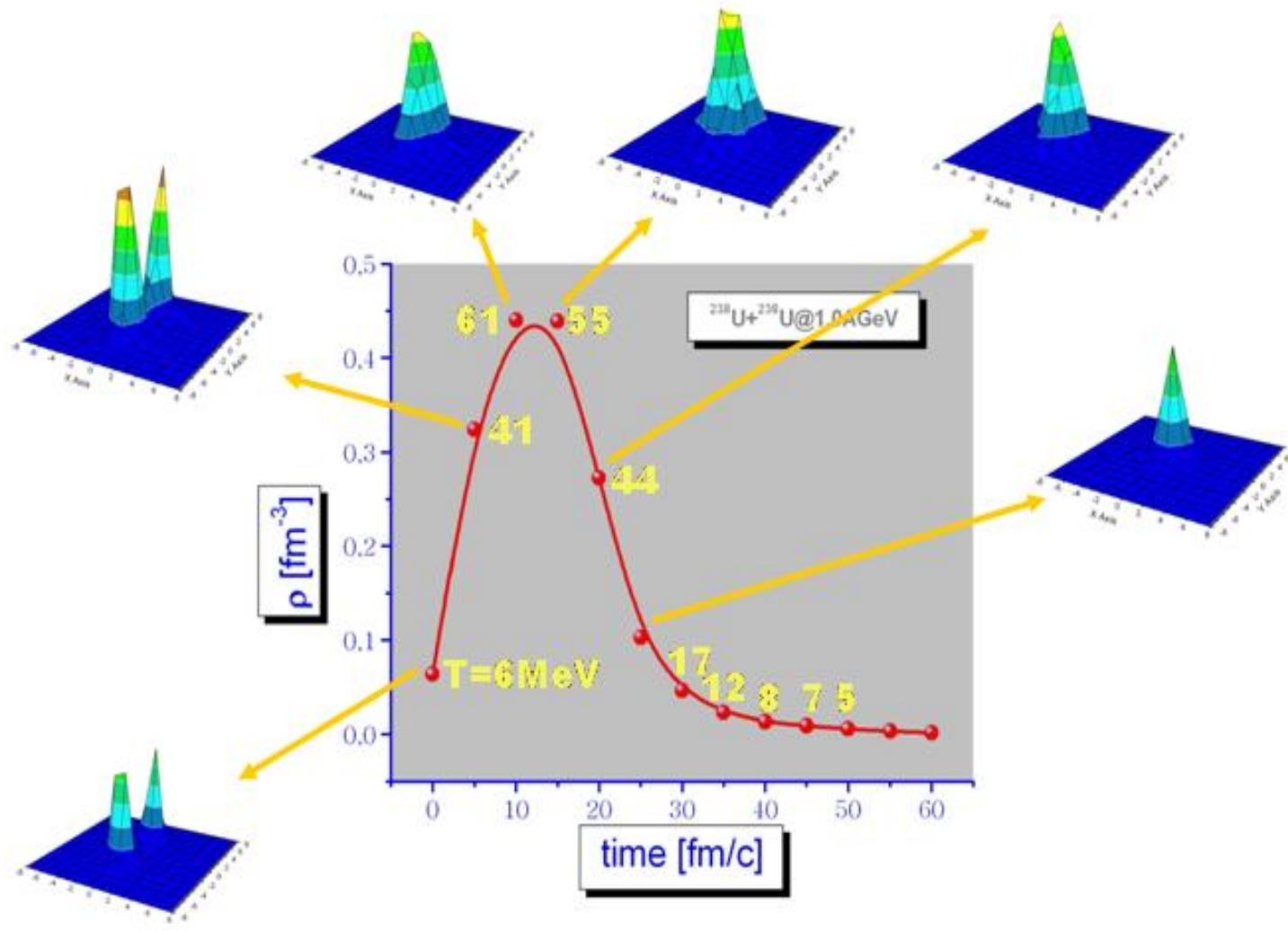
Cosmic matter in the Lab: Heavy-ion collisions



Densities & temperatures in participant matter

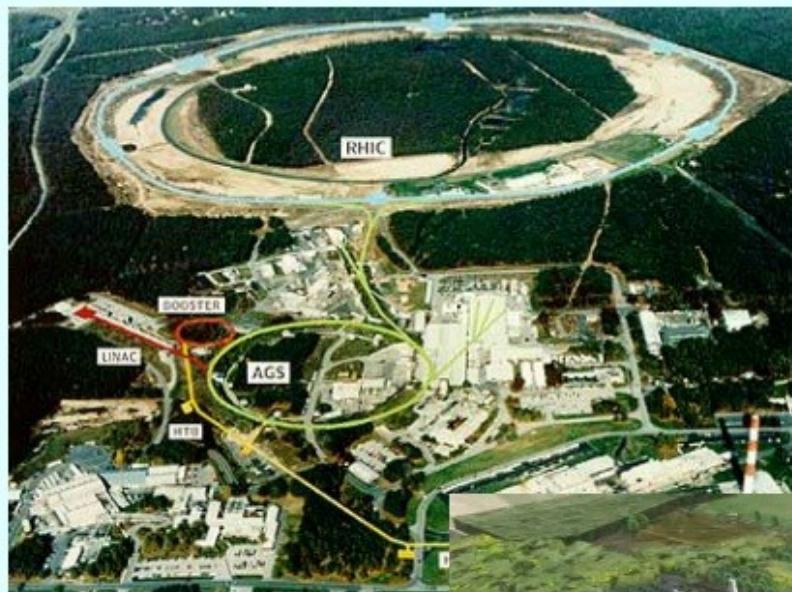


Density & temperature evolution at the center



Introduction...

Experiments on heavy-ion collision



Relativistic Heavy Ion Collider (RHIC), Brookhaven seit 2000

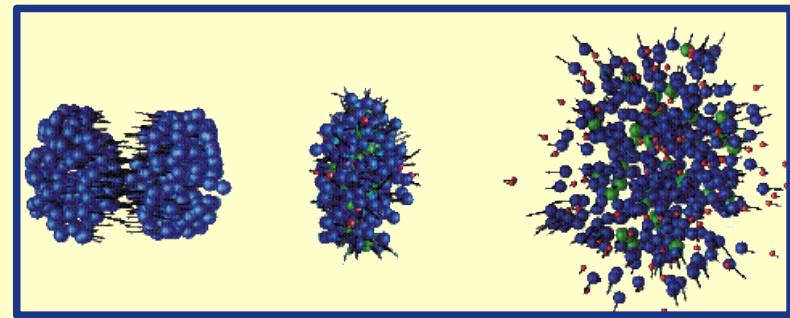
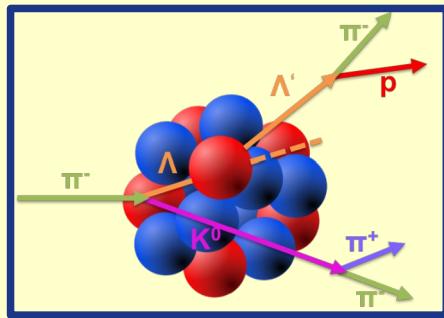
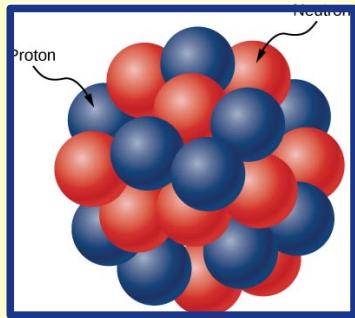


Large Hadron Collider (LHC), Cern, ab 2008



Facility for Antiproton and Ion Research (FAIR), GSI, ab 2011

Theory: Relativistic nuclear/hadron models...



Involves

Elementary particles & interactions → quantum mechanics

Particle velocities $v \sim c$ → relativistic kinematics & dynamics

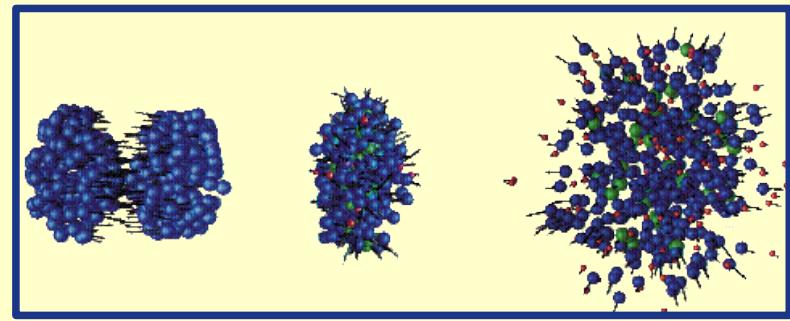
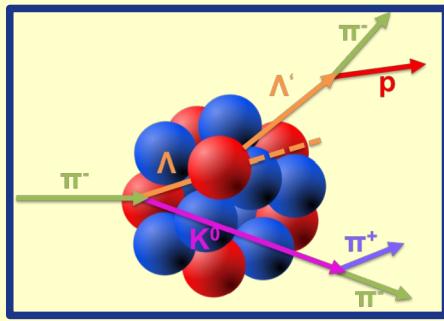
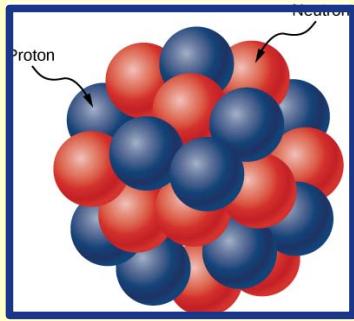
Exact treatment

Relativistic quantum transport theory of a many-body nuclear system

Practical applications...

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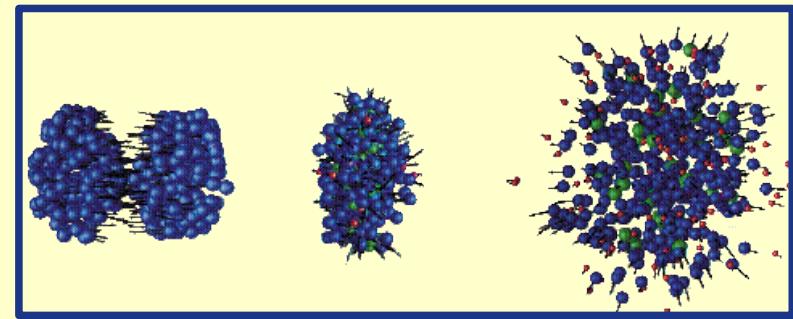
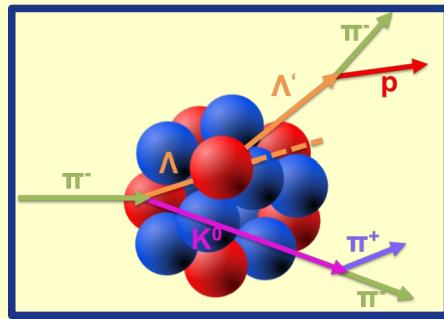
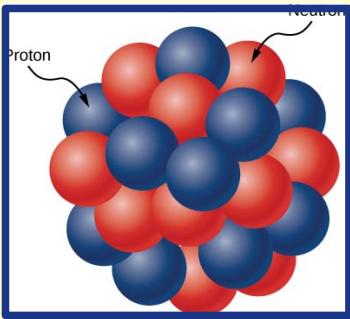
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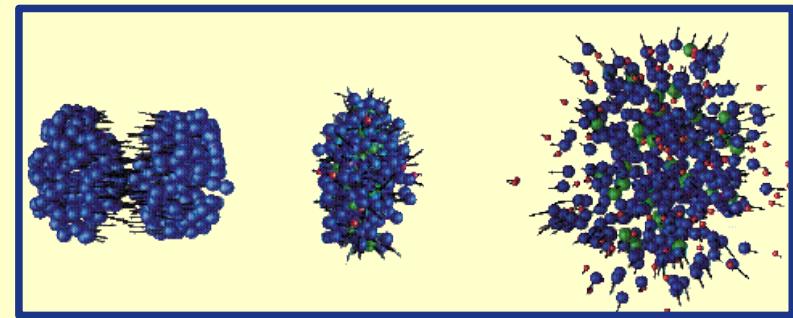
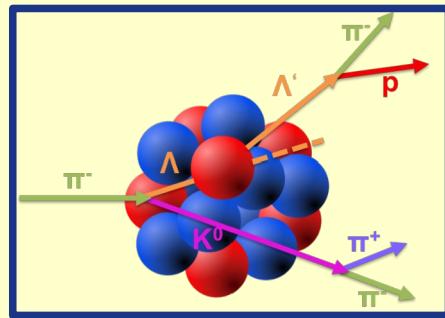
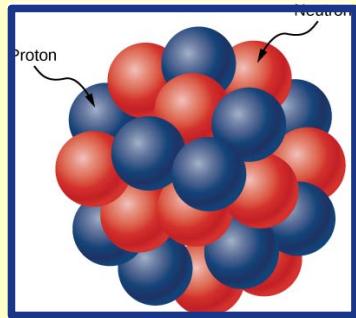
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Practical applications (on computer)

approximations !

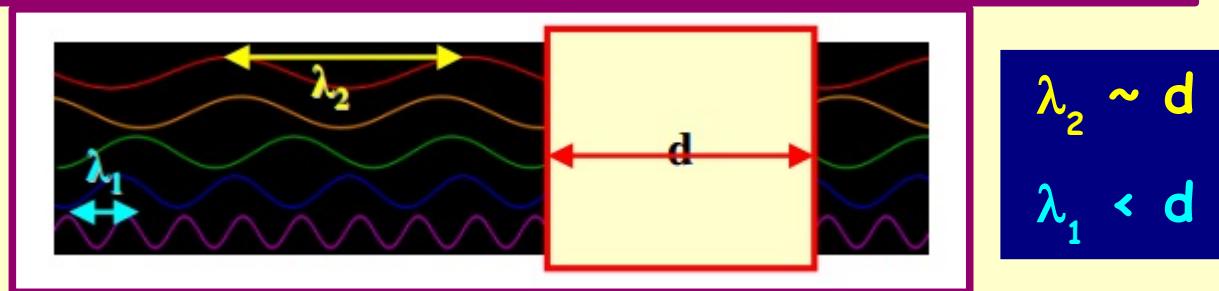
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Theory: Relativistic nuclear/hadron models...

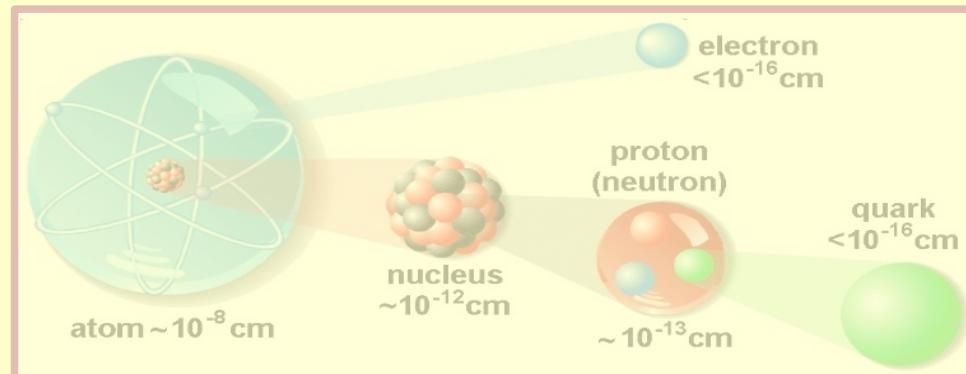
classical mechanics \leftrightarrow quantum mechanics
elem. particles (t, x) \leftrightarrow wave packets $\Psi(t, x)$

Correspondence principle (Niels Bohr, 1923):

Wave nature of particles not visible, when de Broglie wave length $\lambda = h/p$ small compared to considered scales



Considered scales here:
nuclear radius $d \sim 10^{-15} \text{ m} = 1 \text{ fm}$

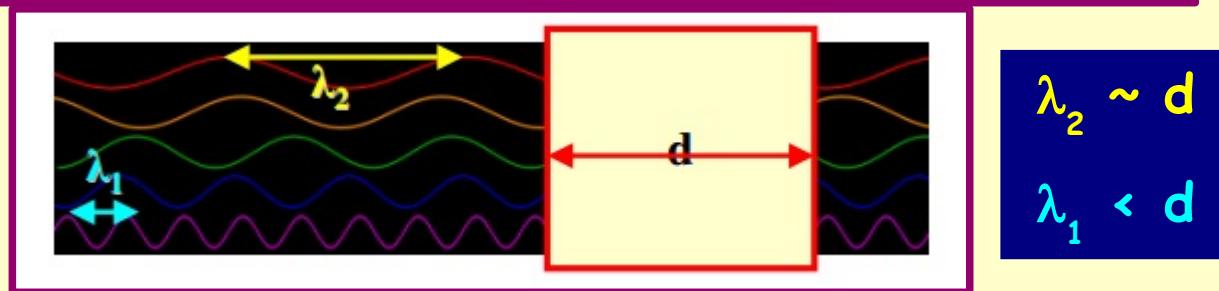


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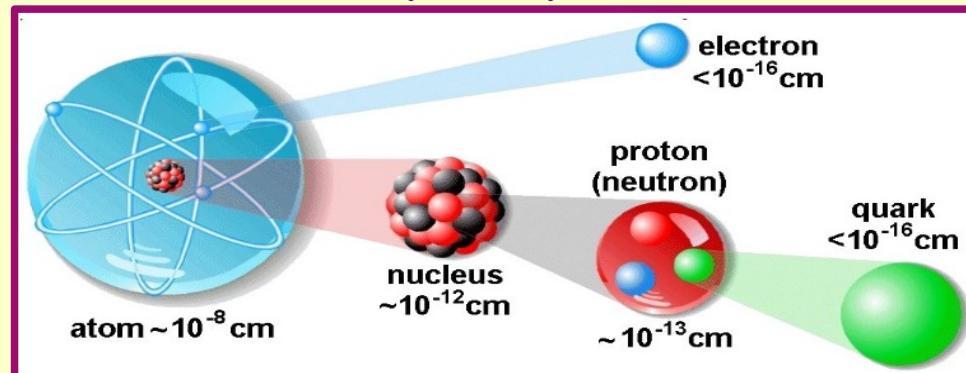
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Considered scales here:
nuclear radius $d \sim (5-10) \times 10^{-15} \text{ m} \sim 10^{-14} \text{ fm}$

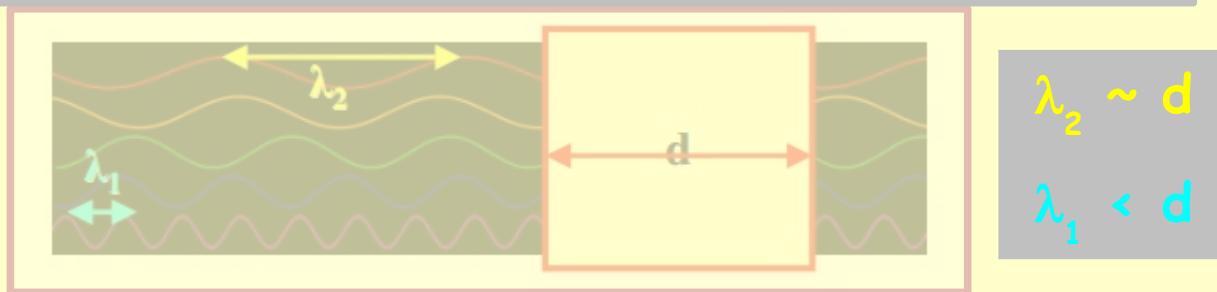


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proton E_{kin} $\lambda = h/p$

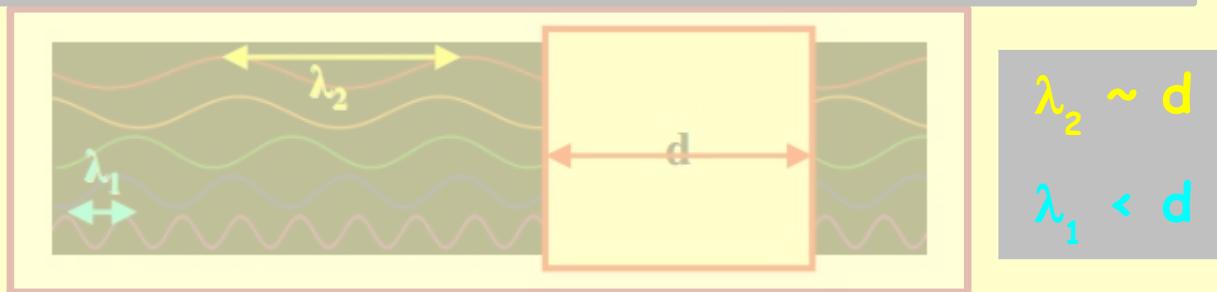
1 MeV	$2,86 \cdot 10^{-14} \text{ m}$
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} classical description possible
Tübingen, 15.11.19

Theory: Relativistic transport & covariant EoS...

Literature on (relativistic) kinetic theory...

- 1. Primer L. Boltzmann, Wien. Ber. **66** (1872) 275
- 2. Primer...L. Nordheim, Proc. R. Soc. London **A119** (1928) 689

E.A. Uehling, G.E. Uhlenbeck, Phys. Rev. **43** (1933) 552

■ Theoretical background

Non-Relativistic kinetic theory

L.P. Kadanoff, G. Baym, „Quantum Statistical Mechanics“ (Benjamin, N.Y. 1962)

Relativistic kinetic theory

S.R. de Groot, W.A. van Leeuwen, C.G. van Weert

„Relativistic kinetic theory“ (North Holland, Amsterdam, 1980)

Modern Relativistic Quantum Transport Theory

W. Botermans, R. Malfliet, Phys. Rep. **198** (1990) 115 (←difficult to understand...)

■ First applications to HIC...

P. Danielewicz, Ann. Phys. **152** (1984) 239 & 305

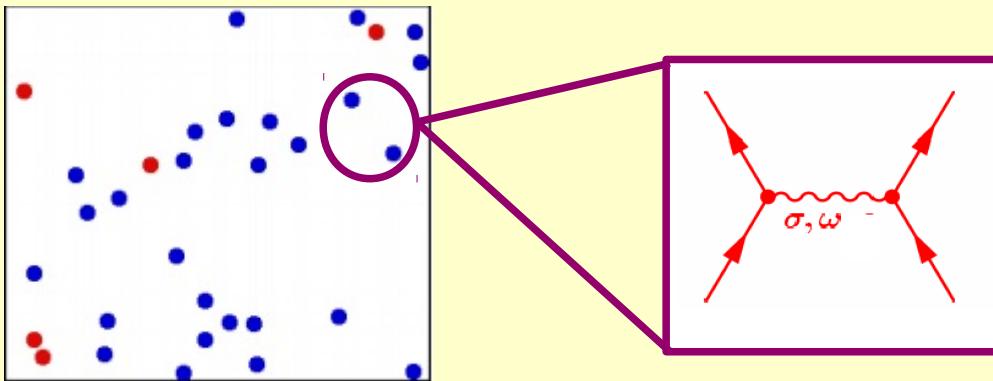
G.F. Bertsch, S. Das Gupta, Phys. Rep. **160** (1988) 189

Relativistic applications to HIC...

B. Blättel, V. Koch, U. Mosel, Rep. Prog. Phys. **56** (1993) 1

Theory: Relativistic Hadro-Dynamics...

Principle: Lagrangian formalism for fields

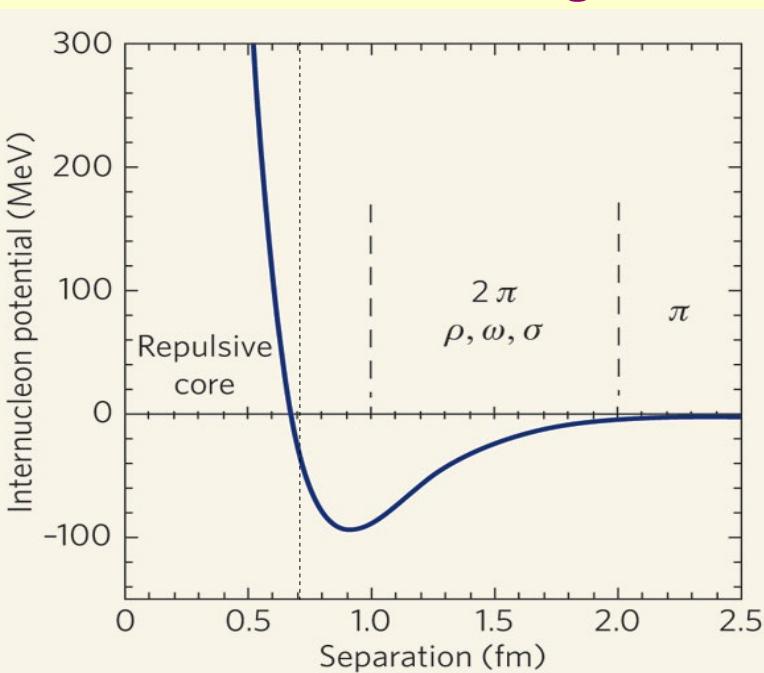


Degrees of freedom

Nucleons \rightarrow Dirac Spinor Ψ

Interaction $\rightarrow \sigma, \omega$ mesons

Interaction: short range attractive & repulsive for $r < 0.8$ fm



Yukawa's concept

screened Coulomb potential

$$V(r) = -g^2 \frac{1}{r} e^{-mr}$$

coupling strength

particle mass

Theory: Relativistic Hadro-Dynamics...

Classical mechanics

versus

relativistic field theory

Point-like particles

relativistic fields

dof: coordinates

$$\mathbf{q}_i$$

$$\varphi(t, \vec{r}) = \varphi(x)$$

Velocities

$$\dot{\mathbf{q}}_i$$

$$\dot{\varphi}^\mu(r, t, \omega)$$

Lagrange function

$$L = q_i \cdot \ddot{q}_i - V(q_i)$$

Lagrangian density

$$\mathcal{L} = \mathcal{L}(q_i, \dot{q}_i)$$

Equations of motion

$$\frac{d\mathcal{L}}{dq_i} - \frac{d}{dt} \frac{d\mathcal{L}}{d\dot{q}_i} = 0$$

Equations of motion

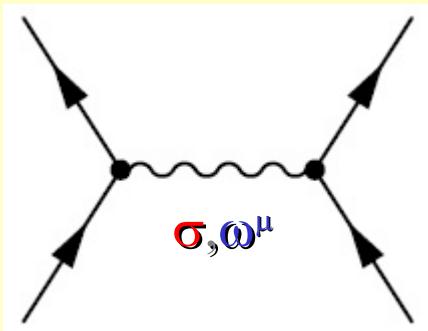
$$\frac{\partial \mathcal{L}}{\partial \varphi} - \partial_\mu \frac{\partial \mathcal{L}}{\partial (\partial_\mu \varphi)} = 0$$

Theory: Relativistic Hadro-Dynamics...

Free Lagrangian for the Dirac field (spin $\frac{1}{2}$ nucleons)

$$\mathcal{L}_{\text{free}} = \bar{\Psi} \gamma^\mu \partial_\mu \Psi - m \bar{\Psi} \Psi$$

Interaction Lagrangian: minimal coupling (analogy to electrodynamics)



$\sigma \rightarrow \text{scalar (spin 0) field}$ $\delta \rightarrow \text{scalar (spin 0) field}$
 $\omega^\mu \rightarrow \text{vector (spin 1) field}$ $\rho^\mu \rightarrow \text{vector (spin 1) field}$

$$\begin{aligned} \mathcal{L}_{\text{int}} = & g_\sigma \bar{\Psi} \Psi \sigma - g_\omega \bar{\Psi} \gamma_\mu \Psi \omega^\mu \\ & + g_\delta \bar{\Psi} \vec{\tau} \Psi \vec{\delta} - g_\rho \bar{\Psi} \gamma_\mu \vec{\tau} \Psi \vec{\rho}^\mu \end{aligned}$$

Full Lagrangian:

$$\begin{aligned} \mathcal{L} = & \bar{\Psi} \gamma_\mu (i \partial^\mu - g_\omega \omega^\mu - g_\rho \vec{\tau} \vec{\rho}^\mu) \Psi - (m - g_\sigma \sigma - g_\delta \vec{\tau} \vec{\delta}) \bar{\Psi} \Psi \\ & + \mathcal{L}_\sigma + \mathcal{L}_\omega + \mathcal{L}_\delta + \mathcal{L}_\rho \end{aligned}$$

The Non-Linear Derivative (NLD) model...

NLD Lagrangian : as in conventional RHD

$$\begin{aligned}\mathcal{L} = & \frac{1}{2} \left[\bar{\Psi} \gamma_\mu i \vec{\partial}^\mu \Psi - \bar{\Psi} i \overleftrightarrow{\partial}^\mu \gamma_\mu \Psi \right] - m \bar{\Psi} \Psi - \frac{1}{2} m_\sigma^2 \sigma^2 + \frac{1}{2} \partial_\mu \sigma \partial^\mu \sigma - U(\sigma) \\ & + \frac{1}{2} m_\omega^2 \omega_\mu \omega^\mu - \frac{1}{4} F_{\mu\nu} F^{\mu\nu} + \frac{1}{2} m_\rho^2 \vec{\rho}_\mu \vec{\rho}^\mu - \frac{1}{4} \vec{G}_{\mu\nu} \vec{G}^{\mu\nu} \\ & + \mathcal{L}_{int}\end{aligned}$$

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Interaction Lagrangian : as in conventional RHD

$$\begin{aligned}\mathcal{L}_{int} = & \frac{g_\sigma}{2} [\bar{\Psi} \Psi \sigma + \sigma \bar{\Psi} \Psi] - \frac{g_\omega}{2} [\bar{\Psi} \gamma^\mu \Psi \omega_\mu + \omega_\mu \bar{\Psi} \gamma^\mu \Psi] \\ & - \frac{g_\rho}{2} [\bar{\Psi} \gamma^\mu \vec{\tau} \Psi \vec{\rho}_\mu + \vec{\rho}_\mu \bar{\Psi} \vec{\tau} \gamma^\mu \Psi]\end{aligned}$$

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Interaction Lagrangian : as in conventional RHD + non-linear derivative operators

$$\begin{aligned}\mathcal{L}_{int} = & \frac{g_\sigma}{2} \left[\bar{\Psi} \overleftarrow{\mathcal{D}} \Psi \sigma + \sigma \bar{\Psi} \overrightarrow{\mathcal{D}} \Psi \right] - \frac{g_\omega}{2} \left[\bar{\Psi} \overleftarrow{\mathcal{D}} \gamma^\mu \Psi \omega_\mu + \omega_\mu \bar{\Psi} \gamma^\mu \overrightarrow{\mathcal{D}} \Psi \right] \\ & - \frac{g_\rho}{2} \left[\bar{\Psi} \overleftarrow{\mathcal{D}} \gamma^\mu \vec{\tau} \Psi \vec{\rho}_\mu + \vec{\rho}_\mu \bar{\Psi} \vec{\tau} \gamma^\mu \overrightarrow{\mathcal{D}} \Psi \right]\end{aligned}$$

The Non-Linear Derivative (NLD) model...

NLD Lagrangian : as in conventional RHD

$$\begin{aligned} \mathcal{L} = & \frac{1}{2} \left[\bar{\Psi} \gamma_\mu i \vec{\partial}^\mu \Psi - \bar{\Psi} i \overleftarrow{\partial}^\mu \gamma_\mu \Psi \right] - m \bar{\Psi} \Psi - \frac{1}{2} m_\sigma^2 \sigma^2 + \frac{1}{2} \partial_\mu \sigma \partial^\mu \sigma - U(\sigma) \\ & + \frac{1}{2} m_\omega^2 \omega_\mu \omega^\mu - \frac{1}{4} F_{\mu\nu} F^{\mu\nu} + \frac{1}{2} m_\rho^2 \vec{\rho}_\mu \vec{\rho}^\mu - \frac{1}{4} \vec{G}_{\mu\nu} \vec{G}^{\mu\nu} \\ & + \mathcal{L}_{int} \end{aligned}$$

Interaction Lagrangian : as in conventional RHD + non-linear derivative operators

$$\begin{aligned} \mathcal{L}_{int} = & \frac{g_\sigma}{2} \left[\bar{\Psi} \overleftarrow{\mathcal{D}} \Psi \sigma + \sigma \bar{\Psi} \overrightarrow{\mathcal{D}} \Psi \right] - \frac{g_\omega}{2} \left[\bar{\Psi} \overleftarrow{\mathcal{D}} \gamma^\mu \Psi \omega_\mu + \omega_\mu \bar{\Psi} \gamma^\mu \overrightarrow{\mathcal{D}} \Psi \right] \\ & - \frac{g_\rho}{2} \left[\bar{\Psi} \overleftarrow{\mathcal{D}} \gamma^\mu \vec{\tau} \Psi \vec{\rho}_\mu + \vec{\rho}_\mu \bar{\Psi} \vec{\tau} \gamma^\mu \overrightarrow{\mathcal{D}} \Psi \right] \end{aligned}$$

Non-linear derivative operators : Taylor expansion of partial derivatives ξ

$$\overrightarrow{\mathcal{D}} := \mathcal{D} \left(\overrightarrow{\xi} \right) = \sum_{j=0}^{n \rightarrow \infty} \frac{\partial^j}{\partial \overrightarrow{\xi}^j} \mathcal{D} \Big|_{\overrightarrow{\xi} \rightarrow 0} \frac{\overrightarrow{\xi}^j}{j!} \quad \overrightarrow{\xi} = - \frac{v^\alpha i \vec{\partial}_\alpha}{\Lambda}$$

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Non-linear derivative operators : Taylor expansion of partial derivatives ξ

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cut-off, will regulate the high-momentum tail of RMF fields

NLD: The formalism...

NLD Lagrangian: contains higher field derivatives: $\mathcal{L}(\varphi_r, \partial_{\alpha_1}\varphi_r, \partial_{\alpha_1\alpha_2}\varphi_r, \dots, \partial_{\alpha_1\dots\alpha_n}\varphi_r)$

→ Generalized Euler-Lagrange equations:

$$\frac{\partial \mathcal{L}}{\partial \varphi_r} + \sum_{i=1}^n (-)^i \partial_{\alpha_1\dots\alpha_i} \frac{\partial \mathcal{L}}{\partial (\partial_{\alpha_1\dots\alpha_i} \varphi_r)} = 0$$

→ Generalized Noether-Theorem: conserved current

$$J^\mu = -i [\mathcal{K}_r^\mu \varphi_r + \mathcal{K}_r^{\mu\sigma_1} \partial_{\sigma_1} \varphi_r + \mathcal{K}_r^{\mu\sigma_1\sigma_2} \partial_{\sigma_1\sigma_2} \varphi_r + \dots + \mathcal{K}_r^{\mu\sigma_1\dots\sigma_n} \partial_{\sigma_1\dots\sigma_n} \varphi_r]$$

with the following tensors

$$\begin{aligned} \mathcal{K}_r^{\mu\nu\dots} &:= \frac{\partial \mathcal{L}}{\partial (\partial_\mu \varphi_r \partial_\nu \varphi_r \dots)} \\ &= \frac{\partial \mathcal{L}}{\partial (\partial_\mu \varphi_r \partial_\nu \varphi_r \dots)} \end{aligned}$$

NLD: The formalism...

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with the following tensors

A handwritten note showing a series of terms involving partial derivatives of the Lagrangian with respect to field derivatives. The terms are enclosed in a blue oval.

infinite series resp. to higher-order field derivatives, but...

NLD: The formalism...

NLD Lagrangian: contains higher field derivatives: $\mathcal{L}(\varphi_r, \partial_{\alpha_1}\varphi_r, \partial_{\alpha_1\alpha_2}\varphi_r, \dots, \partial_{\alpha_1\dots\alpha_n}\varphi_r)$

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with the following tensors

All infinite series can be resummed to compact expressions !

NLD field equations...

→ Dirac equation for nucleons $[\gamma_\mu(i\partial^\mu - \Sigma^\mu) - (m - \Sigma_s)]\Psi = 0$ with selfenergies

$$\Sigma^\mu = g_\omega \omega^\mu \vec{\mathcal{D}} + g_\rho \vec{\tau} \cdot \vec{\rho}^\mu \vec{\mathcal{D}} + \dots$$

$$\Sigma_s = g_\sigma \sigma \vec{\mathcal{D}} + \dots \quad (\text{up to terms containing derivatives of the meson fields})$$

→ Meson field equations:

$$\partial_\alpha \partial^\alpha \sigma + m_\sigma^2 \sigma + \frac{\partial U}{\partial \sigma} = \frac{1}{2} g_\sigma [\bar{\Psi} \overset{\leftarrow}{\mathcal{D}} \Psi + \bar{\Psi} \overset{\rightarrow}{\mathcal{D}} \Psi],$$

$$\partial_\mu F^{\mu\nu} + m_\omega^2 \omega^\nu = \frac{1}{2} g_\omega [\bar{\Psi} \overset{\leftarrow}{\mathcal{D}} \gamma^\nu \Psi + \bar{\Psi} \gamma^\nu \overset{\rightarrow}{\mathcal{D}} \Psi],$$

$$\partial_\mu \vec{G}^{\mu\nu} + m_\rho^2 \vec{\rho}^\nu = \frac{1}{2} g_\rho [\bar{\Psi} \overset{\leftarrow}{\mathcal{D}} \gamma^\nu \vec{\tau} \Psi + \bar{\Psi} \vec{\tau} \gamma^\nu \overset{\rightarrow}{\mathcal{D}} \Psi]$$

→ Noether current (energy-momentum tensor):

$$J^\mu = \bar{\Psi} \gamma^\mu \Psi - \frac{1}{2} g_\sigma [\bar{\Psi} \overset{\leftarrow}{\mathcal{D}}^\mu \Psi - \bar{\Psi} \overset{\rightarrow}{\mathcal{D}}^\mu \Psi] \sigma + \frac{1}{2} g_\omega [\bar{\Psi} \overset{\leftarrow}{\mathcal{D}}^\mu \gamma^\alpha \Psi - \bar{\Psi} \gamma^\alpha \overset{\rightarrow}{\mathcal{D}}^\mu \Psi] \omega_\alpha \\ + \frac{1}{2} g_\rho [\bar{\Psi} \overset{\leftarrow}{\mathcal{D}}^\mu \gamma^\alpha \vec{\tau} \Psi - \bar{\Psi} \gamma^\alpha \overset{\rightarrow}{\mathcal{D}}^\mu \vec{\tau} \Psi] \vec{\rho}_\alpha + \dots$$

RMF approach to infinite asymmetric nuclear matter...

→ Plane wave Ansatz for Ψ and $\bar{\Psi}$ and $\mathcal{D} = \gamma^0 \gamma^1 \gamma^2 \gamma^3$ with $\xi = -\frac{v_\alpha p^\alpha}{\Lambda}$

$$\Sigma_{vi}^\mu = g_\omega \omega^\mu \mathcal{D} + g_\rho \tau_i \rho^\mu \mathcal{D}, \quad \Sigma_{si} = g_\sigma \sigma \mathcal{D}$$

$$m_\sigma^2 \sigma + \frac{\partial U}{\partial \sigma} = g_\sigma \sum_{i=p,n} \langle \bar{\Psi}_i \mathcal{D} \Psi_i \rangle = g_\sigma \rho_s$$

$$m_\omega^2 \omega = g_\omega \sum_{i=p,n} \langle \bar{\Psi}_i \gamma^0 \mathcal{D} \Psi_i \rangle = g_\omega \rho_0$$

$$m_\rho^2 \rho = g_\rho \sum_{i=p,n} \tau_i \langle \bar{\Psi}_i \gamma^0 \mathcal{D} \Psi_i \rangle = g_\rho \rho_I .$$

$$T^{\mu\nu} = \sum_{i=p,n} \frac{\kappa}{(2\pi)^3} \int_{|\vec{p}| \leq p_{F_i}} d^3 p \frac{\Pi_i^\mu p^\nu}{\Pi_i^0} - g^{\mu\nu} \langle \mathcal{L} \rangle$$

$$\Gamma^\mu = \gamma^\mu \gamma^5 \gamma^1 \gamma^2 \gamma^3 \gamma^0 \gamma^4 \gamma^6 \gamma^7 \gamma^8 \gamma^9 \gamma^0 \gamma^1 \gamma^2 \gamma^3 \gamma^4 \gamma^5 \gamma^6 \gamma^7 \gamma^8 \gamma^9$$

RMF approach to infinite asymmetric nuclear matter...

→ Plane wave Ansatz for Ψ and $\bar{\Psi}$ res and $\mathcal{D} = \sigma + i\omega\gamma^0$ with $\xi = -\frac{v_\alpha p^\alpha}{\Lambda}$

$$\Sigma_{vi}^\mu = g_\omega \omega^\mu \mathcal{D} + g_\rho \tau_i \rho^\mu \mathcal{D}, \quad \Sigma_{si} = g_\sigma \sigma \mathcal{D}$$

meson-field equations

$$m_\sigma^2 \sigma + \frac{\partial U}{\partial \sigma} = g_\sigma \sum_{i=p,n} \left\langle \bar{\Psi}_i \mathcal{D} \Psi_i \right\rangle = g_\sigma \rho_s$$

$$m_\omega^2 \omega = g_\omega \sum_{i=p,n} \left\langle \bar{\Psi}_i \gamma^0 \mathcal{D} \Psi_i \right\rangle = g_\omega \rho_0$$

$$m_\rho^2 \rho = g_\rho \sum_{i=p,n} \tau_i \left\langle \bar{\Psi}_i \gamma^0 \mathcal{D} \Psi_i \right\rangle = g_\rho \rho_I$$

Equation of State (EoS)

$$\mathcal{E} = \frac{1}{2} \int d^3 p \left[\frac{1}{2} \vec{p}^2 + \frac{1}{2} \vec{p}^2 \left(\frac{\vec{\Pi}_i \cdot \vec{p}}{\Pi_i^0} \right)^2 + \frac{1}{2} \vec{p}^2 \left(\frac{\vec{\Pi}_i \cdot \vec{p}}{\Pi_i^0} \right)^2 \right]$$

$$P = \frac{1}{3} \sum_{i=p,n} \frac{\kappa}{(2\pi)^3} \int_{|\vec{p}| \leq p_{F_i}} d^3 p \frac{\vec{\Pi}_i \cdot \vec{p}}{\Pi_i^0} + \langle \mathcal{L} \rangle$$

Features of NLD model...

$$\Sigma_{vi}^\mu = g_\omega \omega^\mu \mathcal{D} + g_\rho \tau_i \rho^\mu \mathcal{D}, \quad \Sigma_{si} = g_\sigma \sigma \mathcal{D}$$

meson-field equations

$$m_\sigma^2 \sigma + \frac{\partial U}{\partial \sigma} = g_\sigma \sum_{i=p,n} \left\langle \bar{\Psi}_i \mathcal{D} \Psi_i \right\rangle = g_\sigma \rho_s$$

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cut-off Λ regulates

1) DD & MD of selfenergies

Equation of State (EoS)

$$\epsilon = \frac{1}{3} \sum_{i=p,n} \frac{\kappa}{(2\pi)^3} \int d^3 p \frac{\vec{\Pi}_i \cdot \vec{p}}{\Pi_i^0} + \langle \mathcal{L} \rangle$$

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cut-off Λ regulates

1) DD & MD of selfenergies

2) DD of meson-field sources
(particularly for ω -field)

Equation of State (EoS)

$$\epsilon = \frac{1}{3} \sum_{i=p,n} \frac{\kappa}{(2\pi)^3} \int d^3 p \frac{\vec{\Pi}_i \cdot \vec{p}}{\Pi_i^0} + \langle \mathcal{L} \rangle$$

$$P = \frac{1}{3} \sum_{i=p,n} \frac{\kappa}{(2\pi)^3} \int_{|\vec{p}| \leq p_{F_i}} d^3 p \frac{\vec{\Pi}_i \cdot \vec{p}}{\Pi_i^0} + \langle \mathcal{L} \rangle$$

Features of NLD model...

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meson-field equations

$$m_\sigma^2 \sigma + \frac{\partial U}{\partial \sigma} = g_\sigma \sum_{i=p,n} \left\langle \bar{\Psi}_i \mathcal{D} \Psi_i \right\rangle = g_\sigma \rho_s$$

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$$m_\rho^2 \rho = g_\rho \sum_{i=p,n} \tau_i \left\langle \bar{\Psi}_i \gamma^0 \mathcal{D} \Psi_i \right\rangle = g_\rho \rho_I$$

Equation of State (EoS)

$$\epsilon = \frac{1}{3} \sum_{i=p,n} \frac{\kappa}{(2\pi)^3} \int d^3 p \frac{\vec{\Pi}_i \cdot \vec{p}}{\Pi_i^0} + \langle \mathcal{L} \rangle$$

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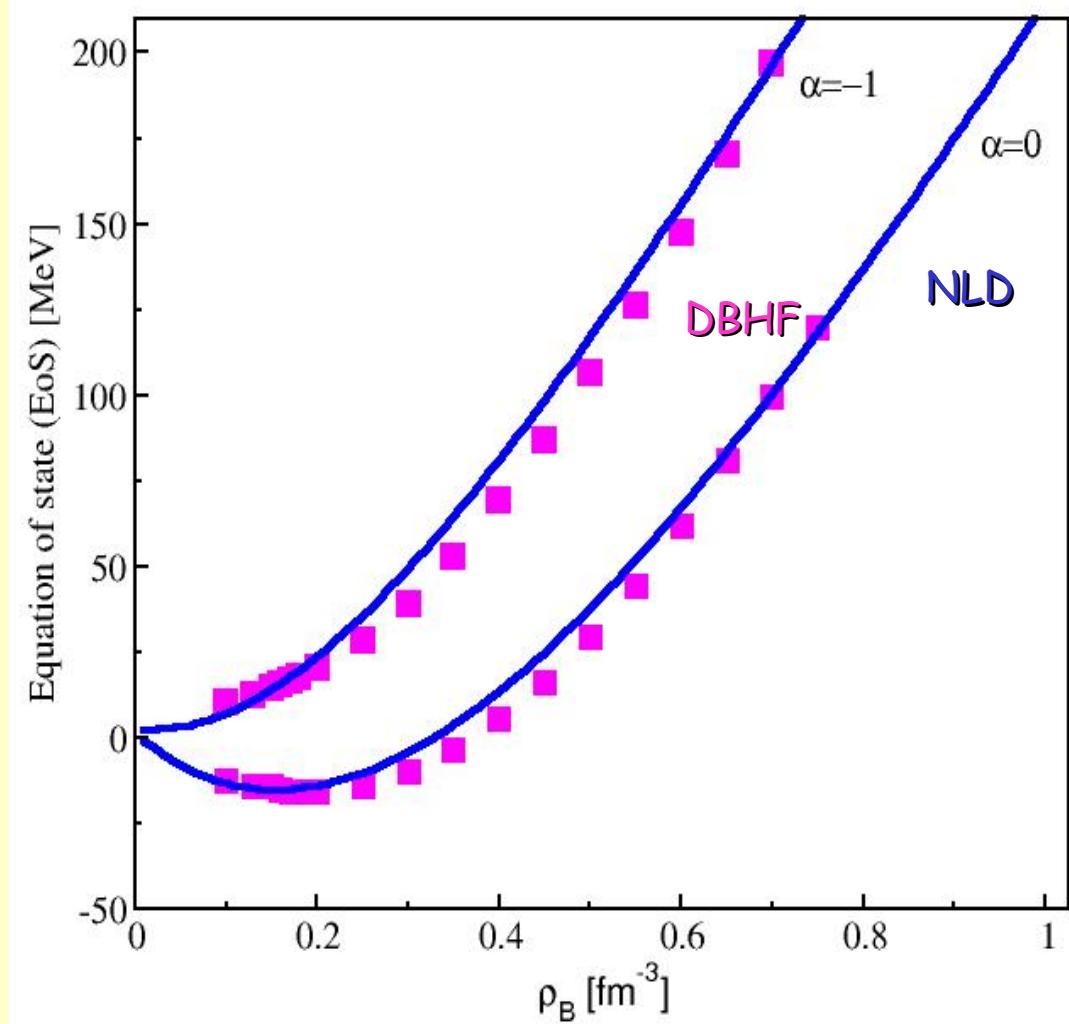
cut-off Λ regulates

1) DD & MD of selfenergies

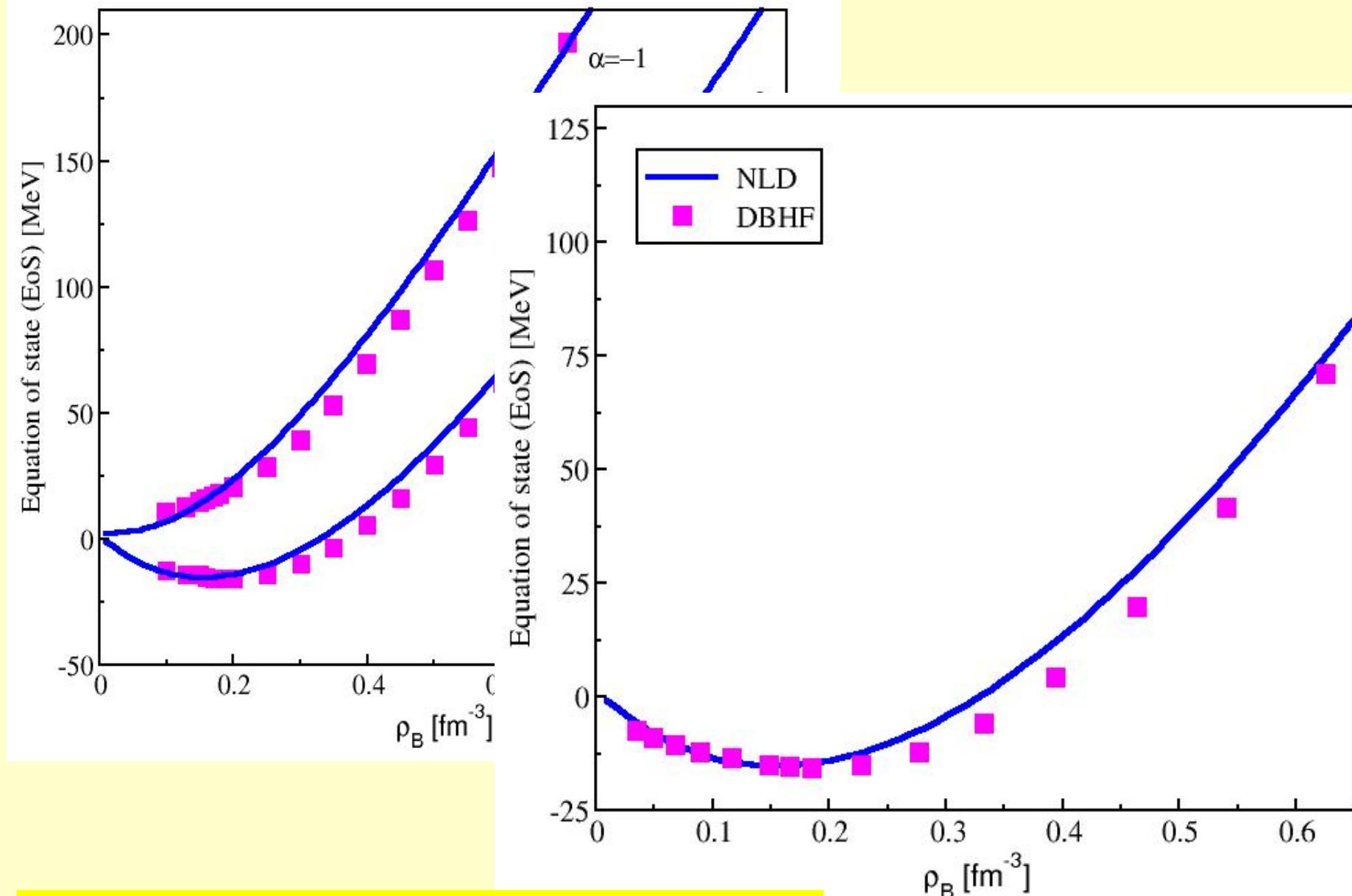
2) DD of meson-field sources
(particularly for ω -field)

3) fully thermodynamic consistent
(important for neutron stars)

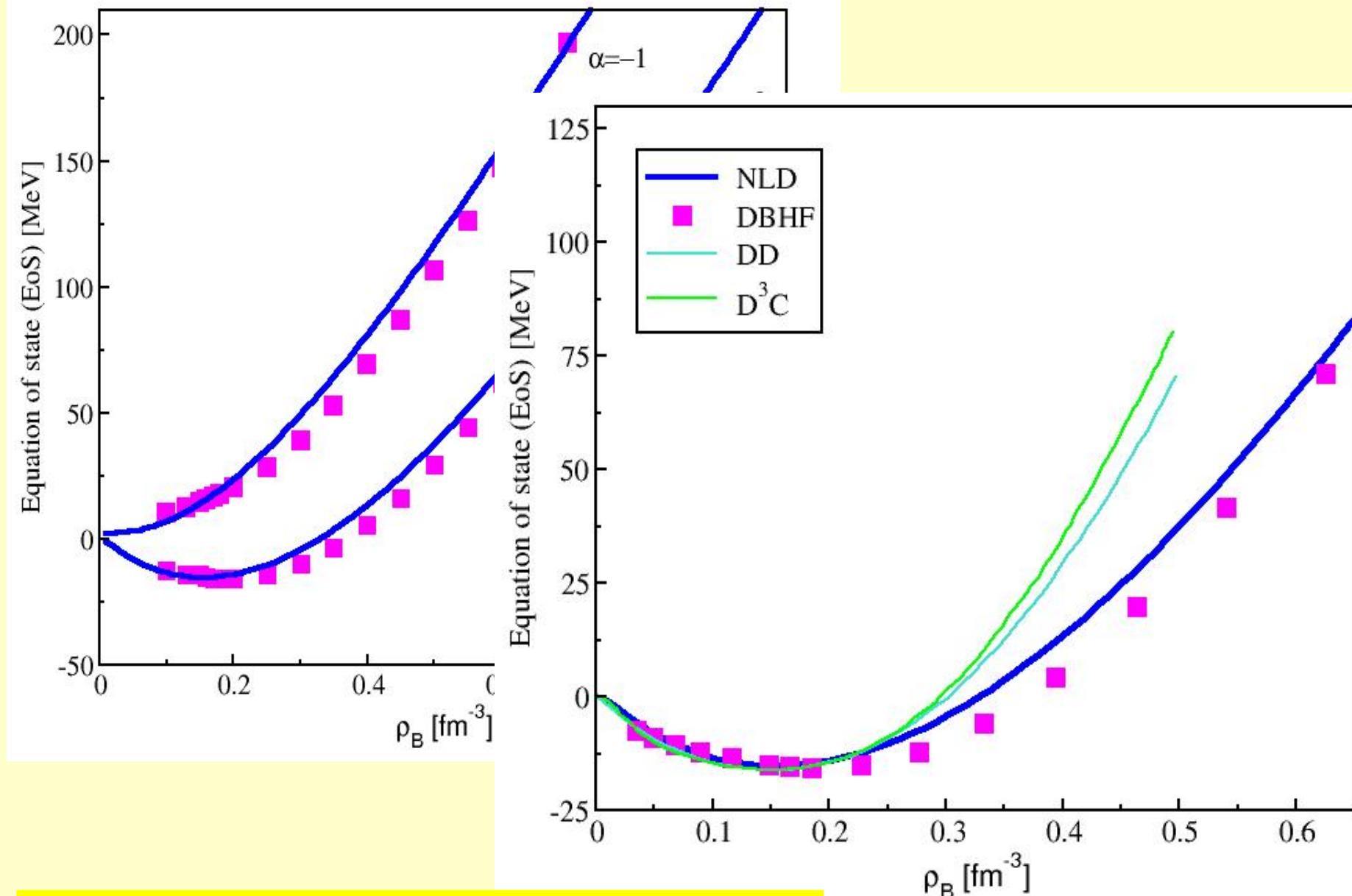
Nuclear EoS...



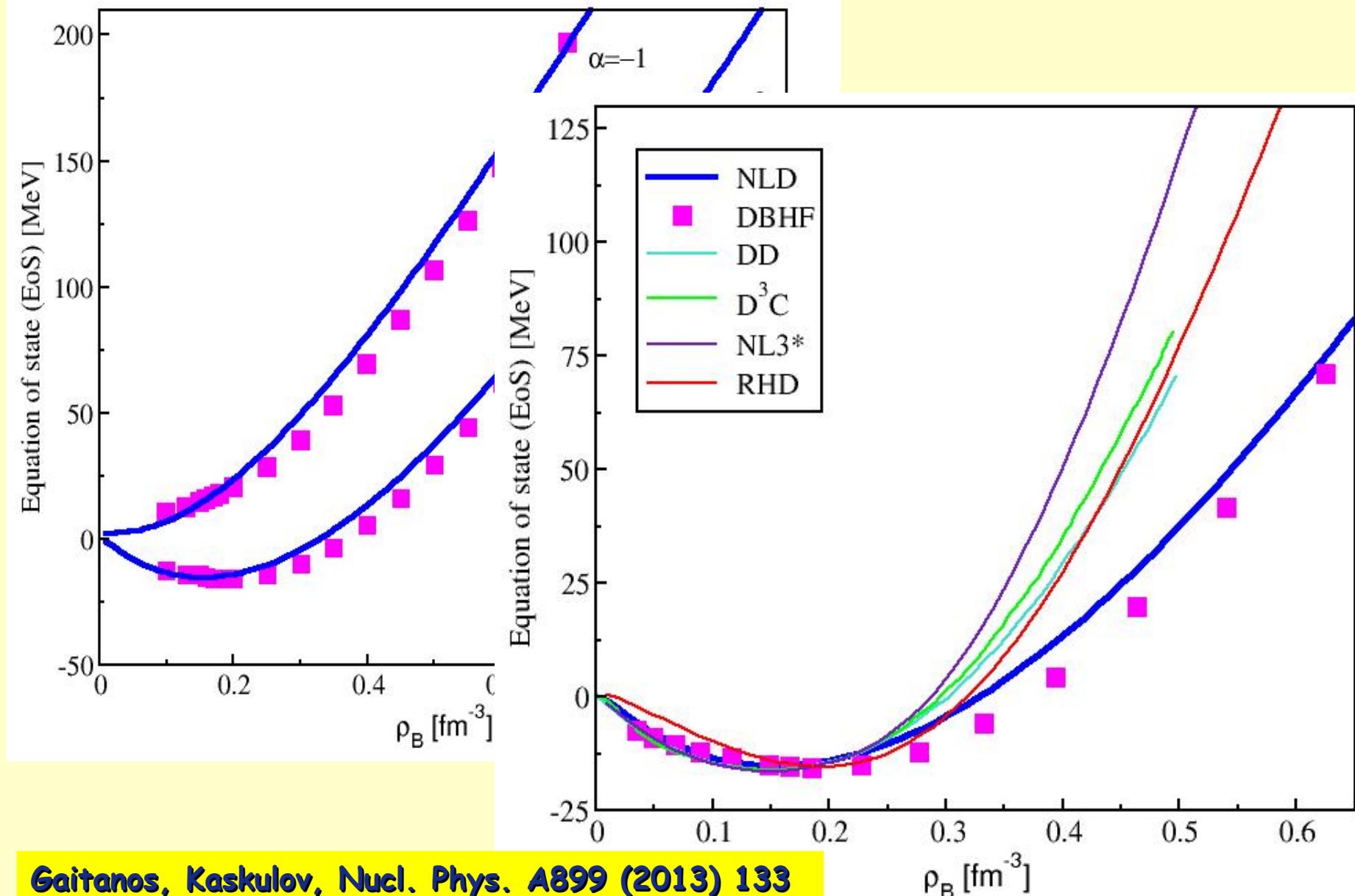
Nuclear EoS...



Nuclear EoS...



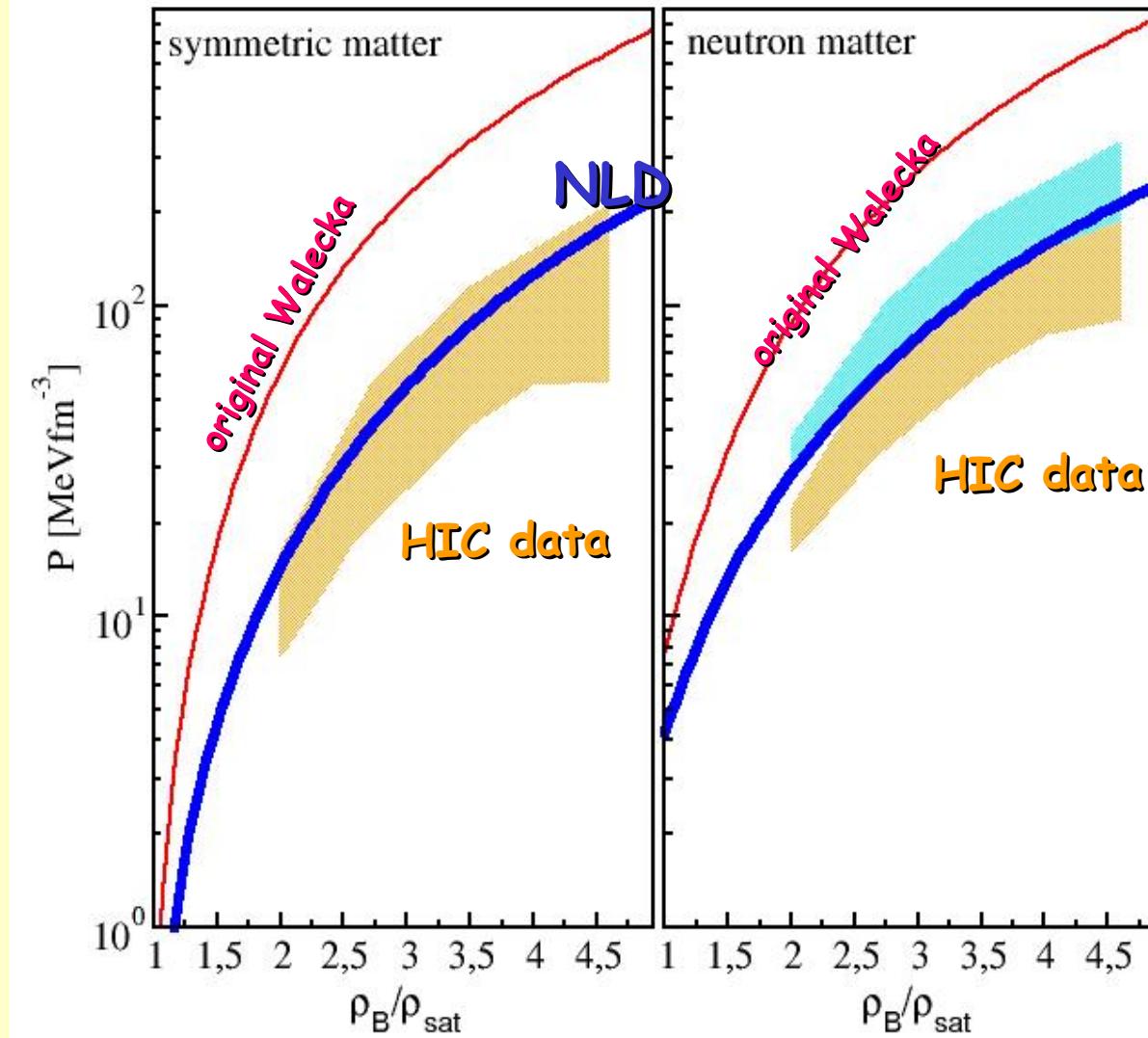
Nuclear EoS...



Gaitanos, Kaskulov, Nucl. Phys. A899 (2013) 133

Gaitanos, Kaskulov, Nucl. Phys. A940 (2015) 181

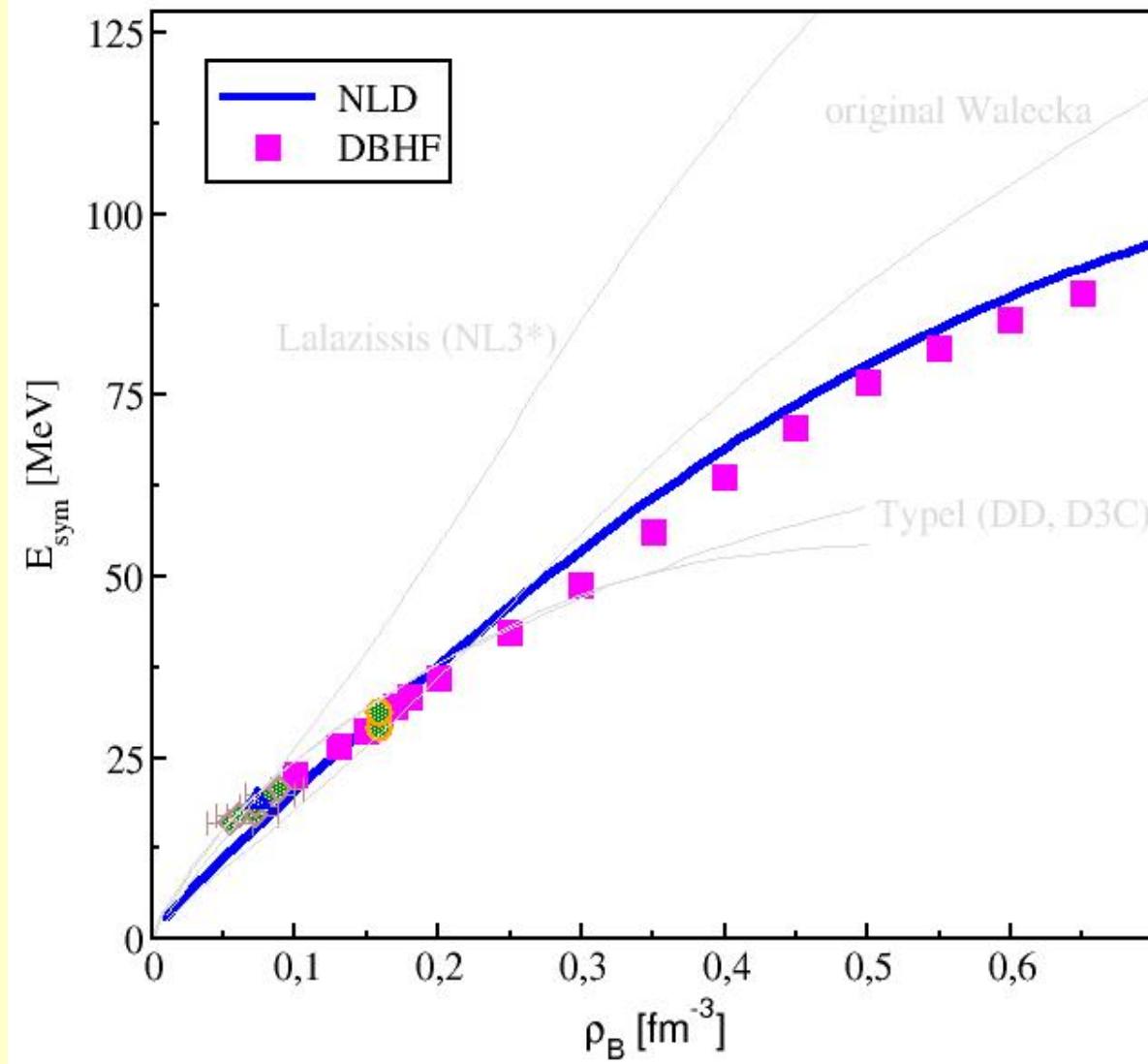
NLD results: EoS...



NLD

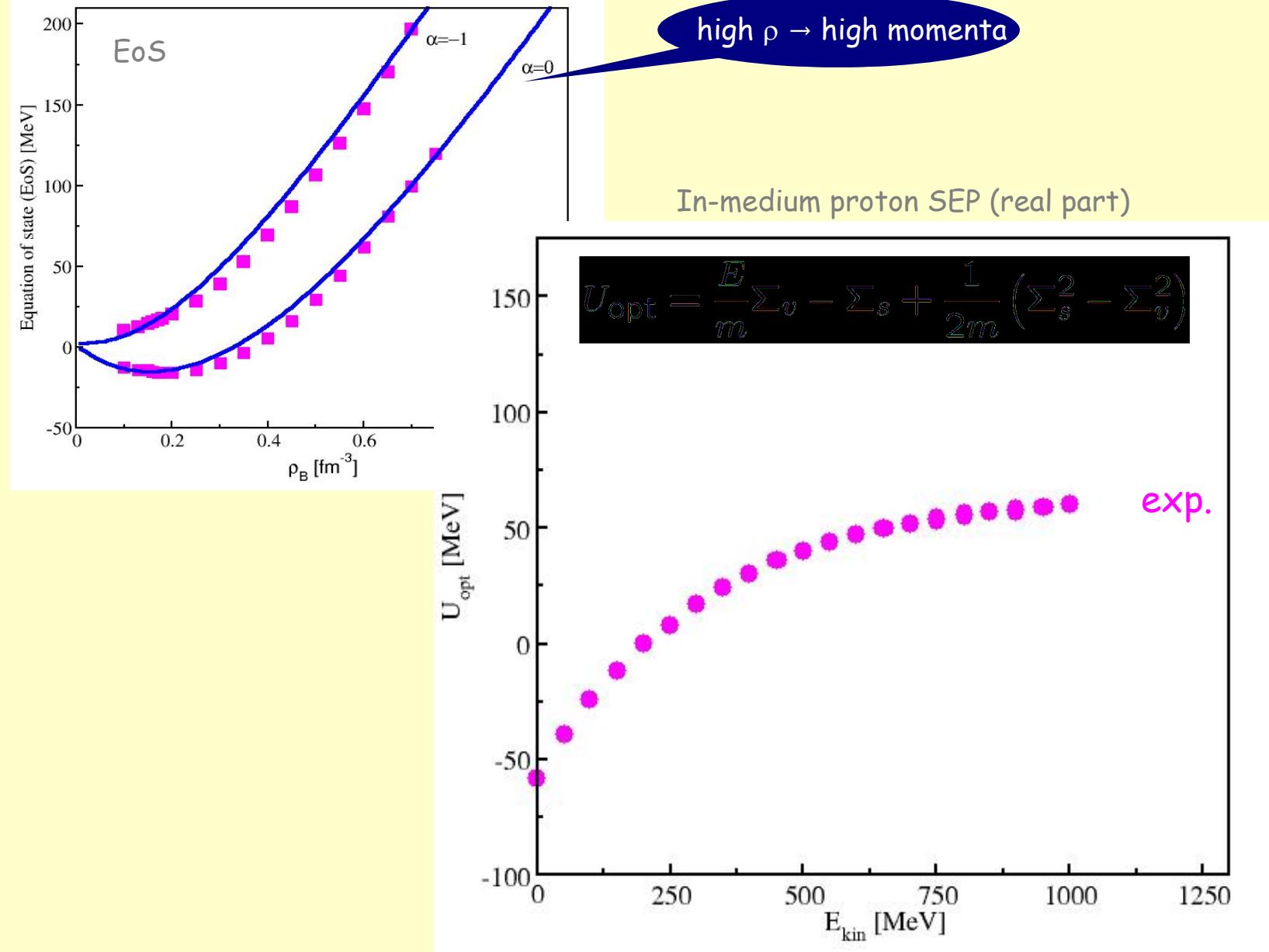
NLD model is consistent with HIC phenomenology

NLD results: symmetry energy...

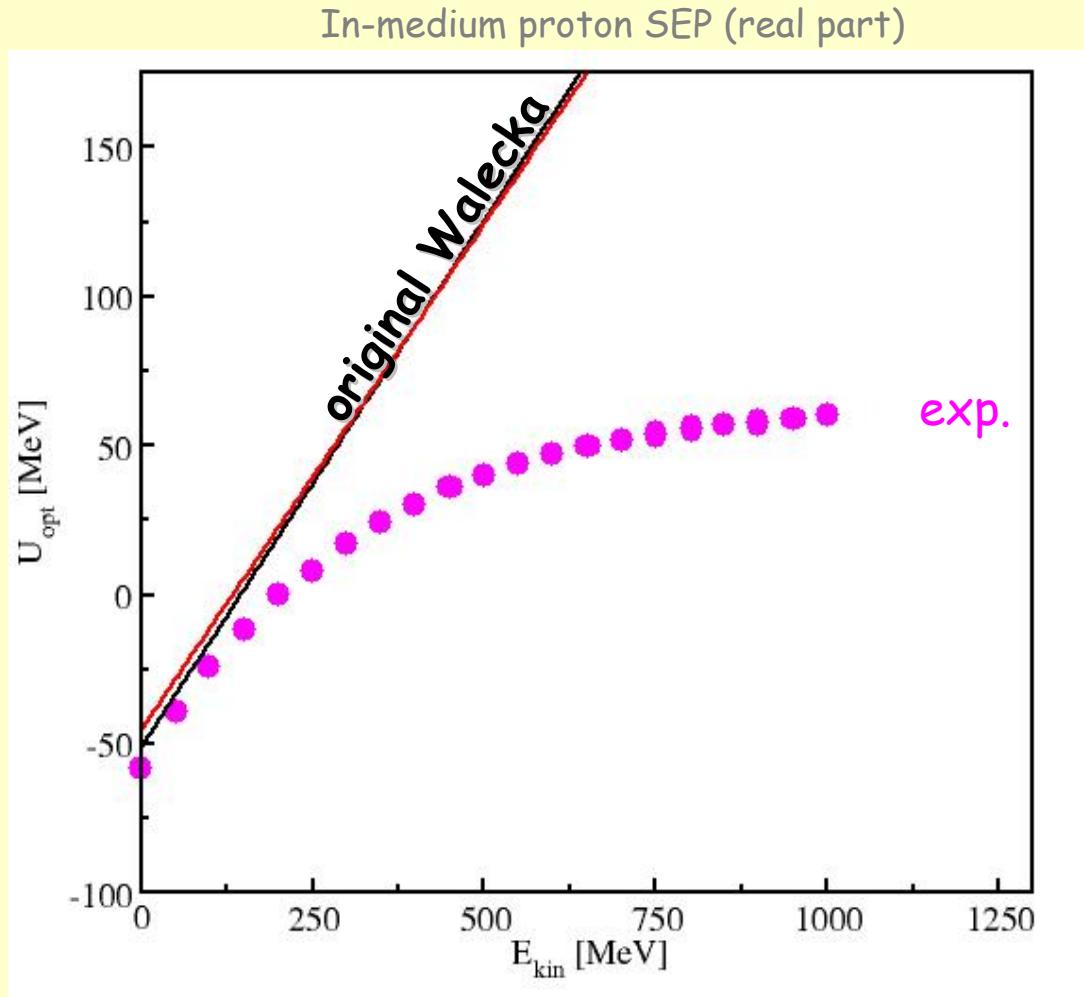


Remarkable agreement with microscopic DBHF

NLD results: MD & optical potentials...

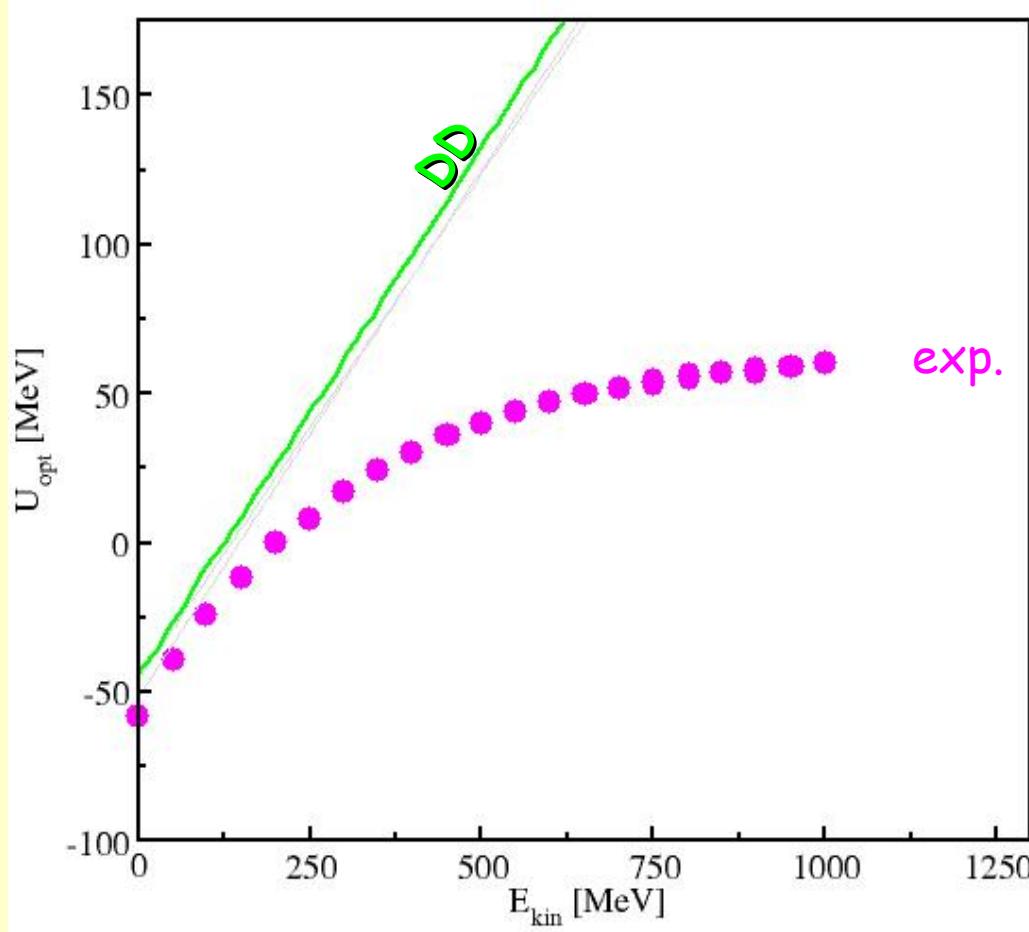


NLD results: MD & optical potentials...

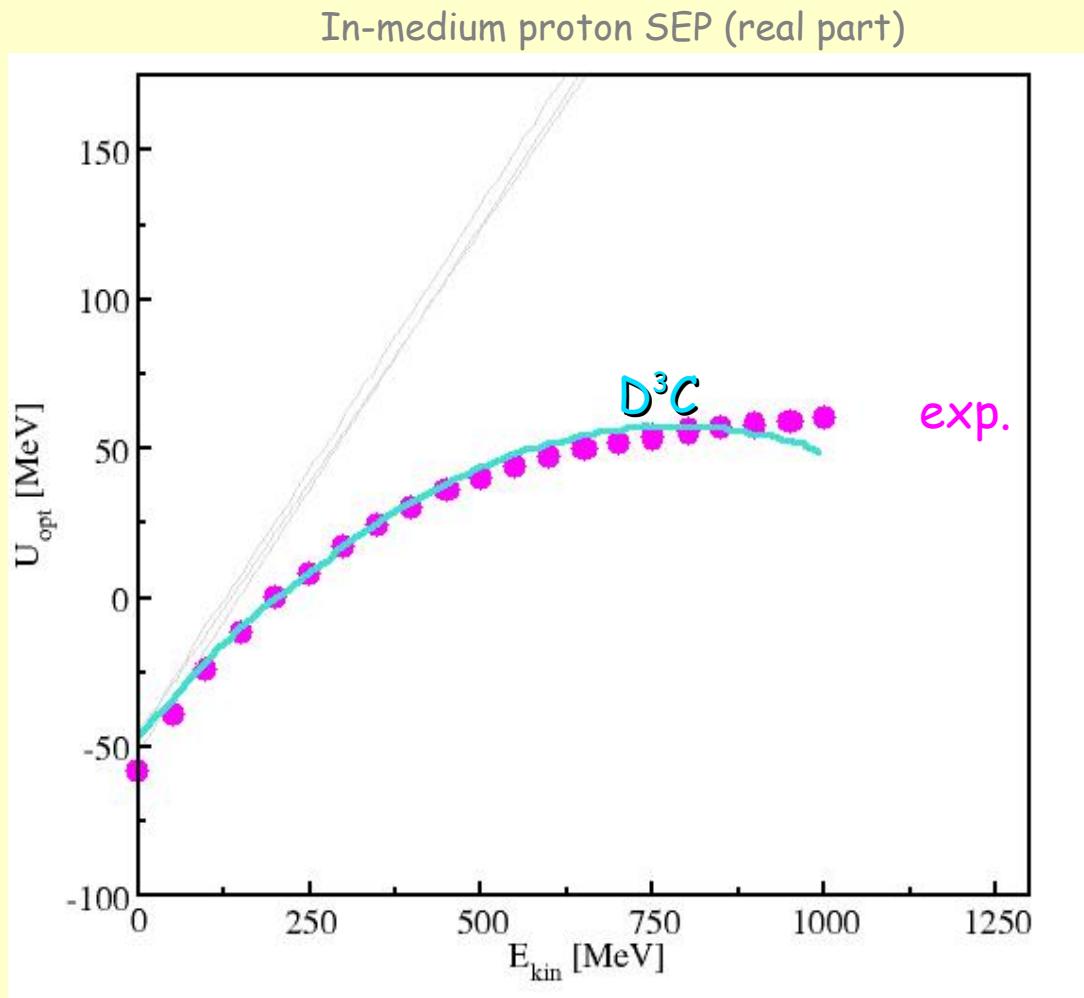


NLD results: MD & optical potentials...

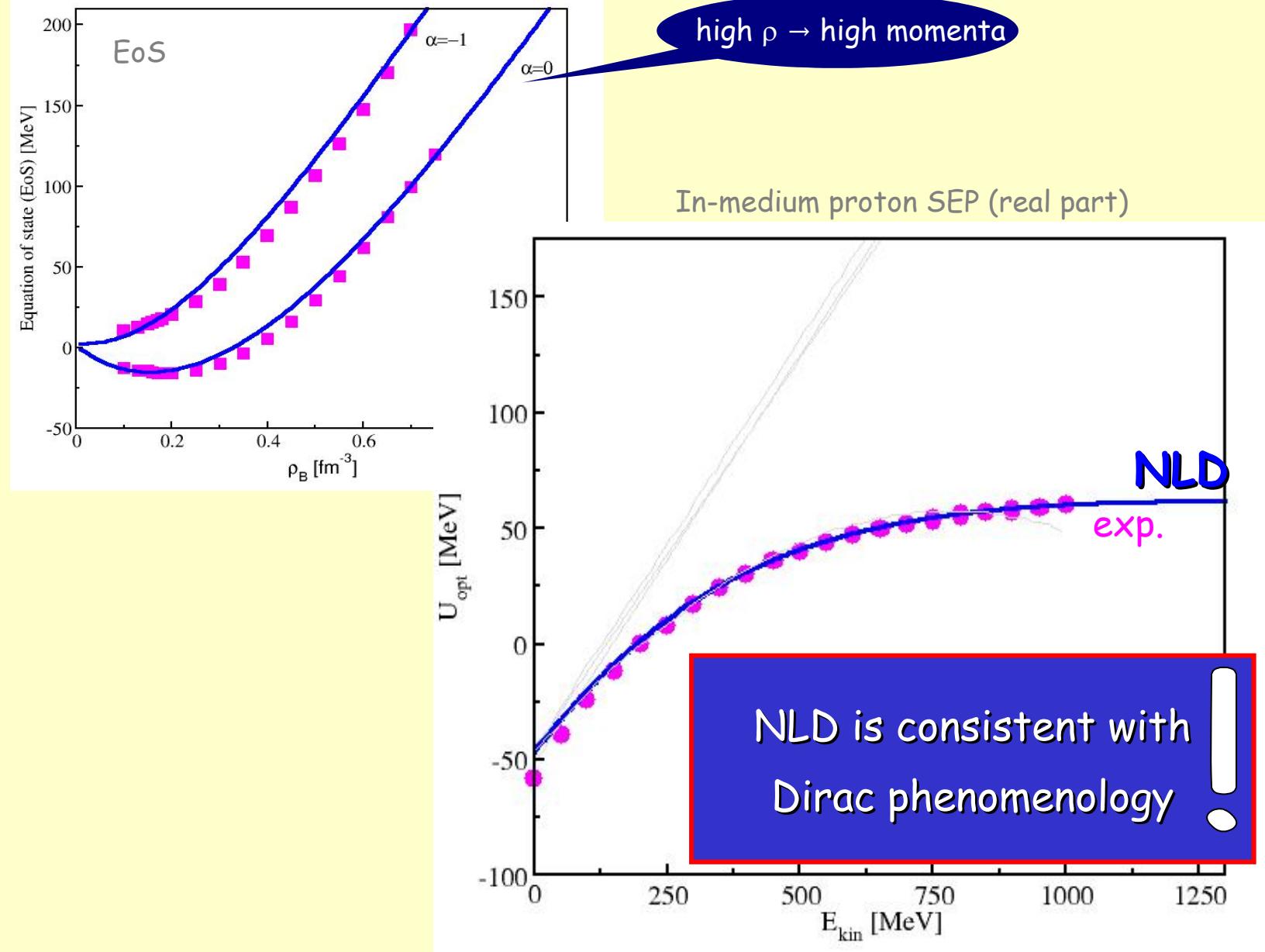
In-medium proton SEP (real part)



NLD results: MD & optical potentials...

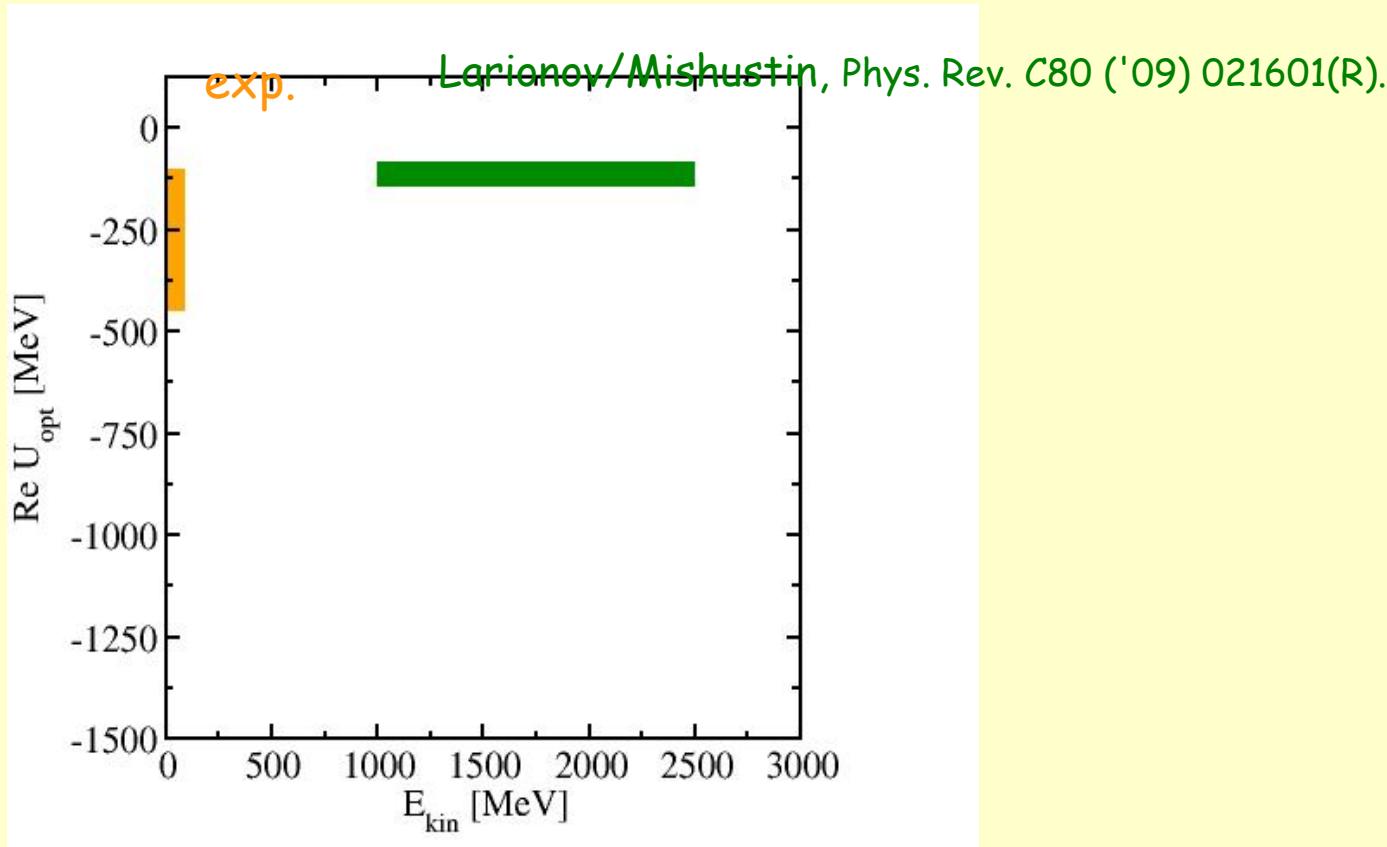


NLD results: MD & optical potentials...



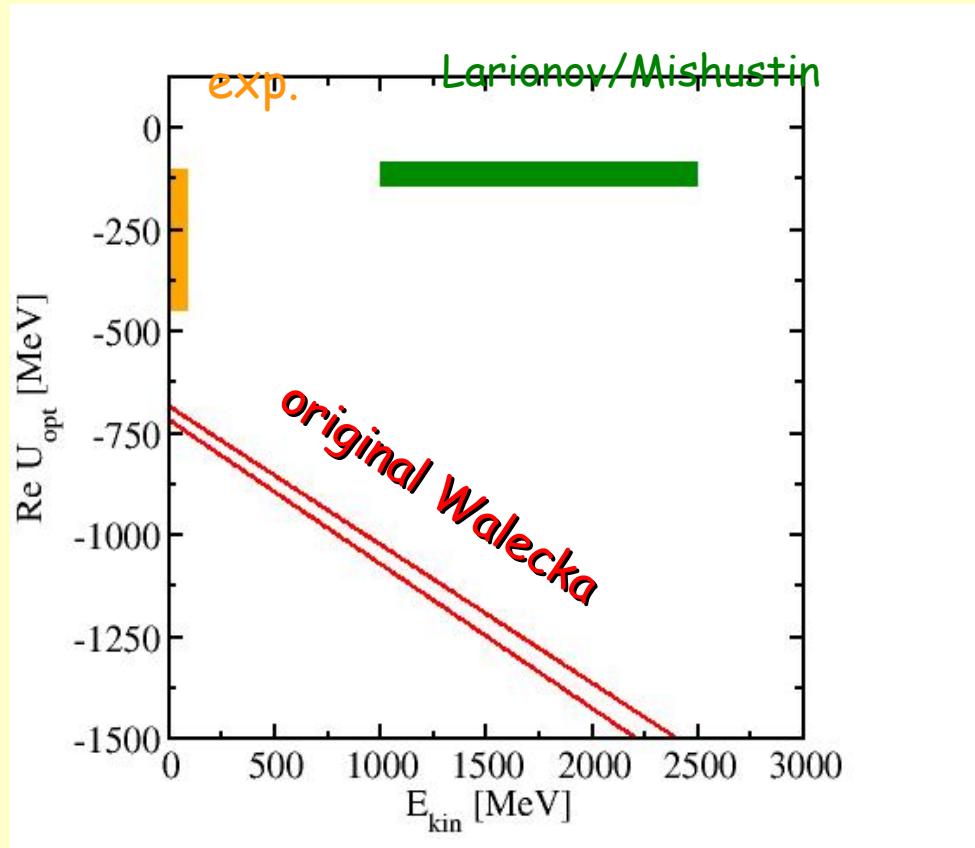
NLD results: MD & optical potentials (anti-proton)...

In-medium anti-proton SEP (real part)



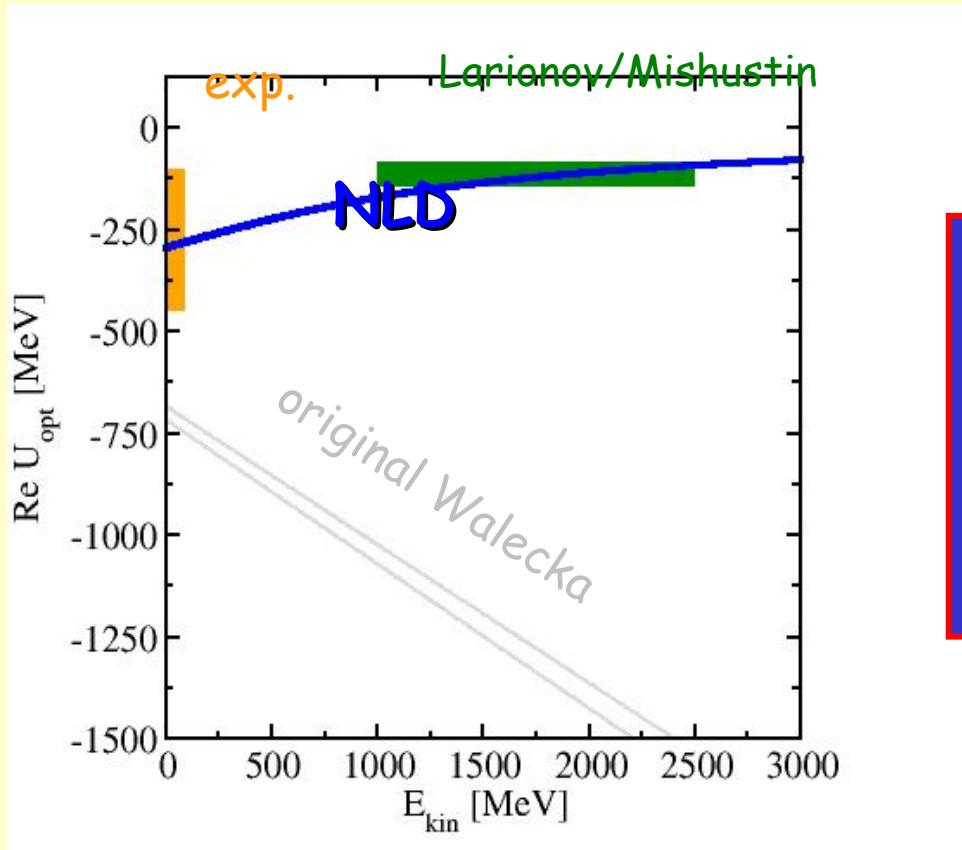
NLD results: MD & optical potentials (anti-proton)...

In-medium anti-proton SEP (real part)



NLD results: MD & optical potentials (anti-proton)...

In-medium anti-proton SEP (real part)



good description of
phenomenology

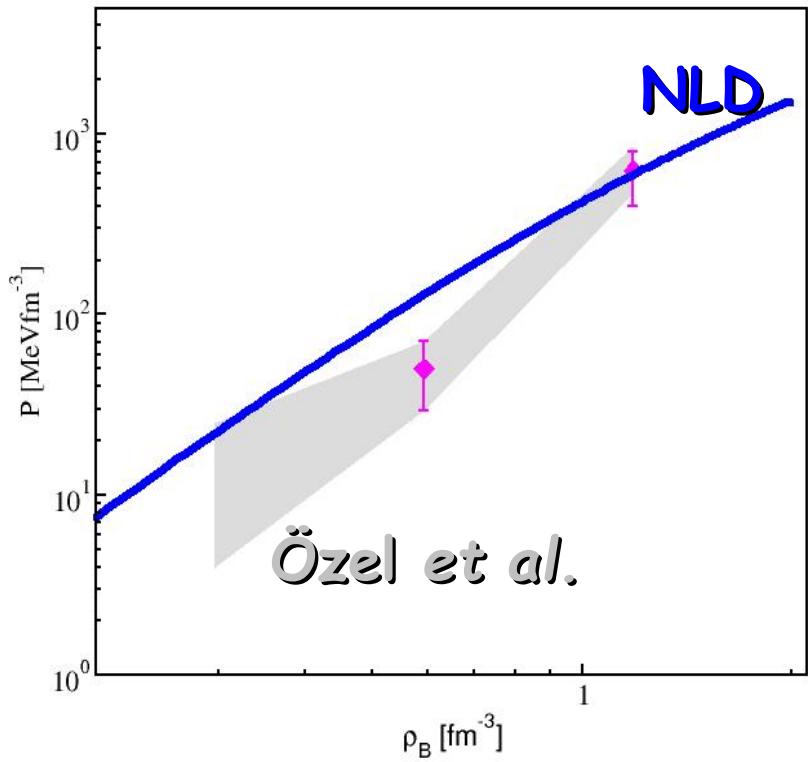
also for in-medium

anti-proton interaction

Also: NLD provides the imaginary part of SEP for anti-proton in-medium interactions using dispersion relation (without subtractions) →

Phys. Lett. B703, ('11) 193

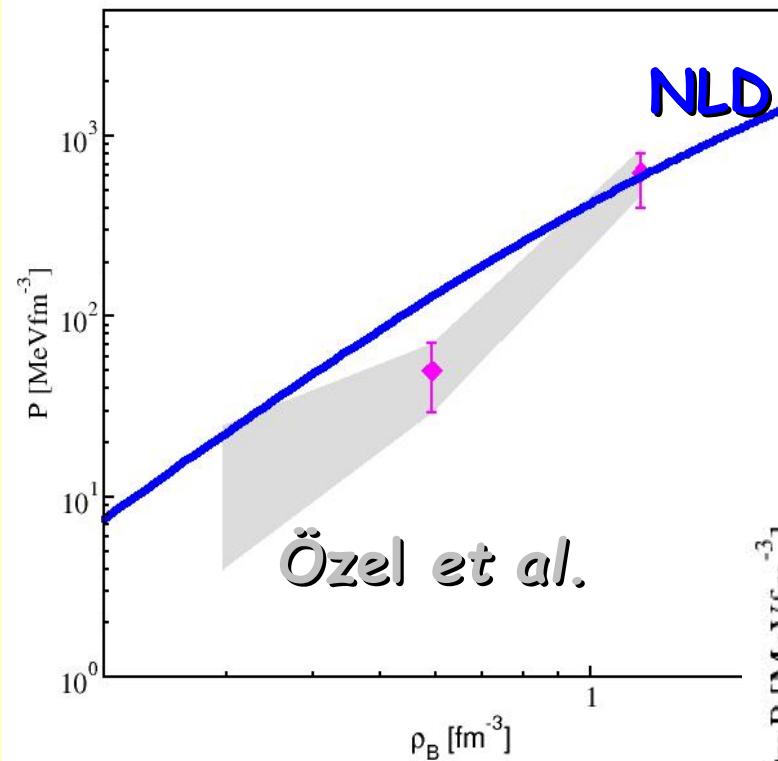
NLD results: high-density EoS at β -equilibrium...



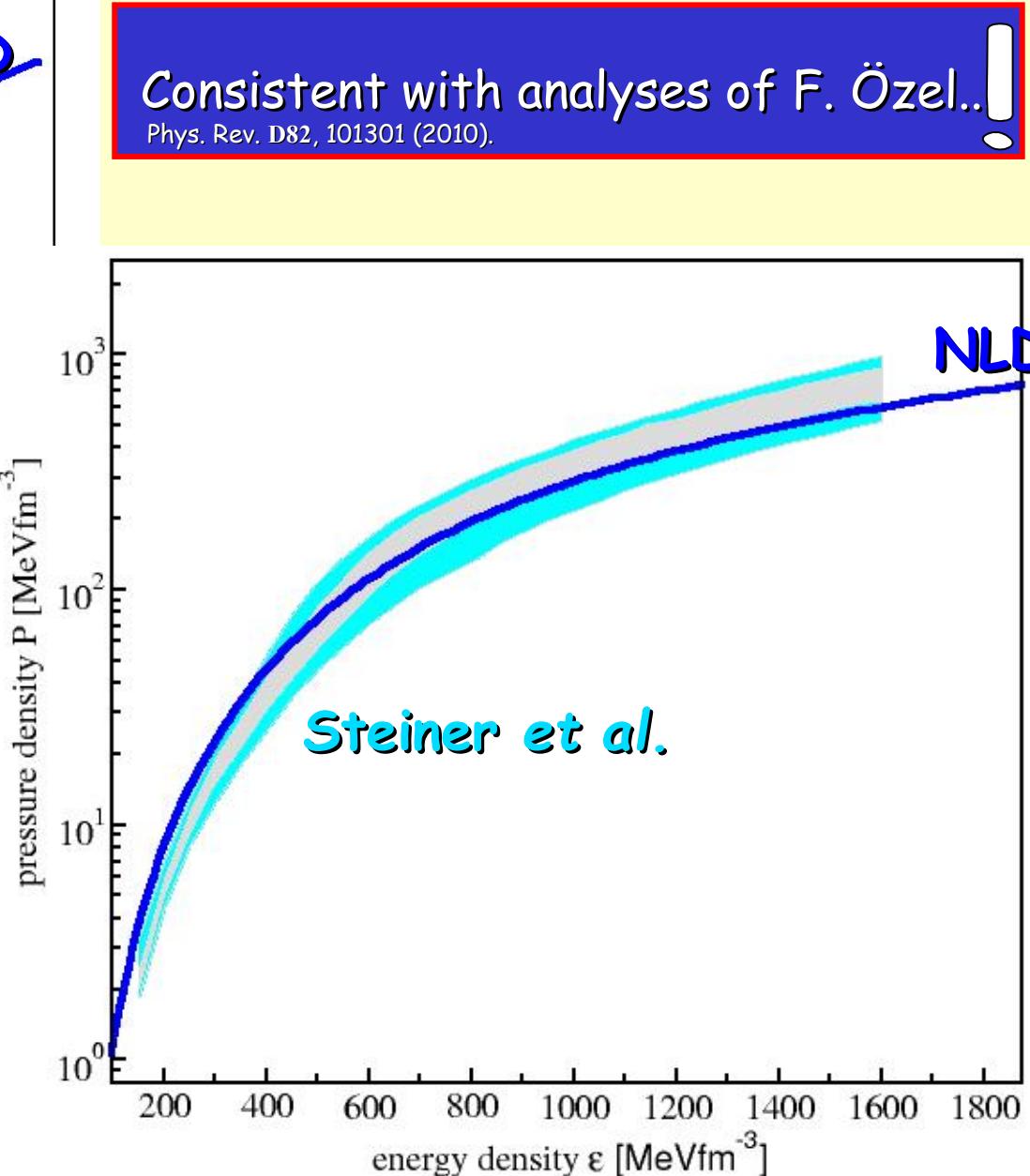
Consistent with analyses of F. Özel..!

Phys. Rev. D82, 101301 (2010).

NLD results: high-density EoS at β -equilibrium...



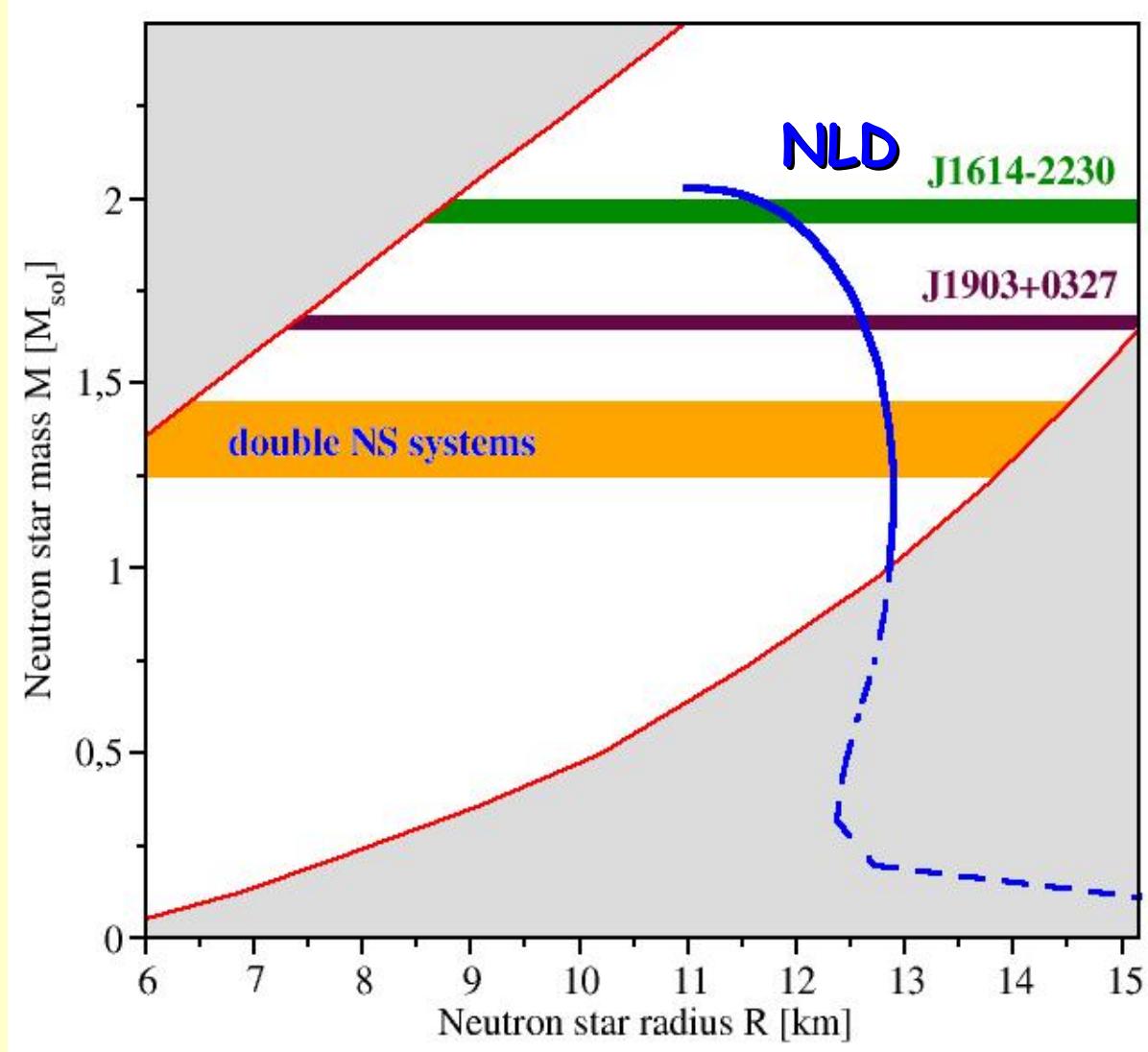
... and A.W. Steiner
Astrophys. J. 722, 33 (2010).



NLD results: NS mass...

Compatible with ...

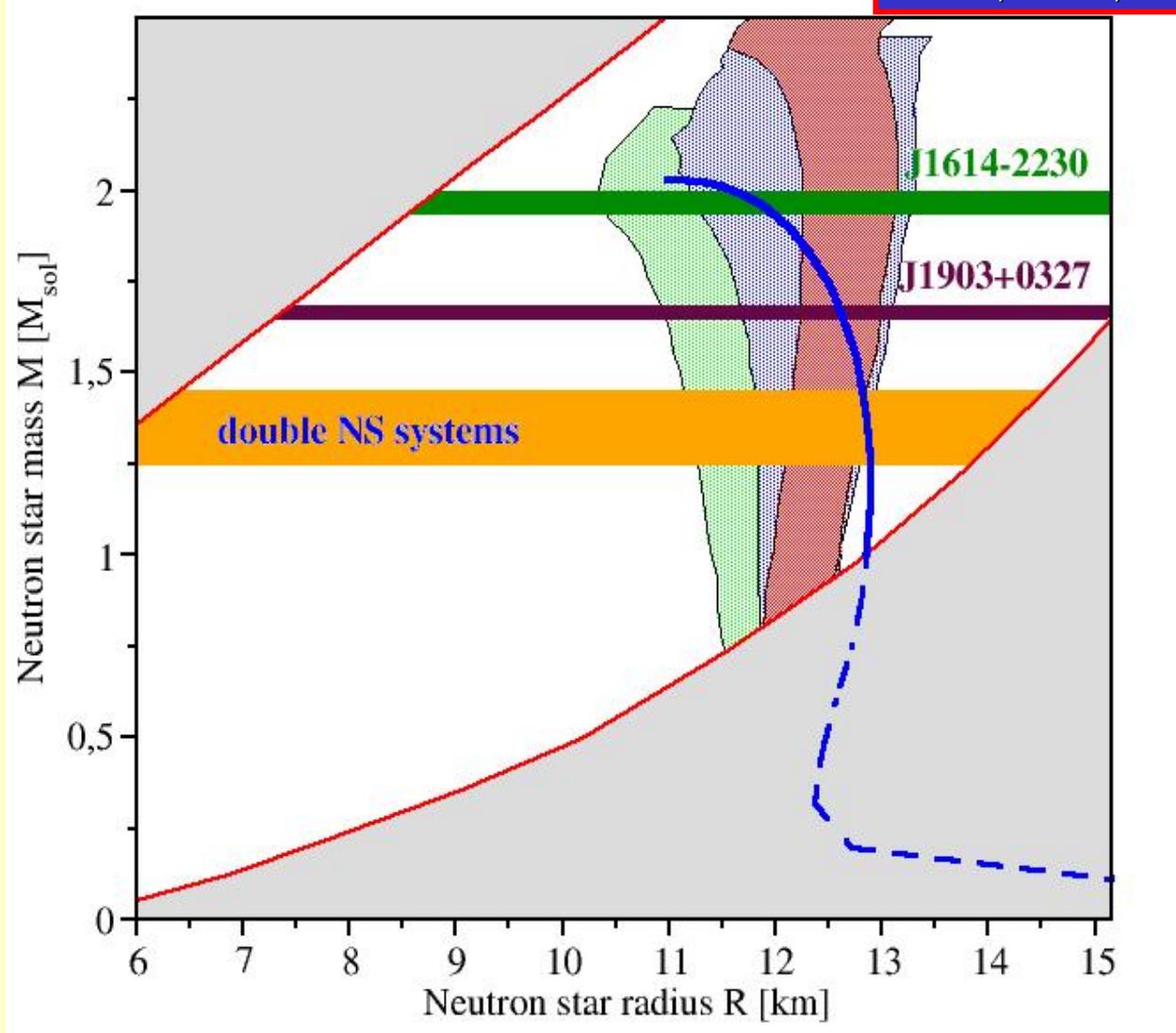
Lattimer/Prakash, Science 304 ('04) 536, Phys. Rep. 442 ('07) 109



NLD results: NS mass...

Compatible with all observations

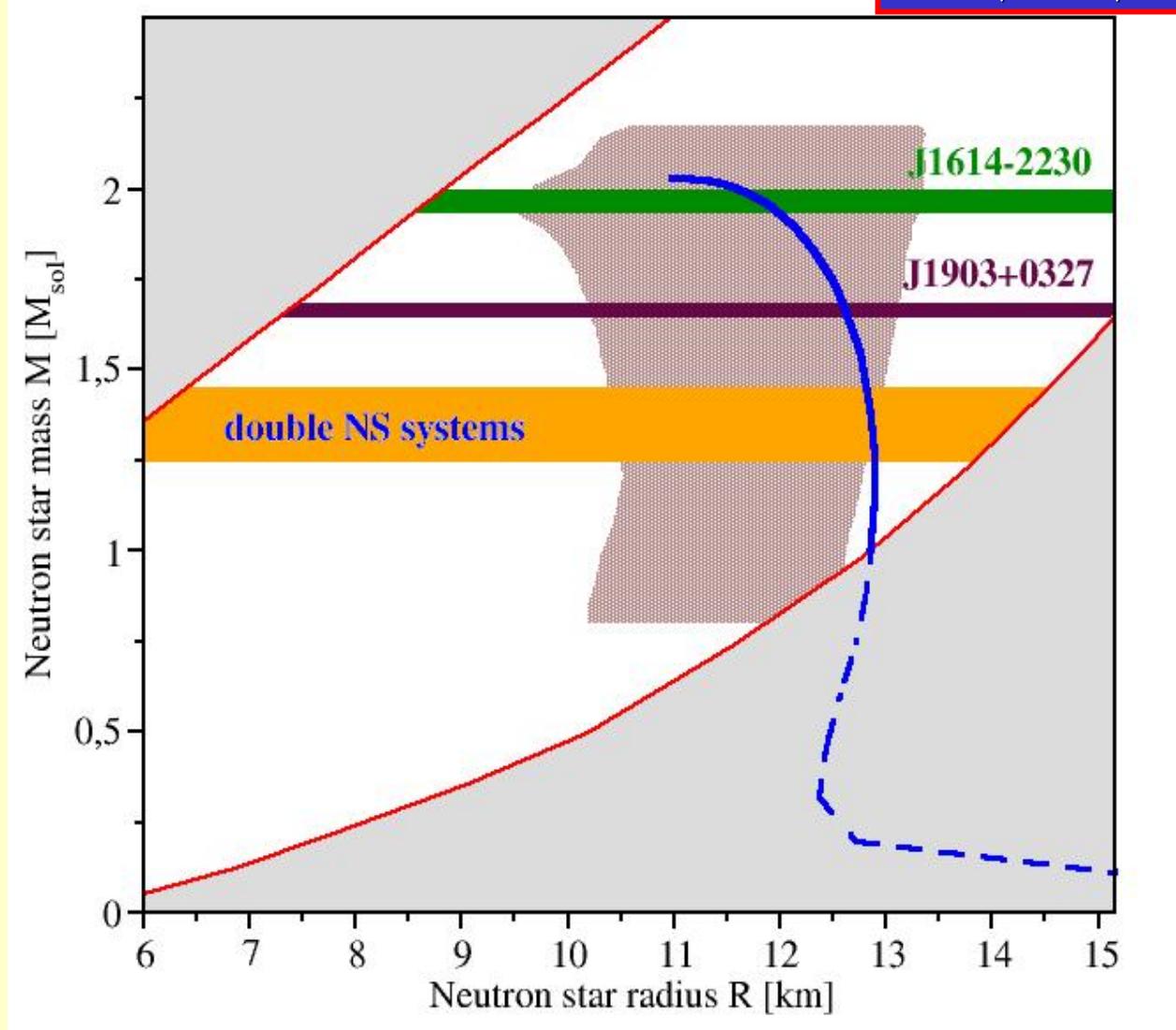
Lattimer/Prakash, Science 304 ('04) 536, Phys. Rep. 442 ('07) 109
Steiner, Lattimer, Brown, arXiv: 1205.6871



NLD results: NS mass...

Compatible with all observations !

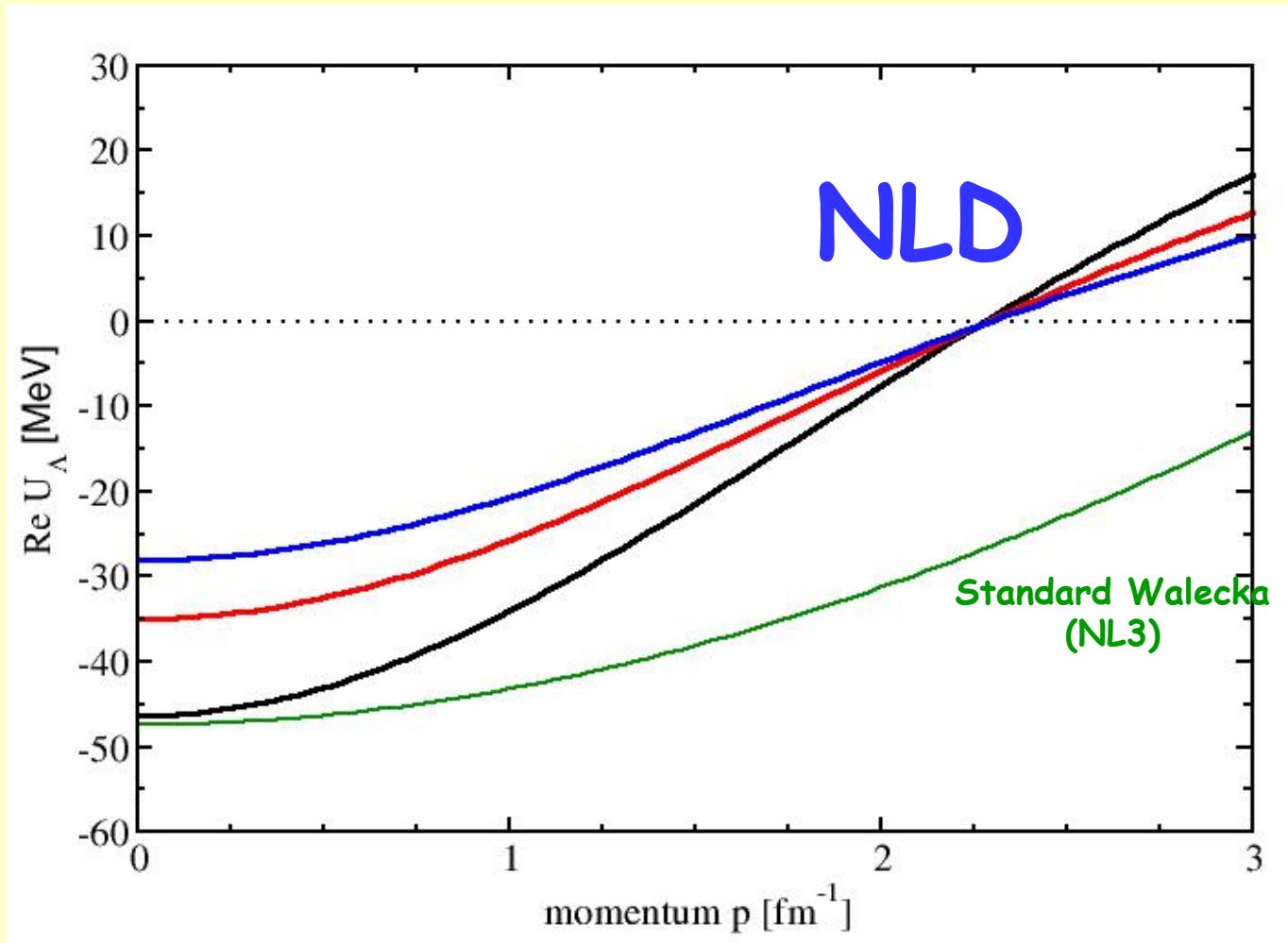
Lattimer/Prakash, Science 304 ('04) 536, Phys. Rep. 442 ('07) 109
Steiner, Lattimer, Brown, arXiv: 1205.6871



NLD results: in-medium Λ -opt. potential...

NLD + SU(3) for standard meson-nucleon couplings

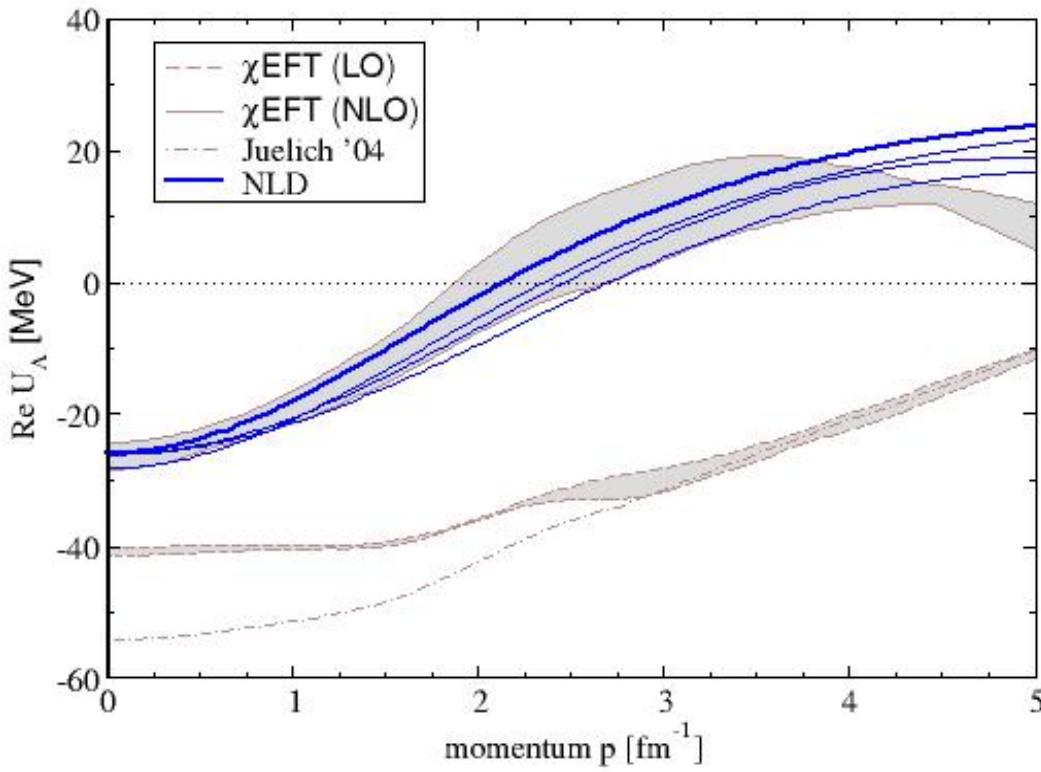
Hyperon cut-off regulates MDI



NLD results: in-medium Λ/Σ -opt. potentials...

NLD + SU(3) for standard meson-nucleon couplings

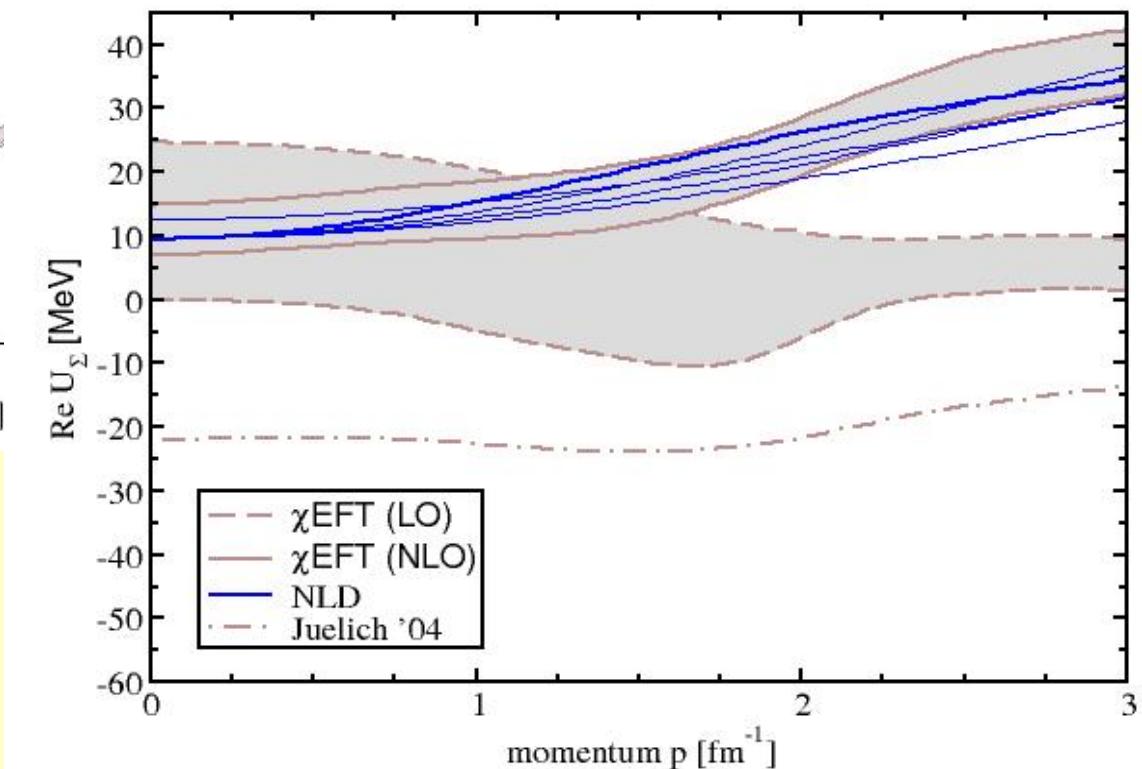
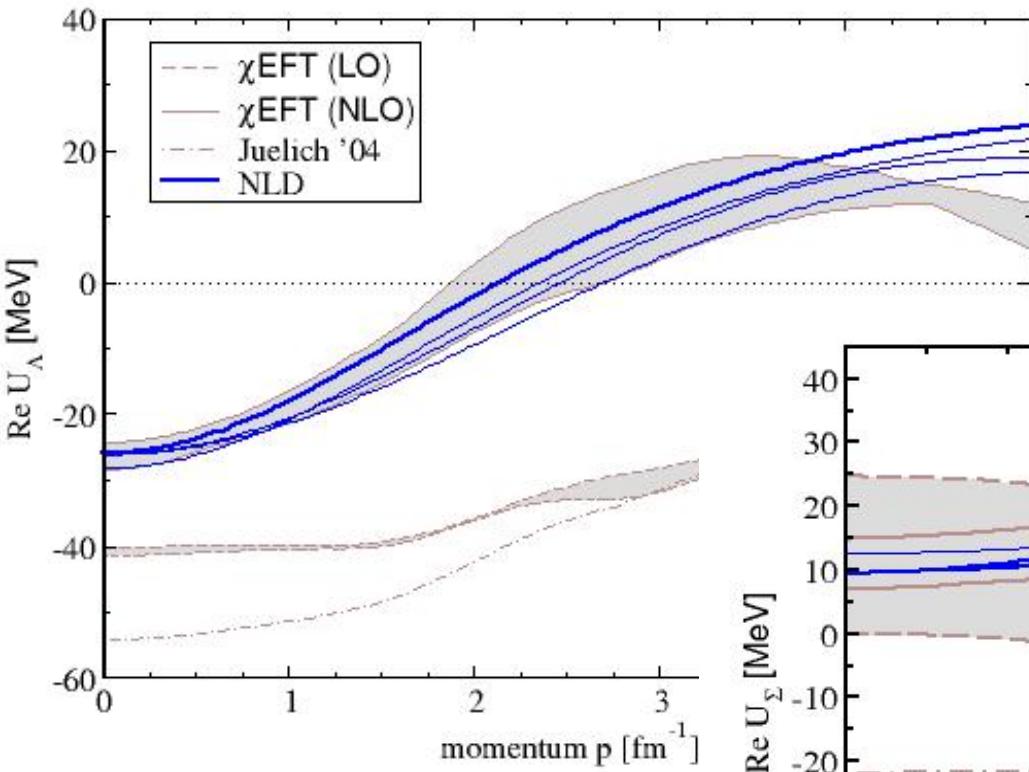
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NLD + SU(3) for standard meson-nucleon couplings

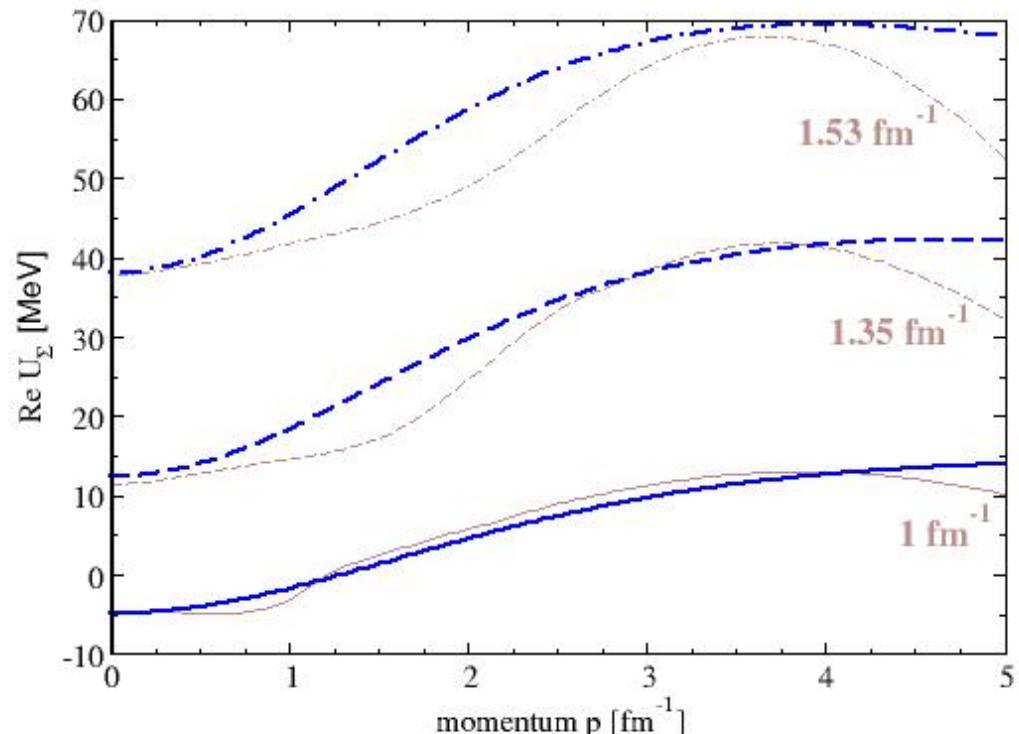
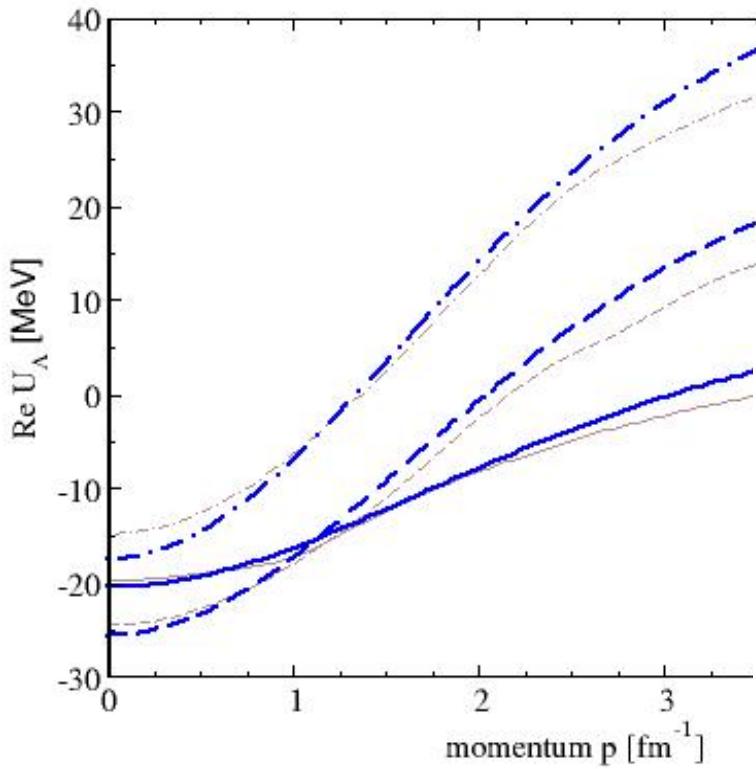
Hyperon cut-off regulates MDI



NLD results: in-medium Λ/Σ -opt. potentials...

NLD + SU(3) for standard meson-nucleon couplings

Hyperon cut-off regulates MDI



Explore in-medium Λ -pot: HADES experiment...

Key Information

- ▶ HADES Spectrometer at GSI
- ▶ Secondary π^- -beam 1.7 GeV/c

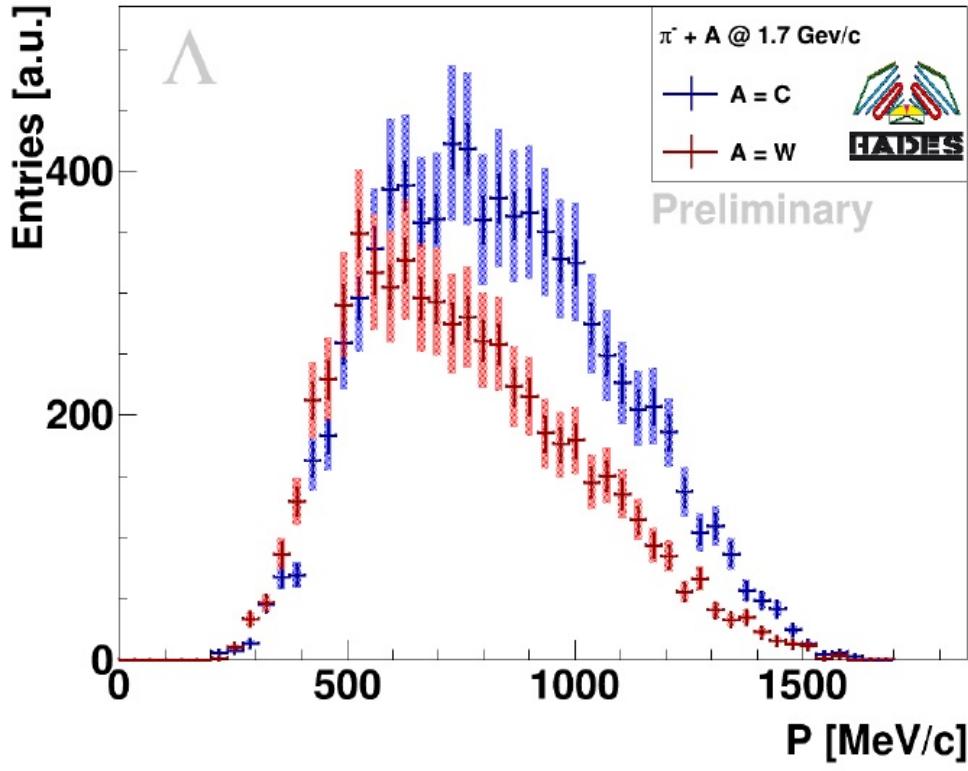
⌚ Target	W	C
➡ Segment Length [mm]	2.4	7.2
📦 ρ [g/cm^3]	19.3	1.85
☀ A	183.84	12.011
📊 Statistics [$\times 10^8$]	1.69	2.00

Idea:

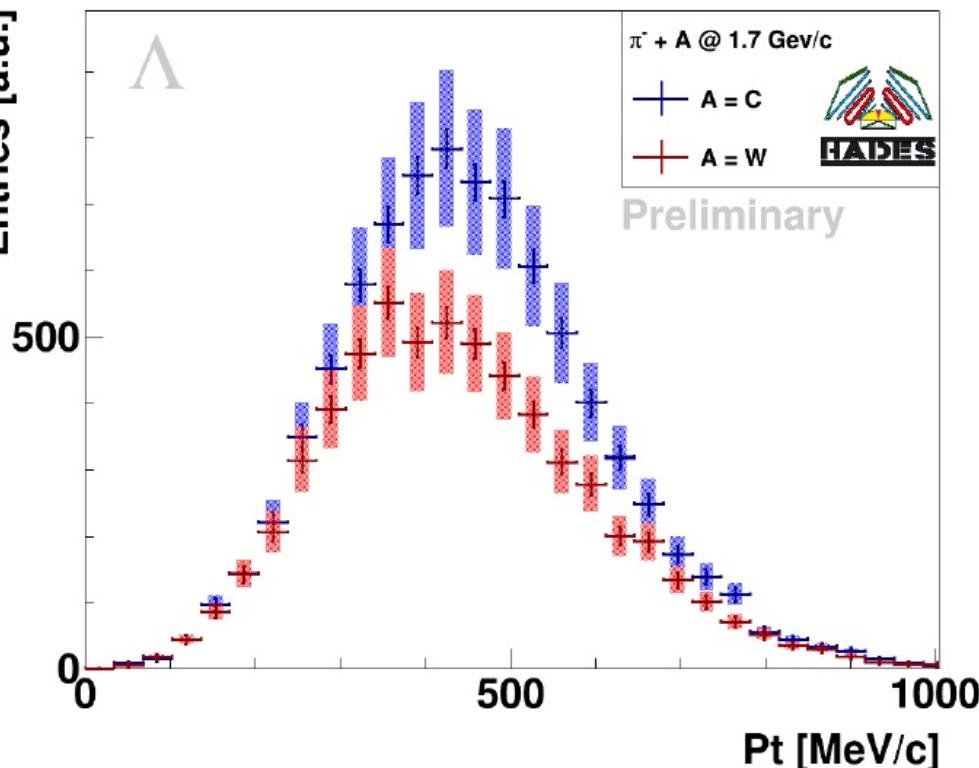
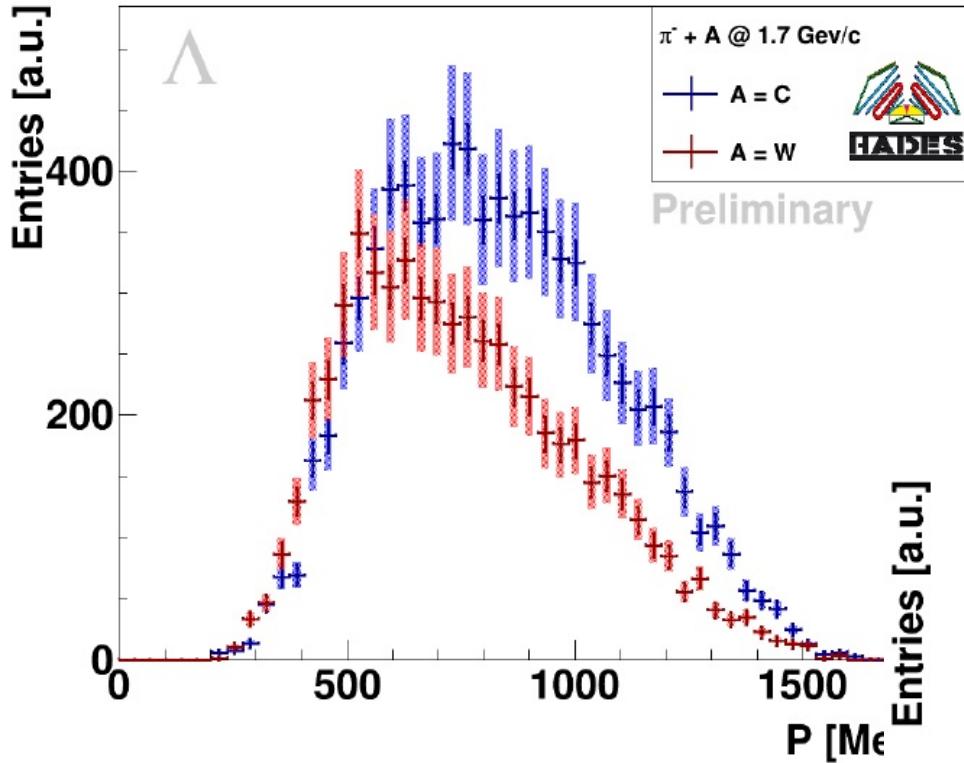
- ▶ Search for Charge Pattern 2+ 2- ($\Lambda \rightarrow p + \pi^-$, $K^0 \rightarrow \pi^+ + \pi^-$)
- ▶ Make best assignment of double π^- occurrence by minimizing:
 $\Delta M_\Lambda = M_{INV}(p + \pi^-) - M(\Lambda)_{PDG}$
 $\Delta M_{K0} = M_{INV}(\pi^+ + \pi^-) - M(K0)_{PDG}$
For all π^- Combination
- ▶ Cut on 2D ΔM_Λ vs. ΔM_{K0}

Icons from: <https://www.flaticon.com/>

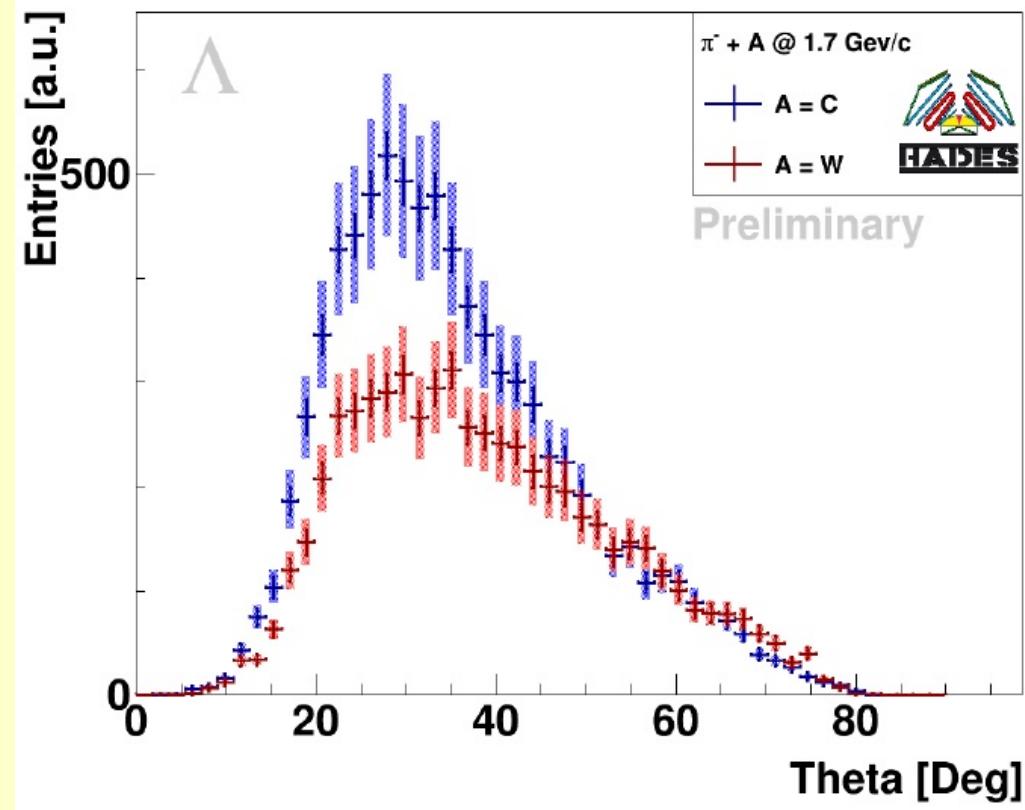
Explore in-medium Λ -pot: HADES new data...



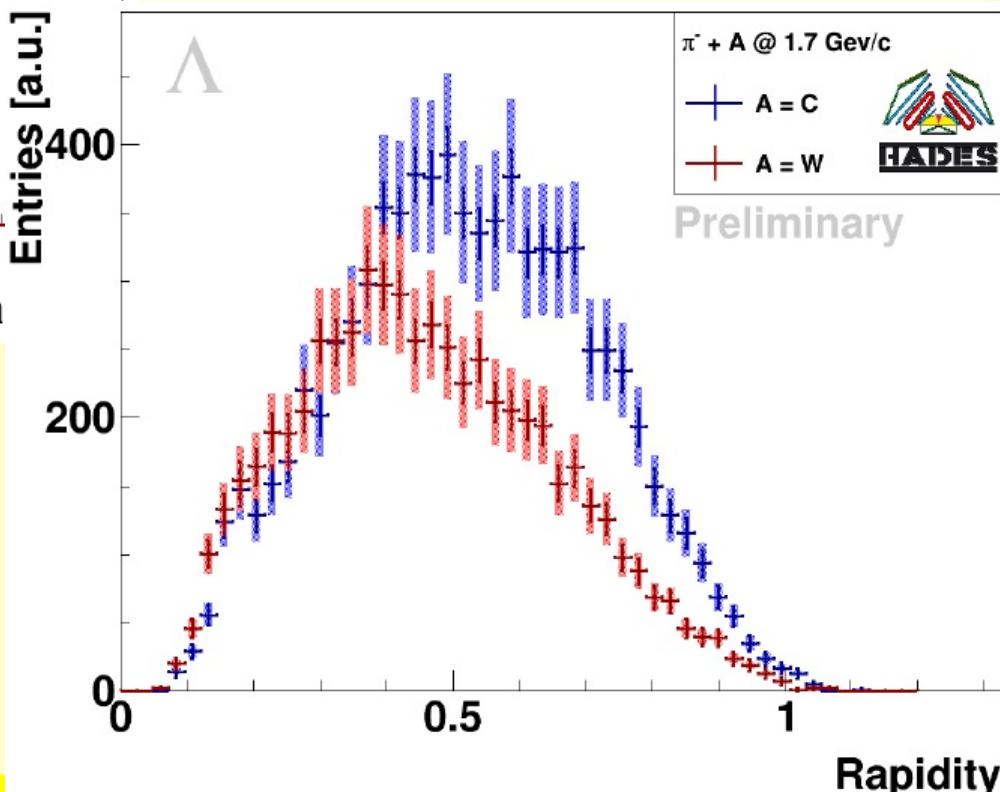
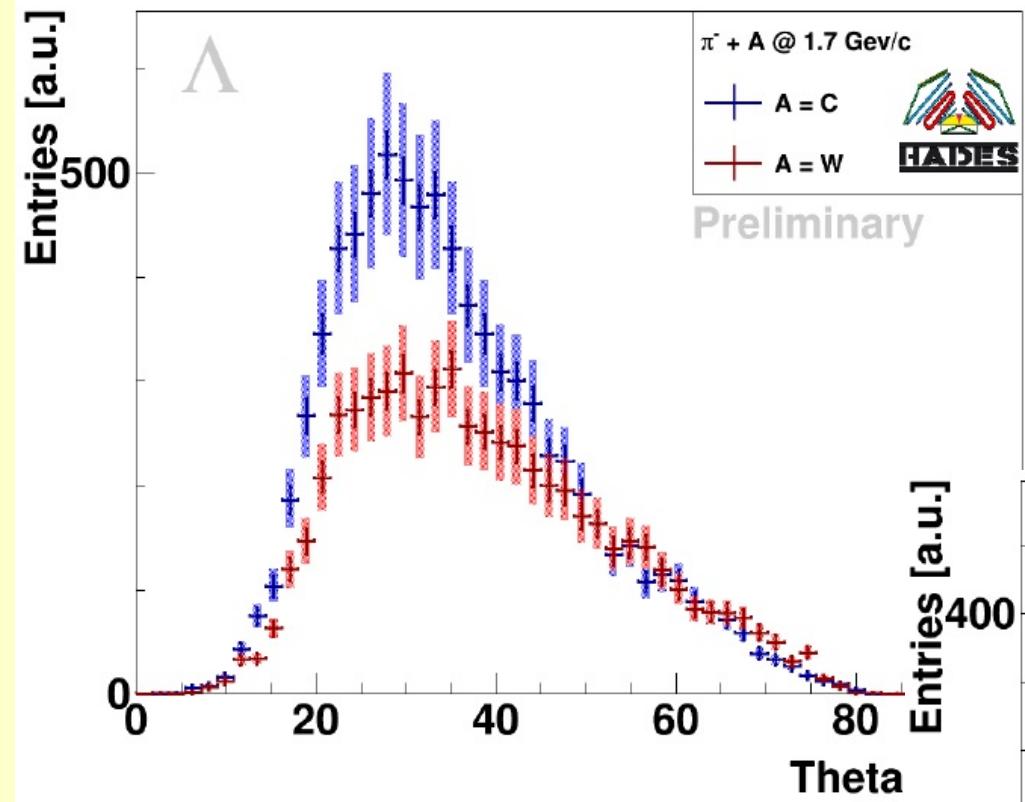
Explore in-medium Λ -pot: HADES new data...



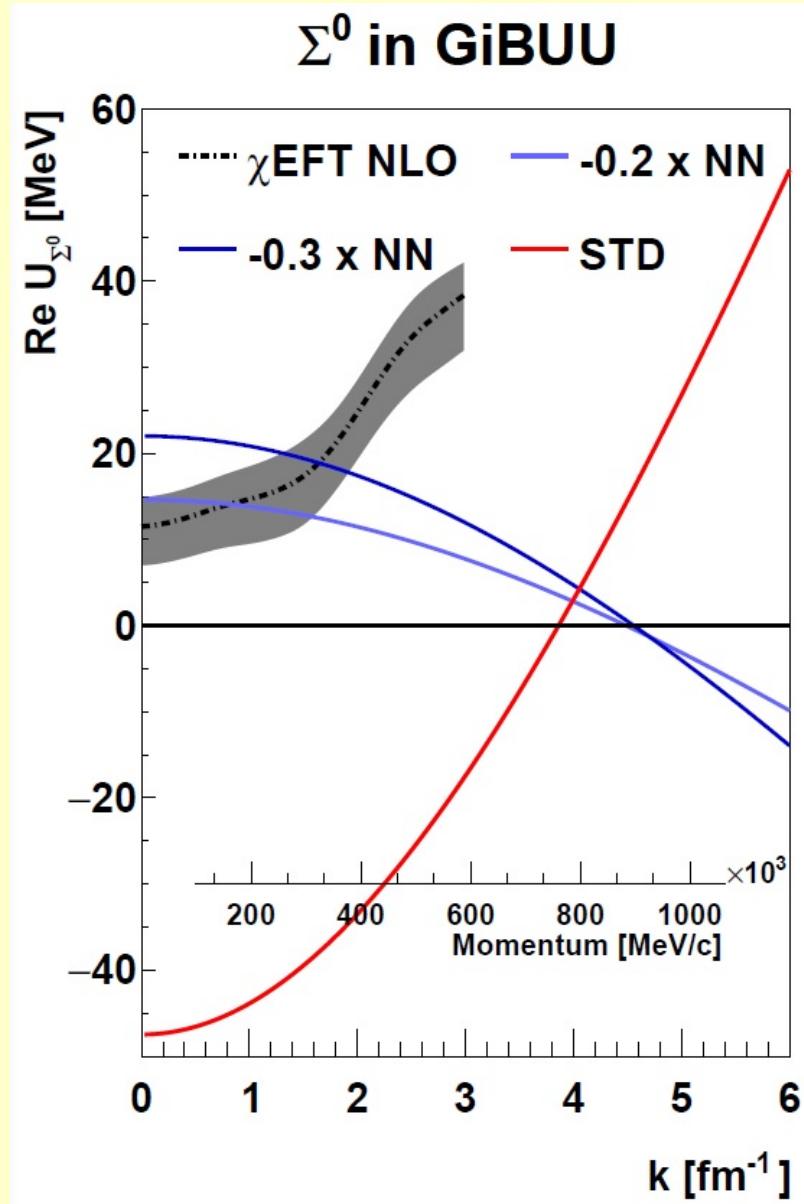
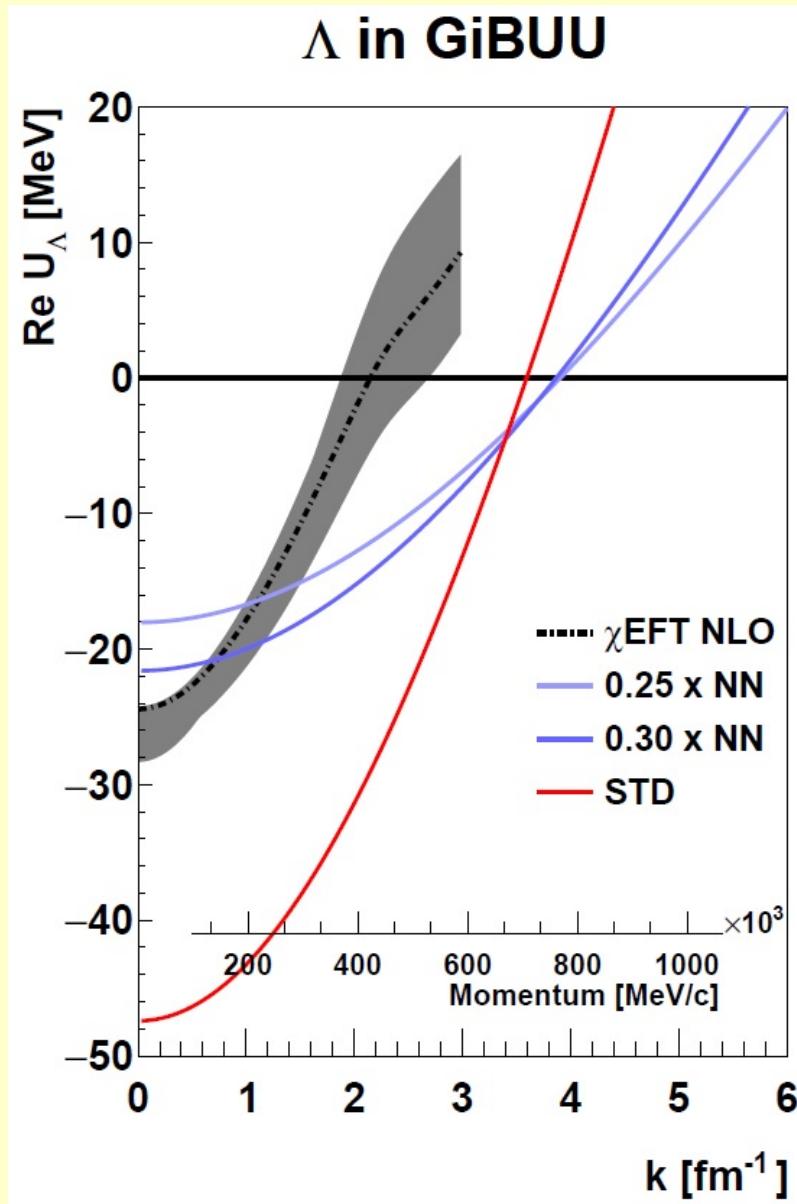
Explore in-medium Λ -pot: HADES new data...



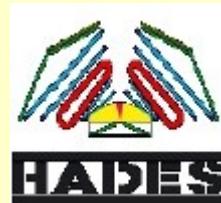
Explore in-medium Λ -pot: HADES new data...



Hyperon potentials & new HADES data: the input...

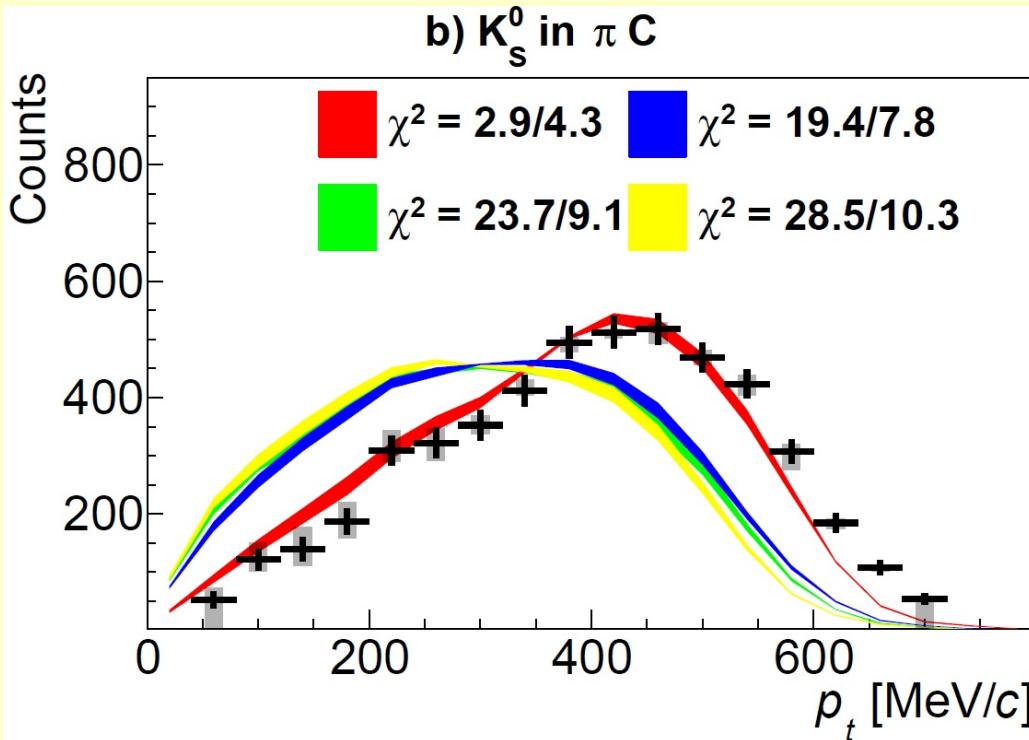


Hyperon potentials & new HADES data: the comparison...



PRELIMINARY

Hyperon potentials & new HADES data: the comparison...



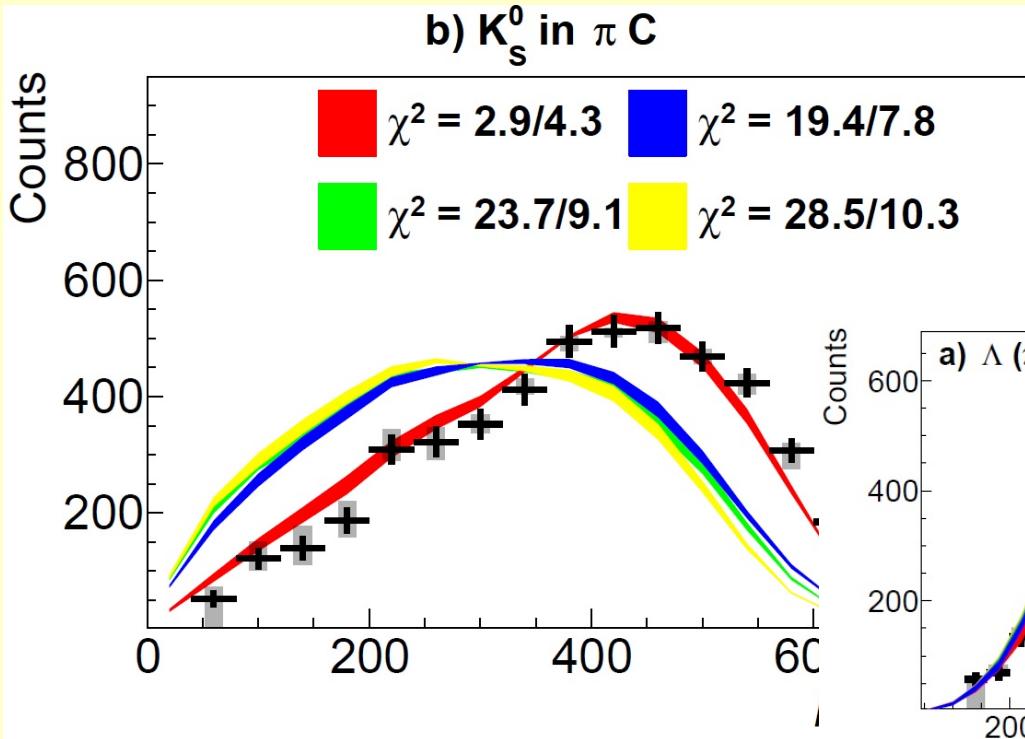
PRELIMINARY

$U_{\Lambda,\Sigma}$ assumptions vs data

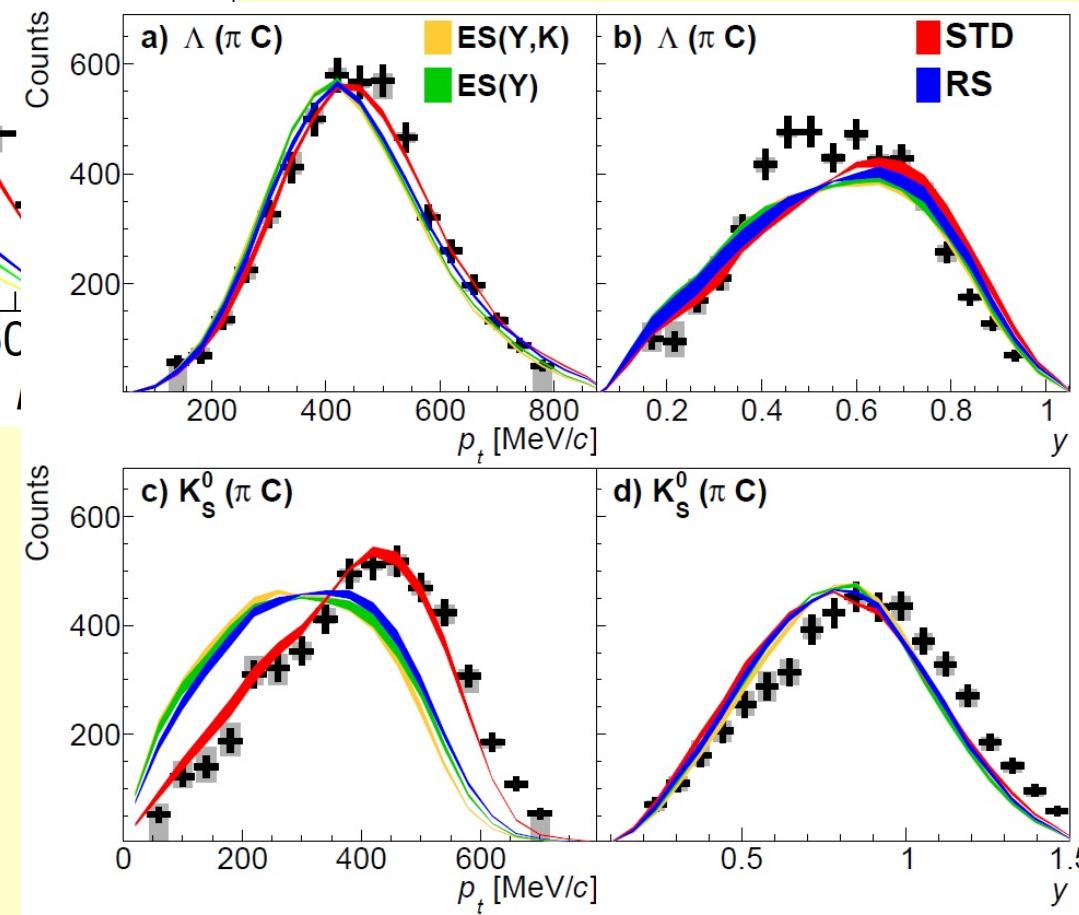
Red (STD) $\rightarrow \Lambda, \Sigma$ attractive

Blue (RS) $\rightarrow \Sigma$ repulsive!

Hyperon potentials & new HADES data: the comparison...



PRELIMINARY

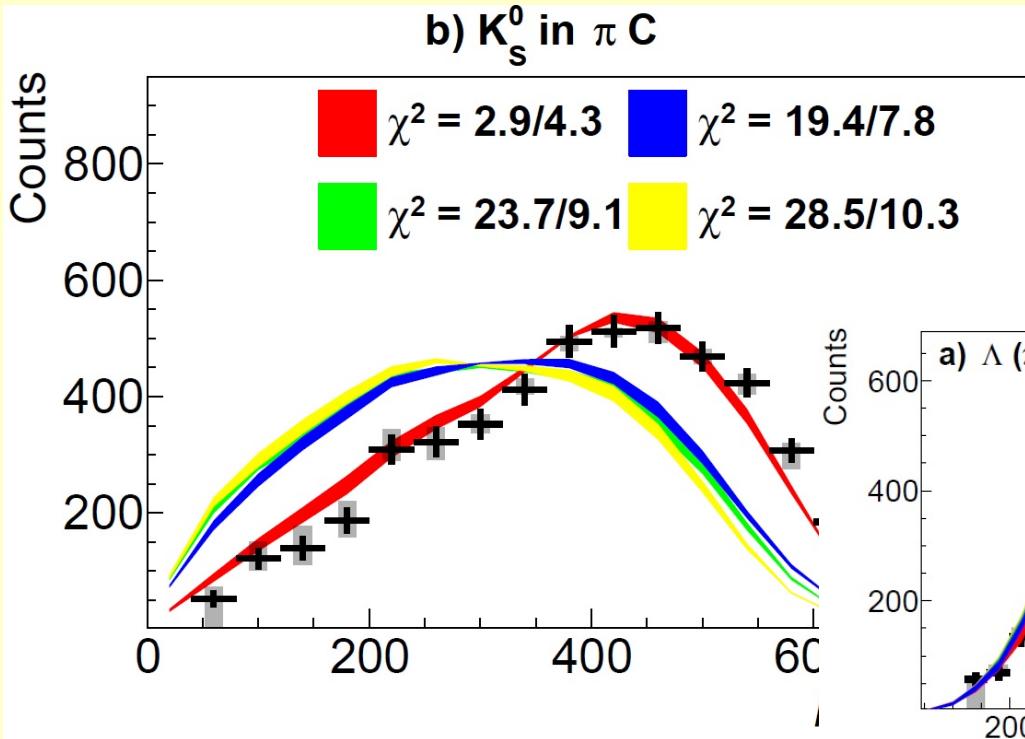


$U_{\Lambda,\Sigma}$ assumptions vs data

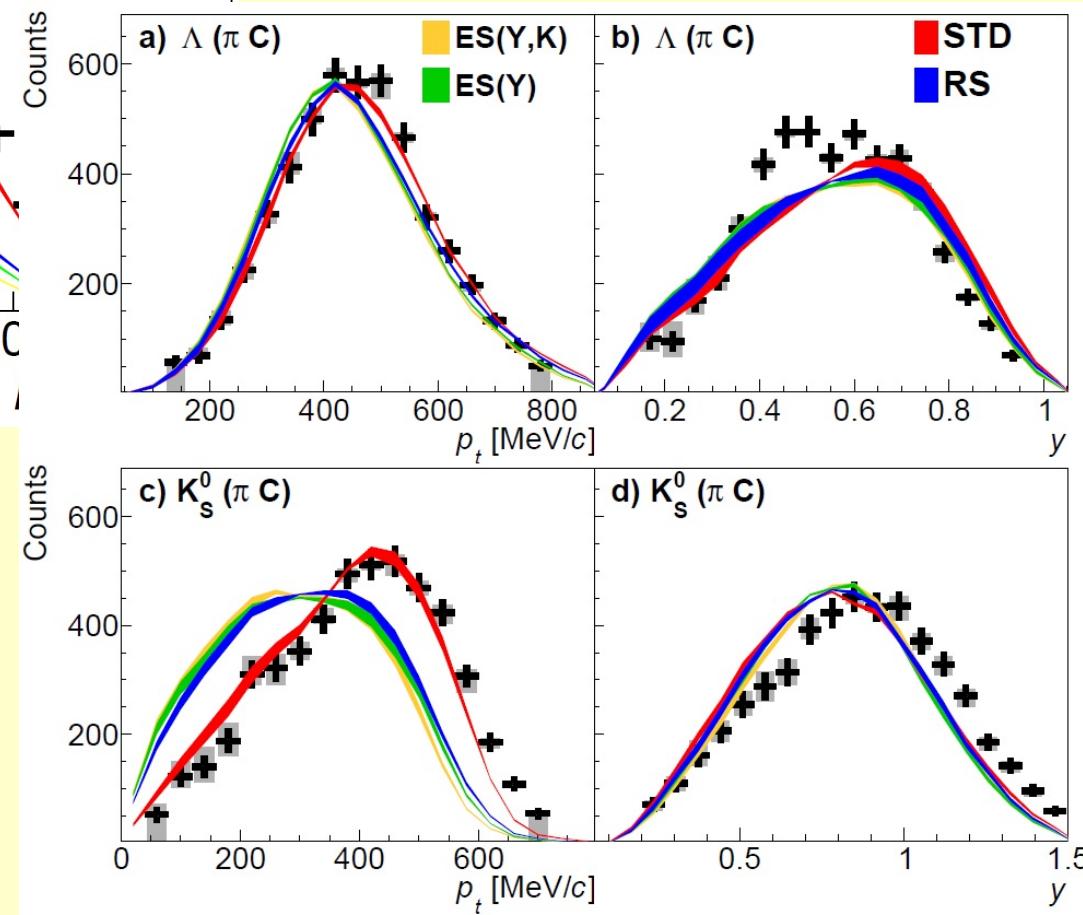
Red (STD) $\rightarrow \Lambda, \Sigma$ attractive

Blue (RS) $\rightarrow \Sigma$ repulsive!

Hyperon potentials & new HADES data: the comparison...



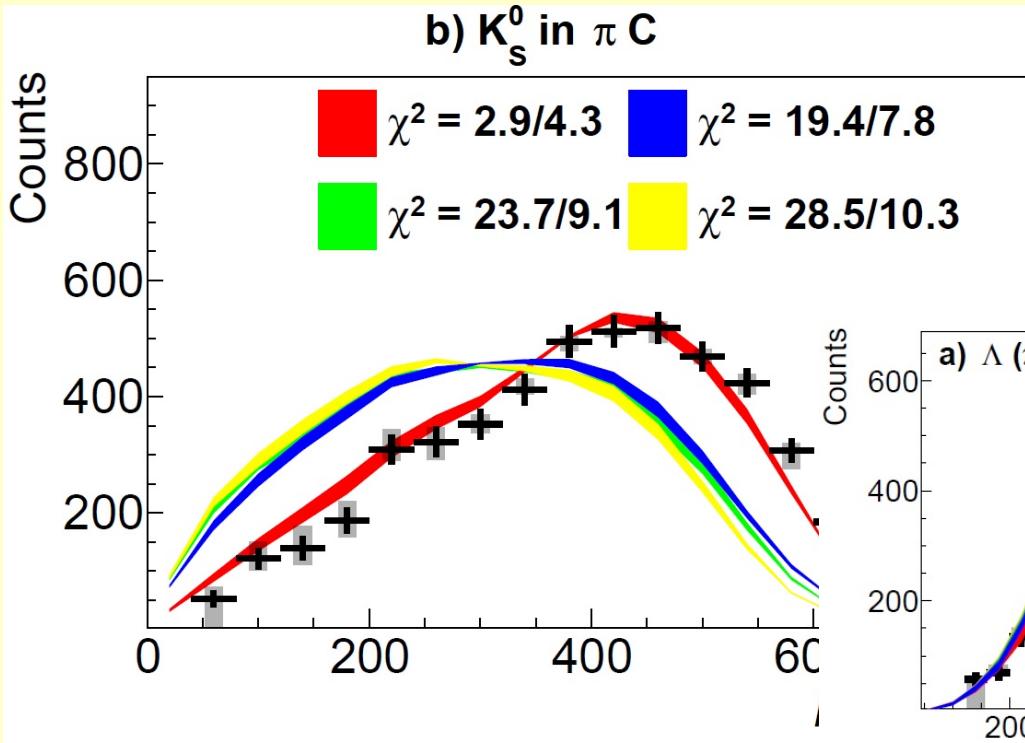
PRELIMINARY



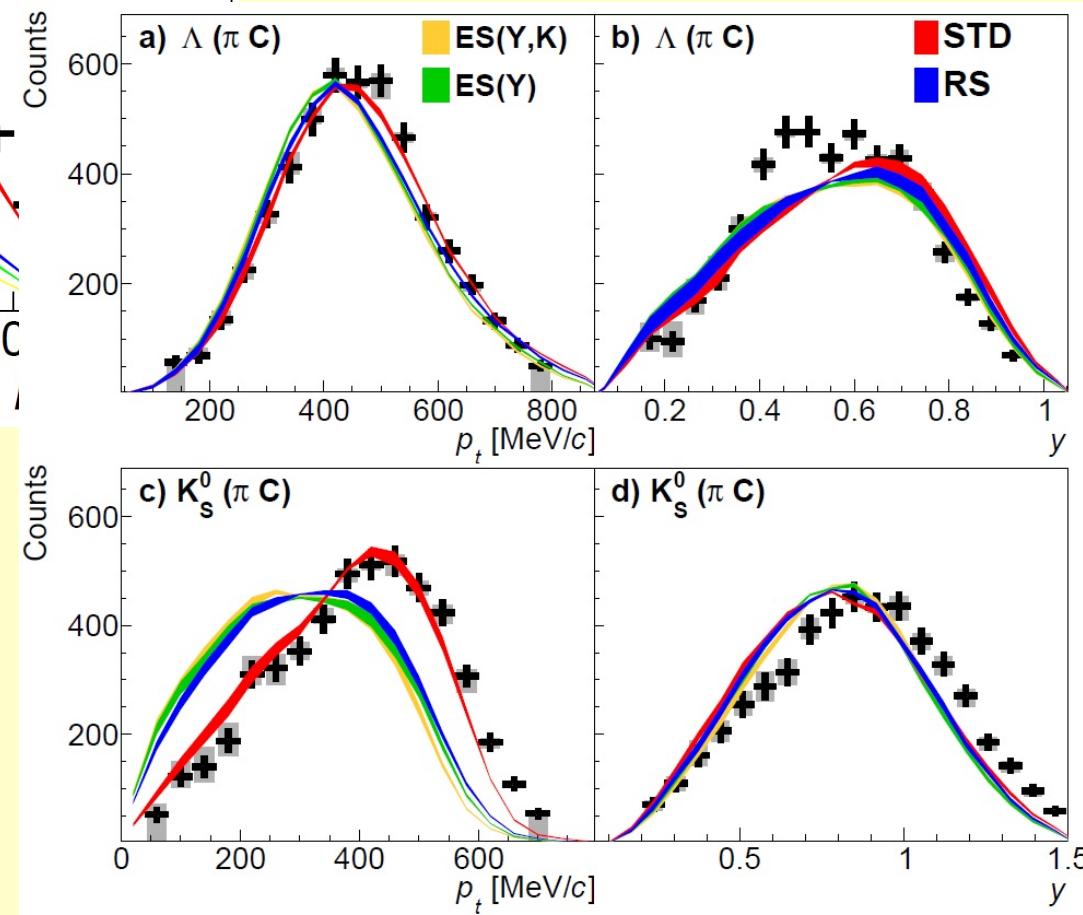
Conclusion

Σ -repulsive potential not supported
by new precise HADES
measurements

Hyperon potentials & new HADES data: the comparison...



PRELIMINARY



Consequences

- EoS with hyperons seems soft
- Soft EoS for neutron stars,
if hyperons included

Hyperon potentials & new HADES data: the comparison...

b) K_s^0 in πC

Counts

800
600
400
200
0

■ $\chi^2 = 2.9/4.3$ ■ $\chi^2 = 19.4/7.8$
■ $\chi^2 = 23.7/9.1$ ■ $\chi^2 = 28.5/10.3$

Counts

600



PRELIMINARY

Hyperon puzzle "supported"
by the new data

STD
s

400 600 800

p_t [MeV/c]

y

Counts

600

d) K_s^0 (πC)

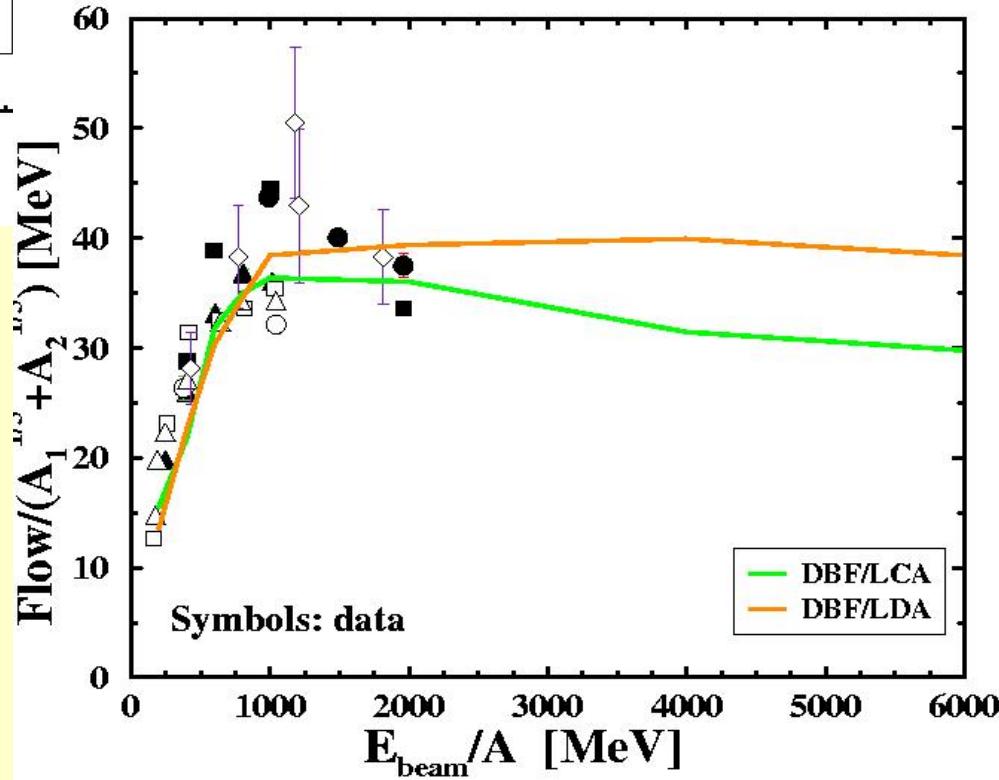
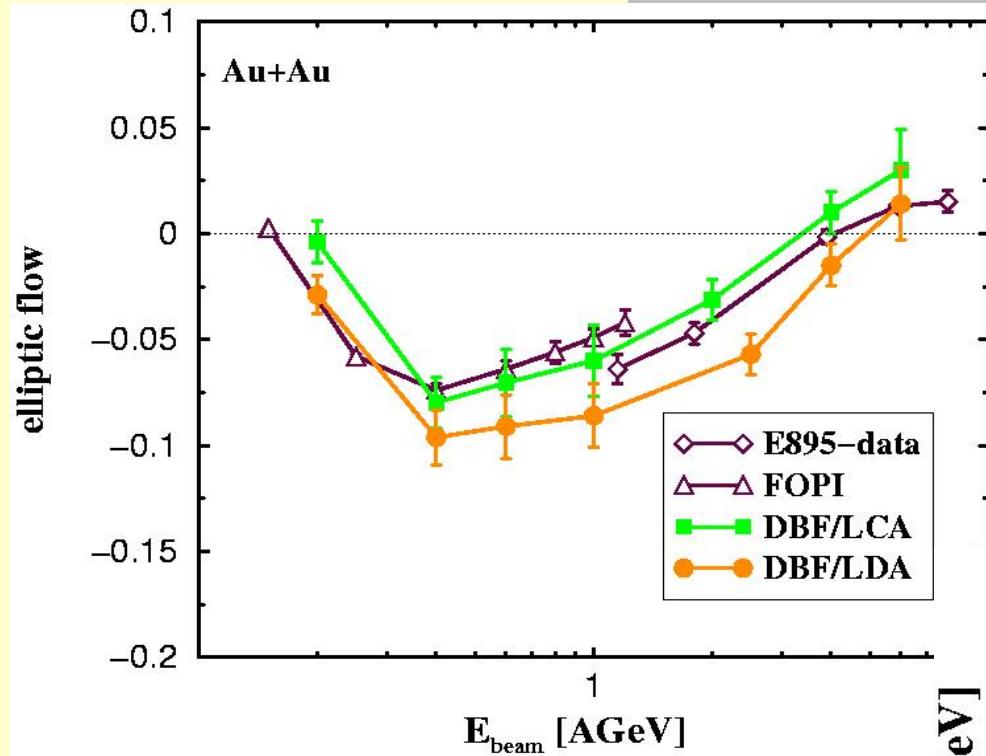
200 400 600

p_t [MeV/c]

y

Backup slides

Nuclear EoS from HIC



Gaitanos, Fuchs, Wolter, Faessler,
Eur. Phys. J. A12 (2001) 421

Gaitanos, Fuchs, Wolter,
Nucl. Phys. A741 (2004) 287